

# MODIS LUT INFORMATION GUIDE

For Level 1B Version 4.3.0 (Terra) and Version 4.3.1 (Aqua)



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# 1 INTRODUCTION AND CHANGES FROM THE PREVIOUS RELEASE

The Level 1B lookup tables (LUTs) contain input parameters to the Level 1B code. Contents of this document include:

- General Information, Conventions and Versioning Strategy (Section 2)
- Reflective Calibration LUTs (Section 3)
- Emissive Calibration LUTs (Section 4)
- Quality Assurance LUTs (Section 5)

There have been several changes to the structure and number of LUTs from previous software releases. Table 1.1, Table 1.2, and Table 1.3 summarize the changes to the Reflective, Emissive, and Quality Assurance LUTs since April 2001.

**Table 1.1 Changes to the Reflective LUT Collection and Structure Since April 2001**

<b>REFLECTIVE LUT CHANGES</b>		
<b>Reflective LUTs added:</b>	<b>Section</b>	
• "RVS_RefSB"	3.3.8	
• "SWIR_OOB_corr_sending_band	3.3.20	
• "B26_B5_Corr_Switch "	3.3.33	
• "B26_B5_Frame_Offset "	3.3.34	
• "B26_B5_Corr"	3.3.35	
<b>Reflective LUTs deleted:</b>	<b>Reason</b>	
• "RVS_250m"	Replaced by RVS_RefSB	
• "RVS_500m"	Replaced by RVS_RefSB	
• "RVS_1km_RefSB"	Replaced by RVS_RefSB	
• "delta_DN_RSB"	ADC Correction eliminated	
• "R_star"	Use in uncertainty calculation replaced by inverse of $m1 * (\text{Solar Irradiance} / \text{Pi})$	
<b>Reflective LUTs changed:</b>	<b>Section</b>	<b>Nature of Change</b>
• "k_inst"	3.3.4	Fill values eliminated
• "k_fpa"	3.3.5	Fill values eliminated
• "m0 "	3.3.6	Fill values eliminated
• "m1"	3.3.7	Fill values eliminated
• "Sigma_m1"	3.3.11	Fill values eliminated
• "Sigma_K_inst"	3.3.12	Fill values eliminated
• "rRSB_NEdl"	3.3.16	Fill values eliminated
• "dn_sat_ev"	3.3.32	Data type changed to float64; fill values eliminated

**Table 1.2: Changes to the Emissive LUT Collection and Structure Since April 2001**

<b>EMISSIVE LUT CHANGES</b>		
<b>Emissive LUTs added:</b>	<b>Section</b>	
• "RVS_TEB"	4.2.31	
• "RVS_BB_SV_Frame_No"	4.2.32	
• "bb_t_sat_switch_aqua"	4.2.41	<i>MODIS/Aqua (FMI) only</i>
• "bb_t_sat_aqua"	4.2.42	<i>MODIS/Aqua (FMI) only</i>
• "bb_t_sat_default_b1_aqua"	4.2.43	<i>MODIS/Aqua (FMI) only</i>
<b>Emissive LUTs deleted:</b>	<b>Reason</b>	
• "RVS_1km_Emiss_SV"	Replaced by RVS_TEB	
• "RVS_1km_Emiss_BB"	Replaced by RVS_TEB	
• "RVS_1km_Emiss_EV"	Replaced by RVS_TEB	
• "delta_DN_TEB"	ADC Correction eliminated	
<b>Emissive LUTs changed:</b>	<b>Section</b>	<b>Nature of Change</b>
• "RSR"	4.2.8	Dimensions changed for <i>MODIS/Aqua (FMI) only</i>

**Table 1.3: Changes to the Quality Assurance LUT Collection and Structure Since April 2001**

<b>QUALITY ASSURANCE LUT CHANGES</b>		
<b>Quality Assurance LUTs added:</b>	<b>Section</b>	
• "Spacecraft_Roll_Threshold_Angle"	5.2.24	
• "Spacecraft_Pitch_Threshold_Angle"	5.2.25	
• "Spacecraft_Yaw_Threshold_Angle"	5.2.26	
<b>Quality Assurance LUTs deleted:</b>	<b>Reason</b>	
• "PROCESSINGCENTER"	Determination moved to Process Control File to allow Direct Broadcast code users to set as they wish.	

Detailed LUT histories and complete information regarding which LUTs are currently being used in MODIS Level 1B data production at the Goddard DAAC are available at the MCST Level 1B Product Information and Status web page (<http://www.mcst.ssai.biz/mcstweb/L1B/product.html>).

## 2 GENERAL INFORMATION, CONVENTIONS, AND VERSIONING STRATEGY

### 2.1 LUT HDF files, LUT Type and Time-Dependent LUTs

The LUTs are organized into three groups:

- Reflective Lookup Tables (Section 3)
- Emissive Lookup Tables (Section 4)
- QA Lookup Tables (Section 5)

There is one Hierarchical Data Format (HDF) file per group, for a total of three LUT files. The three HDF files are treated as a single set of LUTs. The strategy for versioning a set of LUTs is described in Section 2.3.

An individual LUT is implemented in the appropriate HDF file as either:

- A global attribute, or
- A Scientific Data Set (SDS).

This implementation is required by the Level 1B code, which additionally defines the data type and intrinsic LUT dimensions for each LUT. The meaning of "intrinsic" dimensions will become apparent after the next few paragraphs. Intrinsic LUT dimensions are described for each LUT in later sections of this document.

Any LUT that is implemented as an SDS may have a dependence on the data collection time. The type of time dependence is defined by the attribute "algorithm", which is attached to the SDS. The "algorithm" attribute is a scalar, int32 value. There are currently three allowable values of this attribute:

LUT Type	"algorithm" value
Constant	0
Step Function	1
Piecewise Linear	2

A "constant" LUT contains one set of data to be applied within the Level 1B code regardless of the data collection time of the Level 1A data. The array structure of the constant LUT is the same as the "intrinsic" dimensions described for each LUT later in this document. This array structure matches the way that the values are ingested, stored and used within the Level 1B code.

A "step function" LUT can be thought of as the concatenation of a series of constant LUTs, each of which has an effective beginning time associated with it. If an SDS LUT is a step function LUT, it will have an additional leading dimension for the array and an additional SDS attribute named "times" — a float64 array of numbers containing the beginning TAI times for the data sets

contained in the SDS array. For each time, there is an equivalent "constant" LUT contained in the SDS array. The size of the additional leading dimension of the SDS is the same as the size of the "times" attribute.

"Piecewise linear" time dependence is similar to step function dependence except that the LUT values are linearly interpolated from two of the data sets in the file, where the center time of a granule determines the data collection time to which to interpolate the LUT values. If the data collection time is before the first LUT time or after the last LUT time, then the code extrapolates linearly using the first two or last two data sets as appropriate. Currently, only LUTs that have data types of float32 or float64 may have piecewise linear time dependence.

Note that the values of the attached attributes of the SDS LUT determine whether the LUT is constant or time dependent. This implies that a constant LUT may be changed to a time dependent LUT through a LUT update (or vice versa). A code change is not required. Also, LUT updates may add, delete, or change time stamped table pieces.

## 2.2 Conventions Implemented in the Level 1B HDF LUTs

### 2.2.1 Band Groupings and Ordering

- The full set of MODIS bands (NUM\_BANDS = 38):  
1, 2, 3, ..., 12, 13lo, 13hi, 14lo, 14hi, 15, ... 34, 35, 36
- The reflective Solar bands (NUM\_REFLECTIVE\_BANDS = 22):  
1, 2, 3, ..., 12, 13lo, 13hi, 14lo, 14hi, 15, 16, 17, 18, 19, 26
- The thermal emissive bands (NUM\_EMISSIVE\_BANDS = 16):  
20, 21, 22, 23, 24, 25, 27, 28, 29, 30, 31, 32, 33, 34, 35, 36
- The 250m resolution bands (NUM\_250M\_BANDS = 2):  
1, 2
- The 500m resolution bands (NUM\_500M\_BANDS = 5):  
3, 4, 5, 6, 7
- The 1km reflective bands (NUM\_1000M\_REFL\_BANDS = 15):  
8, 9, 10, 11, 12, 13lo, 13hi, 14lo, 14hi, 15, 16, 17, 18, 19, 26
- The SWIR Bands (NUM\_SWIR\_BANDS = 4):  
5, 6, 7, 8
- Thermal emissive bands on MODIS/AQUA which saturate on blackbody warmup (NUM\_AQUA\_BB\_SAT\_BANDS = 3):  
33, 35, 36

### 2.2.2 Intrinsic Ordering of Detectors Within a Band

Within the LUT HDF files, the ordering follows the "product" convention, which is inverted from the SBRS detector layout convention. *Note that the LUT ASCII files are delivered to the software team in "SBRS" detector order.* The software team inverts the detector order inherent in the delivered ASCII files so that the result in the HDF LUT files is "product" order.

### 2.2.3 Intrinsic Order in a Multidimensional Array

A multidimensional array has an intrinsic ordering of elements in the way that the numbers are stored in memory. All L1B LUTs, regardless of dimensionality, are delivered to the software team

in an ASCII file containing a stream of numbers, usually (but not required) one number per record. The stream of numbers should be in the order that they would take in the memory of the computer. In the description of LUTs later in this document, the convention for describing the enumeration of the individual dimensions of a multidimensional array follows the C language convention in that the first dimension is the *least-rapidly varying* and the last dimension is the *most-rapidly varying*.

### 2.3 Strategy for Versioning the Level 1B Code and LUTs

Each LUT file contains the following three data items, which convey versioning information:

- "Serial Number" — formerly conveyed version information about the science content of each file independently of the other files. Presently is unique to the "MCST version" (see below).
- "PGE version" — conveys the version of the Level 1B code itself
- "MCST version" or "Algorithm Package Version" — conveys the version of the three LUT files as a set relative to the PGE version.

Specific formats for the above versions are described in the details in later sections. Here, we present some general comments regarding the meaning of these data items.

The "Serial Number" was formerly used to convey information about the science content of the LUT file. In this regard, the Serial Numbers in the three LUT files were completely independent from each other. A change to values of one of the science LUTs in one file would cause that Serial Number to be updated while the Serial Numbers of the other files remained unchanged. Beginning in mid-2002, the policy was changed so that Serial Numbers are unique to the "MCST version" (see below). This change was made to avoid confusion between different PGE02 versions and LUT updates.

The "PGE\_version" LUT, present in each LUT HDF file, contains a copy of the version of the Level 1B code itself. This version is placed in the ECS core metadata field "PGEVersion". Within the L1B code, the PGE version is hard-coded in the macro "PGE02\_VERSION". When the code changes, the PGE version must also be changed. The code will check that the PGE version set in the LUT file matches the code macro. This will help prevent out-of-date LUT files from being used with a given release of the code. In this regard, the three LUT files form a set that all must have the same PGE version to be valid for a given release of the code.

The "MCST\_version" conveys the version of the LUT files as a set relative to a specific release of the code. This value is placed into the product in the ECS archive metadata field "AlgorithmPackageVersion". Within the L1B products, this version is the most complete single version that describes the calibration used for that particular data set. The code checks the MCST version against an MCST version supplied by the user in the product control file (PCF). This also prevents incorrect or out-of-date LUTs from being used with MOD\_PR02. The MCST\_version LUT is placed in each LUT HDF file and all must agree with each other for the set to be valid. When a LUT update is delivered to the GDAAC, the MCST version changes but the PGE version remains the same. In this situation, all three LUT files will be supplied to the GDAAC as a set even if the science content of only one LUT file actually changed.

### 3 REFLECTIVE LUTS

#### 3.1 Change to “BDSM” LUT structure for MOD\_PR02 Versions 4.0.7, 4.1.1 and higher

A major change to the way that certain reflective LUTs are delivered to the L1B software team and stored in the Reflective LUT HDF file was made for MOD\_PR02 Version 4.0.7 (Terra) and 4.1.1 (Aqua). Those reflective LUTs which are indexed by band, detector, subframe, and mirror side (“BDSM” LUTs) are now delivered and stored without fill values. For each such LUT (or each table piece if the LUT is a time-dependent one), the dimensions of the LUT used within L1B execution remain the same ( $[\text{NUM\_REFLECTIVE\_BANDS}] * [\text{MAX\_DETECTORS\_PER\_BAND}] * [\text{MAX\_SAMPLES\_PER\_BAND}] * [\text{NUM\_MIRROR\_SIDES}]$ ). Each table piece is delivered and stored in the HDF file with dimensions NUM\_REFLECTIVE\_INDICES, where

$$\begin{aligned} \text{NUM\_REFLECTIVE\_INDICES} = & \\ & [\text{NUM\_250M\_BANDS} * \text{DETECTORS\_PER\_250M\_BAND} * \\ & \quad \text{SAMPLES\_PER\_250M\_BAND} * \text{NUM\_MIRROR\_SIDES}] + \\ & [\text{NUM\_500M\_BANDS} * \text{DETECTORS\_PER\_500M\_BAND} * \\ & \quad \text{SAMPLES\_PER\_500M\_BAND} * \text{NUM\_MIRROR\_SIDES}] + \\ & [\text{NUM\_1000M\_BANDS} * \text{DETECTORS\_PER\_1KM\_BAND} * \\ & \quad \text{SAMPLES\_PER\_1KM\_BAND} * \text{NUM\_MIRROR\_SIDES}] \end{aligned}$$

LUTs affected by this change were the following:

LUT	Section
• “K_inst”	3.3.4
• “K_FPA”	3.3.5
• “m0 “	3.3.6
• “m1”	3.3.7
• “Sigma_m1”	3.3.11
• “Sigma_K_inst”	3.3.12
• “RSB_NEdL”	3.3.16
• “dn_sat_ev”	3.3.32

### 3.2 Summary of Reflective LUTs and their Dimensions

Table 3.1 Summary of Reflective Calibration Lookup Tables

LUT name	Section	Meaning
Serial Number of Reflective LUT	3.3.1	Version number of reflective calibration LUTs
PGE_Version	3.3.2	3-digit PGE version number
MCST_Version	3.3.3	4-digit ALGORITHMPACKAGEVERSION
K_inst	3.3.4	Instrument Temperature Correction Factor
K_FPA	3.3.5	Focal Plane Array Temperature Correction Factor
m0	3.3.6	Reflectance Calibration Factors (constant and linear terms)
m1	3.3.7	
RVS_RefSB	3.3.8	RVS (Response vs. Scan Angle) coefficients to calculate correction for the Reflective Solar Bands.
Sigma_RVS_RSB	3.3.9	Uncertainty in RVS at nadir frame
Sigma_RSB_ADC	3.3.10	Uncertainty in the RSB ADCs
Sigma_m1	3.3.11	Uncertainty in m1
Sigma_K_inst	3.3.12	Uncertainty in the instrument temperature correction factor
Sigma_T_inst	3.3.13	Uncertainty in the instrument temperature
Sigma_PV_Resid_Elec	3.3.14	Uncertainty related to electrical cross-talk
Sigma_R_Star_Lin_Resid_Ucoeff	3.3.15	Uncertainty related to deviations from linear behavior in R*
RSB_NEdL	3.3.16	RSB Noise equivalent delta radiances
T_inst_ref	3.3.17	Instrument temperature reference
T_FPA_ref	3.3.18	Focal Plane temperature reference
SWIR_OOB_correction_switch	3.3.19	A switch (0=OFF, 1=ON) for the SWIR OOB leak correction.
SWIR_OOB_corr_sending_band	3.3.20	Number of Emissive band to use as "sending" band for SWIR OOB leak correction.
X_OOB_0	3.3.21	Coefficients of quadratic SWIR band correction formula (zero-, first-, and second-order coefficients respectively).
X_OOB_1	3.3.22	
X_OOB_2	3.3.23	
DN_obc_avg_first_frame_to_use	3.3.24	Index of the first frame to use for computing average OBC DN.
DN_obc_avg_number_of_frames_to_use	3.3.25	Number of frames to use for computing average OBC DN.
Dn_star_Max	3.3.26	Maximum dn** value for scaling to the product scaled integer
Dn_star_Min	3.3.27	Minimum dn** value for scaling to the product scaled integer
RSB_specified_uncertainty	3.3.28	Factor used in computing uncertainty index
RSB_UI_scaling_factor	3.3.29	Factor used in computing uncertainty index
E_sun_over_pi	3.3.30	RSR-weighted Solar irradiance/pi for RSB detectors
RSB_SV_DN_moon_include_frames	3.3.31	Number of frames after sorting if moon in SVP
Dn_sat_ev	3.3.32	Value of EV pixel dn to treat as saturated
B26_B5_Corr_Switch	3.3.33	Switch (0=OFF, 1=ON) for correction of Band 26 by Band 5 aggregated values.
B26_B5_Frame_Offset	3.3.34	Frame offset to use when applying correction of Band 26 by Band 5.
B26_B5_Corr	3.3.35	Correction factor to be applied to Band 26 data.

**Table 3.2 Dimensions used in Reflective LUTs.**

<b>Dimension macro</b>	<b>value</b>	<b>Meaning</b>
<b>MAX_MCST_VERSION_BUFFER</b>	<b>21</b>	<b>Largest allowable string length</b>
<b>MAX_PGE_VERSION_BUFFER</b>	<b>11</b>	<b>Largest allowable string length</b>
<b>MAX_SERIAL_NUMBER_BUFFER</b>	<b>31</b>	<b>Largest allowable string length</b>
<b>NUM_REFLECTIVE_BANDS</b>	<b>22</b>	<b>Number of reflective Solar bands.</b>
<b>MAX_DETECTORS_PER_BAND</b>	<b>40</b>	<b>Maximum number of detectors per band.</b>
<b>MAX_SAMPLES_PER_BAND</b>	<b>4</b>	<b>Maximum number of subsamples per 1km frame.</b>
<b>NUM_MIRROR_SIDES</b>	<b>2</b>	<b>Number of mirror sides.</b>
<b>NUM_FOCAL_PLANES</b>	<b>4</b>	<b>Number of focal plane assemblies (FPAs).</b>
<b>NUM_SWIR_BANDS</b>	<b>4</b>	<b>Number of SWIR bands</b>
<b>MAX_DETECTORS_PER_SWIR_BAND</b>	<b>20</b>	<b>Max. # of detectors in a SWIR band.</b>
<b>MAX_NUM_SWIR_SUBSAMPLES</b>	<b>2</b>	<b>Max. # of subsamples in a SWIR band</b>
<b>NUM_REFL_INDICES</b>	<b>1340</b>	<b>Total over all reflective bands of number detectors per band*number samples per band * number mirror sides</b>
<b>NUM_REFLECTIVE_DETECTORS</b>	<b>330</b>	<b>Total number of reflective band detectors</b>
<b>NUM_2ND_ORDER_COEFFS</b>	<b>3</b>	<b>Number of coefficients in an order 2 polynomial</b>
<b>NUM_4TH_ORDER_COEFFS</b>	<b>5</b>	<b>Number of coefficients in an order 4 polynomial</b>

### 3.3 Reflective LUT listing

---

#### 3.3.1 Serial Number of Reflective LUT

This serial number serves to identify the "science content" version of the reflective lookup tables. It is stored as a string and has the form: "Rvvv yyyy:MM:dd:hh:mm", where "R" is for reflective, "vvv" is an integer version number that gets incremented each time values of any LUT other than PGE\_Version and MCST\_Version are changed, and "yyyy:MM:dd:hh:mm" is the date and time of the last change to any LUT other than PGE\_Version and MCST\_Version. This serial number is updated as lookup tables are modified during operations. It is placed in each of the L1B Earth-view products.

Name: "Serial Number of Reflective LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_SERIAL\_NUMBER\_BUFFER  
Range: N/A  
Fill Value: (none)

---

#### 3.3.2 PGE Version

This LUT contains the value of PGEVERSION, written to the ECS core metadata in each of the L1B products. The format consists of "A.B.C". The PGE version is currently synonymous with the version of the Level 1B code itself. This value only changes if there is an actual change to the code. This value is present as a LUT in the Reflective LUTs file only as a safety check. The value must match the internal code macro for the PGE version for the LUT file to be valid. Otherwise, the code will error-out. Changes in the value of PGE version do not necessarily cause a change in the Serial Number, described above.

Name: "PGE Version LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_PGE\_VERSION\_BUFFER  
Range: N/A  
Fill Value: (none)

---

---

### 3.3.3 MCST Version

This LUT contains the MCST version of Level 1B algorithms, code and LUTs. It is placed in the value of ALGORITHMPACKAGEVERSION, written to the ECS core metadata in each of the L1B products. This value is updated with each change in Level 1B, including any LUT change in any file (not just in the Reflective LUTs file). The format is a four-digit number placed in a string. The first three digits match the PGE\_Version, described above. The fourth digit will indicate the LUT update value since the last code release. Also, the string "\_Terra" or "\_Aqua" will be appended to the LUTs to distinguish between the different satellites. This value is placed in all three LUT files and all must agree for the set of LUT files to be valid. Otherwise, the code will error-out.

Name:	"MCST Version LUT"
Kind:	Global Attribute
Data type:	string
Rank:	1
Dimension:	MAX_MCST_VERSION_BUFFER
Range:	N/A
Fill Value:	(none)

---

### 3.3.4 *K\_inst*

#### Instrument Temperature Correction Factor

This value is used to generate a correction factor to DN. This correction factor accounts for the small dependence of detector responsivity on variations in instrument temperature.

Name:	"K_inst"		
Kind:	SDS	Step Function LUT	
Data type:	float32		
Intrinsic Rank:		L1B Execution*	LUT HDF File*
Dimensions:	dim1:	4	1
	dim2:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim3:	MAX_DETECTORS_PER_BAND	
	dim4:	MAX_SAMPLES_PER_BAND	
Range:	[-0.1, 0.1]		
Fill Value:	-999		
		* See Section 3.1	

### 3.3.5 *K\_FPA*

#### Focal Plane Array Temperature Correction Factor

This value is used to generate a correction factor to DN. This correction factor accounts for the small dependence of detector responsivity on variations in FPA temperature.

Name:	"K_FPA"		
Kind:	SDS	Constant LUT	
Data type:	float32		
Rank:	4	L1B Execution*	LUT HDF File*
Dimensions:	dim1:	1	NUM_REFL_INDICES
	dim2:	NUM_REFLECTIVE_BANDS	
	dim3:	MAX_DETECTORS_PER_BAND	
	dim4:	MAX_SAMPLES_PER_BAND	
Range:	[-0.1, 0.1]		
Fill Value:	-999		
		* See Section 3.1	

### 3.3.6 *m0*

#### Reflectance Calibration Factor – Offset Term

These factors, along with the Earth-Sun distance, are used to convert dn\* to reflectance. These LUTs are also used in Level 1B to compute the product scaled integers and the reflectance calibration scales and offsets.

Name:	"m0"		
Kind:	SDS	Constant LUT	
Data type:	float32		
		L1B Execution*	LUT HDF File*
Rank:	4	1	
Dimensions:	dim1:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim2:	MAX_DETECTORS_PER_BAND	
	dim3:	MAX_SAMPLES_PER_BAND	
	dim4:	NUM_MIRROR_SIDES	
Range:	[0.0, 1.0]		
Fill Value:	-999		

\* See Section 3.1

### 3.3.7 *m1*

#### Reflectance Calibration Factor – Linear Term

These factors, along with the Earth-Sun distance, are used to convert dn\* to reflectance. These LUTs are also used in Level 1B to compute the product scaled integers and the reflectance calibration scales and offsets.

Name:	"m1"		
Kind:	SDS	Piecewise Linear LUT	
Data type:	float32		
		L1B Execution*	LUT HDF File*
Intrinsic Rank:		4	1
Dimensions:	dim1:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim2:	MAX_DETECTORS_PER_BAND	
	dim3:	MAX_SAMPLES_PER_BAND	
	dim4:	NUM_MIRROR_SIDES	
Range:	[0.0, 1.0]		
Fill Value:	-999		

\* See Section 3.1

---

### 3.3.8 *RVS\_RefSB*

This LUT holds the 2<sup>nd</sup> order polynomial coefficients used to compute the response vs. scan angle (RVS) as a function of frame number for the reflective Solar bands. Polynomial evaluation will result in a normalized value of 1.0 at the center frame of the Solar diffuser sector (approximately 50.25 degrees).

Name: "RVS\_RefSB"  
Kind: SDS Piecewise Linear LUT  
Data type: float32  
Intrinsic Rank: 4  
Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
dim2: MAX\_DETECTORS\_PER\_BAND  
dim3: NUM\_MIRROR\_SIDES  
dim4: NUM\_2ND\_ORDER\_COEFFS  
Range: [-2.0E-4, 1.2]  
Fill Value: -999

---

### 3.3.9 *Sigma\_RVS\_RSB*

This LUT holds the uncertainty in the response vs. scan angle (RVS) for the reflective Solar bands at the nadir frame. Units are absolute (same units as calculated RVS).

Name: "Sigma\_RVS\_RSB"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 2  
Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
dim2: NUM\_MIRROR\_SIDES  
Range: [0, 0.03]  
Fill Value: (none)

---

### 3.3.10 *Sigma\_RSB\_ADC*

Uncertainties in reflective Solar band Analog to Digital Converter (ADC).

Name: "Sigma\_RSB\_ADC"  
 Kind: SDS Constant LUT  
 Data type: float32  
 Rank: 2  
 Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
 dim2: MAX\_DETECTORS\_PER\_BAND  
 Range: N/A  
 Fill Value: -999

### 3.3.11 *Sigma\_m1*

Reflectance Calibration Factor Uncertainties

This LUT contains the reflectance calibration factor (m1) uncertainties. These values are calculated off-line and are used in the calculation of uncertainty indices for reflective Solar bands. Units are absolute (same units as m1).

Name: "Sigma\_m1"  
 Kind: SDS Piecewise Linear LUT  
 Data type: float32  
 Intrinsic Rank: L1B Execution\* 4 LUT HDF File\* 1  
 Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
 dim2: MAX\_DETECTORS\_PER\_BAND  
 dim3: MAX\_SAMPLES\_PER\_BAND  
 dim4: NUM\_MIRROR\_SIDES  
 Range: [0.0, 1.0]  
 Fill Value: -999

\* See Section 3.1

### 3.3.12 *Sigma\_K\_inst*

These values are calculated off-line and are the uncertainties of the factor that corrects for effects of instrument temperature on detector responsivity. Units are absolute (same units as K\_inst).

Name:	"Sigma_K_inst"		
Kind:	SDS	Step Function LUT	
Data type:	float32		
Intrinsic Rank:		L1B Execution*	LUT HDF File*
Dimensions:	dim1:	4	1
	dim2:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim3:	MAX_DETECTORS_PER_BAND	
	dim4:	MAX_SAMPLES_PER_BAND	
		NUM_MIRROR_SIDES	
Range:	[0, 0.01]		
Fill Value:	-999		

\* See Section 3.1

### 3.3.13 *Sigma\_T\_inst*

This value is the uncertainty in instrument temperature. Units are degrees.

Name:	"Sigma_T_inst"	
Kind:	SDS	Constant LUT
Data type:	float32	
Rank:	1	
Dimension:	1	
Range:	N/A	
Fill Value:	(none)	

---

### 3.3.14 *Sigma\_PV\_Resid\_Elec*

These values are calculated off-line and are the uncertainties ascribed to residual electronic cross-talk. Units are dn.

Name: "Sigma\_PV\_Resid\_Elec"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 3  
Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
dim2: MAX\_DETECTORS\_PER\_BAND  
dim3: MAX\_SAMPLES\_PER\_BAND  
Range: N/A  
Fill Value: -999

---

### 3.3.15 *Sigma\_R\_Star\_Lin\_Resid\_Ucoeff*

This LUT contains 4<sup>th</sup> order polynomial fits of uncertainty vs. reflectance for the residual of the R\* from a linear fit.

Name: "Sigma\_R\_Star\_Lin\_Resid\_Ucoeff"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 5  
Dimensions: dim1: NUM\_REFLECTIVE\_BANDS  
dim2: MAX\_DETECTORS\_PER\_BAND  
dim3: MAX\_SAMPLES\_PER\_BAND  
dim4: NUM\_MIRROR\_SIDES  
dim5: NUM\_4TH\_ORDER\_COEFFS  
Range: N/A  
Fill Value: -999

---

---

### 3.3.16 RSB\_NEdL

These values are calculated off-line and are the noise equivalent delta radiances of the reflective Solar bands.

Name:	"RSB_NEdL"		
Kind:	SDS	Step Function LUT	
Data type:	float32		
Intrinsic Rank:		L1B Execution*	LUT HDF File*
Dimensions:	dim1:	4	1
	dim2:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim3:	MAX_DETECTORS_PER_BAND	
	dim4:	MAX_SAMPLES_PER_BAND	
		NUM_MIRROR_SIDES	
Range:	[0, 1]		
Fill Value:	-999		

\* See Section 3.1

---

### 3.3.17 T\_inst\_ref

Instrument temperature reference

This value is the instrument reference temperature (in degrees K) to use when correcting DN for effects of instrument temperature on detector responsivity.

Name:	"T_inst_ref"		
Kind:	SDS	Step Function LUT	
Data type:	float32		
Intrinsic Rank:	1		
Dimension:	1		
Range:	[250, 300]		
Fill Value:	(none)		

---

### 3.3.18 *T\_FPA\_ref*

Focal Plane temperature reference

These values are the focal plane assembly (FPA) reference temperatures (in degrees K) to use when correcting DN for effects of FPA temperature on detector responsivity.

Name: "T\_FPA\_ref"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_FOCAL\_PLANES  
Range: [80, 300]  
Fill Value: (none)

Note: the order of the focal planes in the array is: VIS, NIR, SMIR, LWIR.

---

### 3.3.19 *SWIR\_OOB\_correction\_switch*

This LUT defines whether the SWIR out-of-band (OOB) leak correction to dn is applied. If this correction switch is set to 0 (OFF), then the correction will not be applied regardless of the values in the LUTs relating to this correction (see X\_OOB\_0, Section 3.3.21, X\_OOB\_1, Section 3.3.22, and X\_OOB\_2, Section 3.3.23).

Name: "SWIR\_OOB\_correction\_switch"  
Kind: SDS Step Function LUT  
Data type: int16  
Intrinsic Rank: 1  
Dimension: 1  
Range: [0, 1]  
Fill Value: (none)

---

### 3.3.20 SWIR\_OOB\_corr\_sending\_band

This LUT defines the number of the “sending band” for the SWIR out-of-band (OOB) leak correction. It is normally set to Band 25 for MODIS/Aqua and Band 28 for MODIS/Terra. See X\_OOB\_0, Section 3.3.21, X\_OOB\_1, Section 3.3.22, and X\_OOB\_2, Section 3.3.23.

Name: "SWIR\_OOB\_corr\_sending\_band "  
 Kind: SDS Step Function LUT  
 Data type: int16  
 Intrinsic Rank: 1  
 Dimension: 1  
 Range: [20, 36]  
 Fill Value: (none)

### 3.3.21 X\_OOB\_0

This LUT contains the constant coefficient of the quadratic relation empirically derived between the SWIR OOB correction sending band (“Band X”, see Section 3.3.20 above)  $dn$ ,  $dn_x$ , and the correction to the SWIR band,  $dn_{corr}$ . The correction is applied as:

$$dn = dn - dn_{corr} \quad \text{where}$$

$$dn_{corr} = X^{OOB}_0 + X^{OOB}_1 * dn_x + X^{OOB}_2 * dn_x^2$$

Name: "X\_OOB\_0"  
 Kind: SDS Constant LUT  
 Data type: float32  
 Rank: 4  
 Dimensions: dim1: NUM\_SWIR\_BANDS  
 dim2: MAX\_DETECTORS\_PER\_SWIR\_BAND  
 dim3: MAX\_NUM\_SWIR\_SUBSAMPLES  
 dim4: NUM\_MIRROR\_SIDES  
 Range: [-100, 100]  
 Fill Value: -999

### 3.3.22 X\_OOB\_1

This LUT contains the first order coefficient of the quadratic relation empirically derived between the SWIR OOB correction sending band (“Band X”)  $dn$ ,  $dn_x$ , and the correction to the SWIR band,  $dn_{corr}$ . See X\_OOB\_0 above (Section 3.3.21) for how the correction is applied.

Name: "X\_OOB\_1"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 4  
Dimensions: dim1: NUM\_SWIR\_BANDS  
dim2: MAX\_DETECTORS\_PER\_SWIR\_BAND  
dim3: MAX\_NUM\_SWIR\_SUBSAMPLES  
dim4: NUM\_MIRROR\_SIDES  
Range: [-100, 100]  
Fill Value: -999

### 3.3.23 X\_OOB\_2

This LUT contains the second order coefficient of the quadratic relation empirically derived between the SWIR OOB correction sending band (“Band X”)  $dn$ ,  $dn_x$ , and the correction to the SWIR band,  $dn_{corr}$ . See X\_OOB\_0 above (Section 3.3.21) for how the correction is applied.

Name: "X\_OOB\_2"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 4  
Dimensions: dim1: NUM\_SWIR\_BANDS  
dim2: MAX\_DETECTORS\_PER\_SWIR\_BAND  
dim3: MAX\_NUM\_SWIR\_SUBSAMPLES  
dim4: NUM\_MIRROR\_SIDES  
Range: [-100, 100]  
Fill Value: -999

---

### 3.3.24 *DN\_obc\_avg\_first\_frame\_to\_use*

Name: "DN\_obc\_avg\_first\_frame\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 49]  
Fill Value: (none)  
Value: Presently set to 10.

---

### 3.3.25 *DN\_obc\_avg\_number\_of\_frames\_to\_use*

This LUT holds the index (0 through N-1) of the first frame to use in computing the average OBC DN, used as a zero-point value in the reflective Solar band calibration algorithms. The space-view (SV) DN is typically used in calculating this number. However, if the moon is in the SV keep-out-box or there is some other problem with the SV data, the blackbody (BB) DNs are used.

Name: "DN\_obc\_avg\_number\_of\_frames\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [1, 50]  
Fill Value: (none)  
Value: Presently set to 30.

---

---

### 3.3.26 *dn\_star\_Max*

This value defines the upper limit of the dynamic range of  $dn^{**}$  for the purpose of scaling to the product scaled integer. The name of this is somewhat misleading since we scale  $dn^{**}$ , not  $dn^*$ , to the product scaled integer.

Name: "dn\_star\_Max"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_REFLECTIVE\_BANDS  
Range: [4095, 4095]  
Fill Value: (none)  
Values: Derived pre-launch. The value for each band is 4095.

---

### 3.3.27 *dn\_star\_Min*

This value defines the lower limit of the dynamic range of  $dn^{**}$  for the purpose of scaling to the product scaled integer. (See other comments above for *dn\_star\_Max*).

Name: "dn\_star\_Min"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_REFLECTIVE\_BANDS  
Range: [-40, 0]  
Fill Value: (none)  
Values: Derived pre-launch. For Bands 1-7, the value is 0 (zero). For other bands, the value is -40. If the nadir aperture door is closed on any scan in the granule, then the values in this LUT for Bands 1-7 will be dynamically over-written by the L1B code and re-set to -40.

---

---

### 3.3.28 *RSB\_specified\_uncertainty*

This LUT contains the specified uncertainty factor for each reflective Solar band, which is used along with the scaling factor described by the LUT "RSB\_UI\_scaling\_factor" (see Section 3.3.29) to convert percent uncertainty to an uncertainty index.

Name: "RSB\_specified\_uncertainty"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_REFLECTIVE\_BANDS  
Range: N/A  
Fill Value: (none)

---

### 3.3.29 *RSB\_UI\_scaling\_factor*

This LUT contains the uncertainty scaling factor for each reflective Solar band, which is used along with the calculated uncertainty described by the LUT "RSB\_specified\_uncertainty" (see Section 3.3.28) to convert percent uncertainty to an uncertainty index. The formula for uncertainty index (UI) in terms of uncertainty\_in\_percent, specified\_uncertainty, and scaling\_factor is:

$$UI = \text{scaling\_factor} * \ln (\text{uncertainty\_in\_percent} / \text{specified\_uncertainty})$$

where "ln" is the natural logarithm. The uncertainty in percent is computed within the Level 1B code dynamically.

Name: "RSB\_UI\_scaling\_factor"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_REFLECTIVE\_BANDS  
Range: N/A  
Fill Value: (none)

---

---

### 3.3.30 *E\_sun\_over\_pi*

For each RSB detector, this LUT contains the Solar irradiance at 1 AU divided by and weighted by the detector's relative spectral response (RSR).

Name: "E\_sun\_over\_pi"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_REFLECTIVE\_DETECTORS  
Range: N/A  
Fill Value: (none)

---

### 3.3.31 *RSB\_SV\_DN\_moon\_include\_frames*

This LUT holds the number of frames to use when calculating average space-view (SV) DN when the moon is in the SV keep-out box (KOB). If the moon is determined to be in the SV KOB, the 50 SV DN frames are sorted from the lowest to the highest values. The lowest "RSB\_SV\_DN\_moon\_include\_frames" are used to calculate average SV DN using the same logic as is done for the case when the moon is not in the SV KOB.

Name: "RSB\_SV\_DN\_moon\_include\_frames"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 50]  
Fill Value: (none)

---

### 3.3.32 *dn\_sat\_ev*

Value of EV pixel dn to treat as saturated

For some RSB detectors, saturation occurs before the ADC saturates at DN=4095. This premature saturation is best detected by using the space-view subtracted digital number, dn. For those cases, the value of this LUT will be less than 4095. If no premature saturation occurs for a detector, then the LUT values are set to 4095.

Name:	"dn_sat_ev"		
Kind:	SDS	Step Function LUT	
Data type:	float64		
		L1B Execution*	LUT HDF File*
Intrinsic Rank:		4	1
Dimensions:	dim1:	NUM_REFLECTIVE_BANDS	NUM_REFL_INDICES
	dim2:	MAX_DETECTORS_PER_BAND	
	dim3:	MAX_SAMPLES_PER_BAND	
	dim4:	NUM_MIRROR_SIDES	
Range:	[0, 4095]		
Fill Value:	-999		

\* See Section 3.1

### 3.3.33 *B26\_B5\_Corr\_Switch*

This LUT defines whether the Band 5 correction to Band 26 scaled integers is applied. If this correction switch is set to 0 (OFF), then the correction will not be applied regardless of the values in the next two LUTs relating to this correction. See "B26\_B5\_Corr" (Section 3.3.35) for the correction formula. This LUT was originally added for MODIS/Terra only but has subsequently been inserted for MODIS/Aqua as well.

Name:	"B26_B5_Corr_Switch"	
Kind:	SDS	Step Function LUT
Data type:	int16	
Intrinsic Rank:	1	
Dimension:	1	
Range:	[0, 1]	
Fill Value:	(none)	

---

### 3.3.34 B26\_B5\_Frame\_Offset

This LUT gives the frame offsets used in calculating the Band 5 correction to Band 26 scaled integers (SIs). See “B26\_B5\_Corr” (Section 3.3.35) for the correction formula. This LUT was originally added for MODIS/Terra only but has subsequently been inserted for MODIS/Aqua as well.

Name: “B26\_B5\_Frame\_Offset”  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: DETECTORS\_PER\_1KM\_BAND  
Range: [-10, 10]  
Fill Value: -999

---

### 3.3.35 B26\_B5\_Corr

This LUT gives detector-by-detector correction factors to be used for correction of the Band 26 scaled integers (SIs) by the the aggregated Band 5 SIs. This LUT was originally added for MODIS/Terra only but has subsequently been inserted for MODIS/Aqua as well. The correction formula is

$$SI_{26}(D, F)_{Corr} = SI_{26}(D, F) \cdot SI_{5_{Aggr}}(D, F + F_{Off}(D)) \cdot B26\_B5\_Corr(D) \cdot Rad\_Sc\_5 / Rad\_Sc\_26$$

where

SI<sub>26</sub>  
 =  
 Band 26 scaled integer after reflective Solar band calibration.

SI<sub>5<sub>Aggr</sub></sub>  
 =  
 Band 5 scaled integer aggregated to 1KM resolution after reflective Solar band calibration.

Name: B26\_B5\_Corr Step Function LUT  
 Kind: SDS  
 Data type: float32  
 Intrinsic Rank: 1  
 Detector Number: DETECTORS\_PER\_1KM\_BAND  
 Range: [0., 1.]  
 Fill Value: -999

=  
 Frame Number (0 - 1353)

F<sub>Off</sub>  
 =  
 Frame Offset from "B26\_B5\_Frame\_Offset\_Terra" above

Rad\_Sc\_5  
 =  
 The Band 5 Radiance scale (calculated dynamically in L1B)

Rad\_Sc\_26  
 =  
 The Band 26 Radiance scale (calculated dynamically in L1B)

B26\_B5\_Corr  
 =  
 The correction value from this table.

## 4 EMISSIVE LUTS

### 4.1 Summary of Emissive LUTs and their Dimensions

Table 4.1 Summary of Emissive Calibration Lookup Tables

LUT name	Section	Meaning
Serial Number of Emissive LUT	4.2.1	Version number of emissive calibration LUTs
PGE_Version	4.2.2	3-digit PGE version number
MCST_Version	4.2.3	4-digit version, placed in ALGORITHMPACKAGEVERSION
Epsilon_cav	4.2.4	Effective cavity emissivity.
epsilon_bb	4.2.5	Blackbody emissivity.
Delta_T_bb_beta	4.2.6	The " $\beta$ " term in the equation for calculating $\Delta T_{bb}$ .
Delta_T_bb_delta	4.2.7	The " $\Delta$ " term in the equation for calculating $\Delta T_{bb}$ .
RSR	4.2.8	Relative spectral responses
WAVELENGTH	4.2.9	Wavelengths at points of RSRs
NWL	4.2.10	Number of values in RSR distribution.
A0, A2	4.2.11, 4.2.12	Quadratic coefficients for calculating a0 and a2.
Ucoeff	4.2.13	Coefficients of polynomial fit of uncertainty weight vs. DN.
Sigma_TEB_PV_resid_elec	4.2.14	PV bands residual electronic cross-talk uncertainty
Sigma_TEB_ADC	4.2.15	ADC uncertainty
Ucoeff_Calibr_resid	4.2.16	Residual uncertainty to the calibration polynomial fit.
BB_DN_first_frame_to_use	4.2.17	Index of 1 <sup>st</sup> frame for computing BB DN averages
BB_DN_number_of_frames_to_use	4.2.18	Number of frames for computing BB DN averages
SV_DN_first_frame_to_use	4.2.19	Index of 1 <sup>st</sup> frame for computing SV DN averages
SV_DN_number_of_frames_to_use	4.2.20	Number of frames for computing SV DN averages
SV_DN_moon_include_frames	4.2.21	Number of frames after sorting if moon in SVP
Num_overlap_scans_b1	4.2.22	Number of scans in leading and trailing granules for cross-granule averaging of b1
T_ins_function_flag	4.2.23	Identifies suitable instrument temperature thermistors.
T_ins_default	4.2.24	Default value of instrument temperature in °K
T_ins_offset	4.2.25	Instrument temperature offset in °K.
T_cav_function_flag	4.2.26	Identifies suitable cavity temperature thermistors.
T_cav_default	4.2.27	Default value of cavity temperature in °K
T_mir_function_flag	4.2.28	Identifies suitable mirror temperature thermistors.
T_mir_default	4.2.29	Default value of mirror temperature in °K
BB_Weight	4.2.30	Weight factor used for computing average BB temperature.
RVS_TEB	4.2.31	Polynomial coefficients to calculate RVS correction for thermal Emissive Bands
Band_21_b1	4.2.33	The value of b1 for each Band 21 detector.
Band_21_Uncert_Lsat	4.2.34	Part of extra uncertainty term for Band 21.
L_Max	4.2.35	Top end of radiance dynamic range
L_Min	4.2.36	Bottom end of radiance dynamic range
TEB_specified_uncertainty	4.2.37	Factor used in computing uncertainty index
TEB_UI_scaling_factor	4.2.38	Factor used in computing uncertainty index
PC_XT	4.2.39	PC bands cross-talk correction parameters.
PCX_correction_switch	4.2.40	Switch (0=OFF, 1=ON) for crosstalk correction
BB_T_sat_switch_aqua	4.2.39	<i>MODIS/AQUA only.</i> On/off (1=ON; 0=OFF)switch for blackbody warmup saturation correction.
BB_T_sat_aqua	4.2.42	<i>MODIS/AQUA only.</i> Temperatures at which affected bands saturate on blackbody warmup.)
BB_T_sat_default_b1_aqua	4.2.43	<i>MODIS/AQUA only.</i> Default b1 values to use when blackbody temperature is over saturation values.



Table 4.2 Dimensions used in Emissive LUTs

Dimension macro	value	meaning
DETECTORS_PER_1KM_BAND	10	Number of detectors per 1km band
MAX_MCST_VERSION_BUFFER	21	Largest allowable string length
MAX_PGE_VERSION_BUFFER	11	Largest allowable string length
MAX_SERIAL_NUMBER_BUFFER	31	Largest allowable string length
MAX_NUM_RSR_vs_LAMBDA	49 (Terra) 66 (Aqua)	Maximum number of RSR vs. wavelength samples of an RSR distribution.
NUM_2ND_ORDER_COEFFS	3	Number of polynomial coefficients in 2 <sup>nd</sup> order polynomial.
NUM_4TH_ORDER_COEFFS	5	Number of polynomial coefficients in 4 <sup>th</sup> order polynomial.
NUM_a0_vs_T_inst_COEFF, NUM_a2_vs_T_inst_COEFF	3	Number of coefficients of a polynomial representation of a0 or a2 vs. instrument temperature.
NUM_AQUA_BB_SAT_BANDS	3	Number of thermal bands on MODIS/AQUA where saturation on blackbody warmup is observed.
NUM_BB_THERMISTORS	12	Number of blackbody thermistors.
NUM_EMISSIVE_BANDS	16	Number of emissive bands.
NUM_EMISSIVE_DETECTORS	160	Number of emissive detectors.
NUM_MIRROR_SIDES	2	Number of mirror sides.
NUM_PC_XT_BANDS	5	Number of bands affected by electronic cross-talk.
NUM_PC_XT_PARAMETERS	4	Number of electronic cross-talk parameters.
NUM_T_CAV_THERMISTORS	4	Number of thermistor points that may be used for determining cavity temperature.
NUM_T_INS_THERMISTORS	4	Number of thermistor points that may be used for determining instrument temperature.
NUM_T_MIR_THERMISTORS	2	Number of thermistor points that may be used for determining average mirror temperature.
NUM_UI_PARAMETERS	8	Number of algorithmic parameters influencing the uncertainty index.
NUM_UI_POLYNOMIAL_COEFF	2	Number of polynomial coefficients of fit vs. L.

## 4.2 Emissive LUT Listing

---

### 4.2.1 Serial Number of Emissive LUT

This serial number serves to identify the "science content" version of the emissive lookup tables. It is stored as a string and has the form: "Evvv yyyy:MM:dd:hh:mm", where "E" is for emissive, "vvv" is an integer version number that gets incremented each time values of any LUT other than PGE\_Version and MCST\_Version are changed, and "yyyy:MM:dd:hh:mm" is the date and time of the last change to any LUT other than PGE\_Version and MCST\_Version. This serial number is updated as lookup tables are modified during operations. It is placed in each of the L1B Earth-view products.

Name: "Serial Number of Emissive LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_SERIAL\_NUMBER\_BUFFER  
Range: N/A  
Fill Value: (none)

---

### 4.2.2 PGE Version

This LUT contains the value of PGEVERSION, written to the ECS core metadata in each of the L1B products (see amplified description in Section 3.3.2).

Name: "PGE Version LUT"  
Kind: Global Attribute  
Data type: String  
Rank: 1  
Dimension: MAX\_PGE\_VERSION\_BUFFER  
Range: N/A  
Fill Value: (none)

---

---

#### 4.2.3 MCST Version

This LUT contains the MCST version of Level 1B algorithms, code and LUTs (see amplified description in Section 3.3.3).

Name: "MCST Version LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_MCST\_VERSION\_BUFFER  
Range: N/A  
Fill Value: (none)

---

#### 4.2.4 *epsilon\_cav*

This lookup table holds values of the MODIS scan cavity effective emissivity,  $\epsilon_{cav}$ .

Name: "epsilon\_cav"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [0.5, 1.0]  
Fill Value: (none)

---

#### 4.2.5 *epsilon\_bb*

This lookup table holds values of the OBC blackbody emissivity,  $\epsilon_{bb}$ .

Name: "epsilon\_bb"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [0.9, 1.1]  
Fill Value: (none)

---

#### 4.2.6 *delta\_T\_bb\_beta*

This value represents " ", used for calculating a correction to the OBC blackbody temperature.

Name: "delta\_T\_bb\_beta"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [-0.5, 0.5]  
Fill Value: (none)

#### 4.2.7 *delta\_T\_bb\_delta*

This value represents the additional " " term at the end of the equation used for calculating a correction to the OBC blackbody temperature.

Name: "delta\_T\_bb\_delta"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [-0.5, 0.5]  
Fill Value: (none)

#### 4.2.8 *RSR*

This LUT holds a discrete approximation to the relative spectral response (RSR) curve for each of the emissive band detectors. Each emissive detector's RSR is given by these values for a given number of wavelengths (specified by the "NWL" LUT; see Section 4.2.10) and the wavelength values are given by the "Wavelength" LUT (see Section 4.2.9).

Name: "RSR"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 2  
Dimensions: dim1: NUM\_EMISSIVE\_DETECTORS  
dim2: MAX\_NUM\_RSR\_vs\_LAMBDA  
Range: [ 0.01, 1.0]  
Fill Value: 0.0

---

#### 4.2.9 WAVELENGTH

This lookup table holds the wavelength values used to compute the relative signal response (RSR) for each of the emissive detectors. (See LUTs “RSR”, Section 4.2.8, and “NWL”, Section 4.2.10).

Name: "WAVELENGTH"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 2  
Dimensions: dim1: NUM\_EMISSIVE\_DETECTORS  
dim2: MAX\_NUM\_RSR\_vs\_LAMBDA  
Range: [3.0, 15.5]  
(MODIS/AQUA)  
Fill Value: -999.  
[3.0, 15.0]  
(MODIS/TERRA)

---

#### 4.2.10 NWL

This lookup table holds number of relative signal responses (RSR) vs. wavelength values for each of the emissive detectors that L1B should actually use.. For example, if NWL is 20 for detector 100, L1B will use the first 20 wavelengths for detector 100 from the “Wavelength” LUT (see Section 4.2.9) and the first 20 RSR values for detector 100 from the “RSR” LUT (see Section 4.2.8) to compute the integrated RSR.

Name: "NWL"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [24, 67] (MODIS/AQUA)  
Fill Value: [14, 50] (MODIS/TERRA)

---

---

#### 4.2.11 A0

This LUTs holds the coefficients of polynomial fits of  $a_0$  vs. instrument temperature, where  $a_0$  is the constant coefficient of fits of radiance to dn.

Name: "A0"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 3  
Dimensions: dim1: NUM\_a0\_vs\_T\_inst\_COEFF  
dim2: NUM\_MIRROR\_SIDES  
dim3: NUM\_EMISSIVE\_DETECTORS  
Fill Value: (none)

---

#### 4.2.12 A2

This LUT holds the coefficients of polynomial fits of  $a_2$  vs. instrument temperature, where  $a_2$  is the quadratic coefficient of fits of radiance to dn.

Name: "A2"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 3  
Dimensions: dim1: NUM\_a2\_vs\_T\_inst\_COEFF  
dim2: NUM\_MIRROR\_SIDES  
dim3: NUM\_EMISSIVE\_DETECTORS  
Range: [-1., 1.]  
Fill Value: (none)

---

---

#### 4.2.13 *Ucoeff*

This LUT contains polynomial coefficients of a fit of uncertainty in various parameters to radiance level.

Name: "Ucoeff"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 4  
Dimensions: dim1: NUM\_EMISSIVE\_DETECTORS  
dim2: NUM\_UI\_PARAMETERS  
dim3: NUM\_UI\_POLYNOMIAL\_COEFF  
dim4: NUM\_FI\_POLYNOMIAL\_COEFF  
Range: N/A  
Fill Value: -999

---

#### 4.2.14 *Sigma\_TEB\_PV\_resid\_elec*

This LUT contains fractional uncertainties for the PV bands residual electronic cross-talk.

Name: "Sigma\_TEB\_PV\_resid\_elec"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: N/A  
Fill Value: -999

---

---

#### 4.2.15 *Sigma\_TEB\_ADC*

This LUT contains the uncertainty in dn ascribed to the detector ADCs.

Name: "Sigma\_TEB\_ADC"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: N/A  
Fill Value: -999

---

#### 4.2.16 *Ucoeff\_Calibr\_resid*

This LUT contains the residual uncertainty to the calibration fit expressed in terms of 4<sup>th</sup> order polynomial coefficients of the retrieved radiance.

Name: "Ucoeff\_Calibr\_resid"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 2  
Dimensions: dim1: NUM\_EMISSIVE\_DETECTORS  
dim2: NUM\_4TH\_ORDER\_COEFFS  
Range: N/A  
Fill Value: -999

---

#### 4.2.17 *BB\_DN\_first\_frame\_to\_use*

This LUT holds the index (0 through N-1) of the first frame to use in computing the average blackbody (BB) DN.

Name: "BB\_DN\_first\_frame\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 49]  
Fill Value: (none)

---

---

#### 4.2.18 *BB\_DN\_number\_of\_frames\_to\_use*

This LUT holds the number of frames to use in computing the average blackbody (BB) DN.

Name: "BB\_DN\_number\_of\_frames\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [1, 50]  
Fill Value: (none)

---

#### 4.2.19 *SV\_DN\_first\_frame\_to\_use*

This LUT holds the index (0 through N-1) of the first frame to use in computing the average space-view (SV) DN.

Name: "SV\_DN\_first\_frame\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 49]  
Fill Value: (none)

---

#### 4.2.20 *SV\_DN\_number\_of\_frames\_to\_use*

This LUT holds the number of frames to use in computing the average space-view (SV) DN.

Name: "SV\_DN\_number\_of\_frames\_to\_use"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [1, 50]  
Fill Value: (none)

---

---

#### 4.2.21 *SV\_DN\_moon\_include\_frames*

This LUT holds the number of frames to use when calculating average space-view (SV) DN, when the moon is in the SV keep-out-box. If the moon is determined to be in the SV KOB, the 50 SV DN frames are sorted from the lowest to the highest values. The lowest "SV\_DN\_moon\_include\_frames" are used to calculate average SV DN using the same logic as is done for the case when the moon is not in the SV KOB.

Name: "SV\_DN\_moon\_include\_frames"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 50]  
Fill Value: (none)

---

#### 4.2.22 *num\_overlap\_scans\_b1*

This LUT defines the number of scans in leading and trailing granules to use for cross-granule averaging of b1. In the leading granule, these scans come from the end of the granule. In the trailing granule these scans from the beginning of the granule.

Name: "num\_overlap\_scans\_b1"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 1  
Range: [0, 100]  
Fill Value: (none)  
Values: Presently 40.

---

#### 4.2.23 *T\_ins\_function\_flag*

There are 4 thermistors, which may be used to determine the instrument temperature. This LUT holds the off-line determination of whether any of these may be used to perform this function. A value of 1 (one) means that the thermistor may be used to determine instrument temperature. A value of zero means that the thermistor should not be used.

Name: "T\_ins\_function\_flag"  
Kind: SDS Constant LUT  
Data type: int32  
Rank: 1  
Dimension: NUM\_T\_INS\_THERMISTORS  
Range: [0, 1]  
Fill Value: (none)

<b>index</b>	<b>telemetry point</b>
0	TP_AO_SMIR_OBJ
1	TP_AO_LWIR_OBJ
2	TP_AO_SMIR_LENS
3	TP_AO_LWIR_LENS

#### 4.2.24 *T\_ins\_default*

This LUT holds the default value of instrument temperature in °K to use if, on any scan, none of the four thermistors (described above in the *T\_ins\_function\_flag* LUT, Section 4.2.23) was found suitable for providing the on line instrument temperature.

Name: "T\_ins\_default"  
Kind: SDS Constant LUT  
Data type: Float32  
Rank: 1  
Dimension: 1  
Range: [200., 300.]  
Fill Value: (none)

#### 4.2.25 *T\_ins\_offset*

This LUT holds the thermistor-dependent offset value of instrument temperature in °K to in the equation for computing the instrument temperature. These offsets are determined off-line.

Name: "T\_ins\_offset"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_T\_INS\_THERMISTORS  
Range: [-15., 15.]  
Fill Value: (none)

index	telemetry point
0	TP_AO_SMIR_OBJ
1	TP_AO_LWIR_OBJ
2	TP_AO_SMIR_LENS
3	TP_AO_LWIR_LENS

#### 4.2.26 *T\_cav\_function\_flag*

There are 4 thermistors that may be used for determining the instrument scan cavity temperature, *T\_cav*. This LUT holds the off-line determination of whether any of these may be used to perform this function. A value of 1 (one) means that the thermistor may be used to determine instrument temperature. A value of zero means that the thermistor should not be used.

Name: "T\_cav\_function\_flag"  
Kind: SDS Constant LUT  
Data type: int32  
Rank: 1  
Dimension: NUM\_T\_CAV\_THERMISTORS  
Range: [0, 1]  
Fill Value: (none)

Index	telemetry point
0	TP_MF_CALBKHD_SR
1	TP_SR_SNOUT
2	TP_MF_Z_BKHD_BB
3	TP_MF_CVR_OP_SR

#### 4.2.27 *T\_cav\_default*

This LUT holds the default value of instrument scan cavity temperature in °K to use if, on any scan, none of the four thermistors (described above in the T\_cav\_function\_flag LUT) was found suitable for providing the on-line scan cavity temperature.

Name: "T\_cav\_default"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [230, 300]  
Fill Value: (none)

#### 4.2.28 *T\_mir\_function\_flag*

There are 2 thermistors that may be used for determining the average mirror temperature, T\_mir. This LUT holds the off-line determination of whether any of these may be used to perform this function. A value of 1 (one) means that the thermistor may be used to determine temperature. A value of zero means that the thermistor should not be used.

Name: "T\_mir\_function\_flag"  
Kind: SDS Constant LUT  
Data type: int32  
Rank: 1  
Dimension: NUM\_T\_MIR\_THERMISTORS  
Range: [0, 1]  
Fill Value: (none)

Index	telemetry point
0	TP_SA_RCT1_MIRE
1	TP_SA_RCT2_MIRE

---

#### 4.2.29 *T\_mir\_default*

This LUT holds the default value of average mirror temperature in °K to use if, on any scan, neither of the two mirror thermistors (described above) was found suitable for providing the on-line scan average mirror temperature.

Name: "T\_mir\_default"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [230, 300]  
Fill Value: (none)

---

#### 4.2.30 *BB\_Weight*

There are 12 blackbody (BB) thermistors that may be used for determining the average blackbody temperature. This LUT holds a weighting factor to be applied in computing the average. This array should only be populated with 1's and 0's.

Name: "BB\_Weight"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_BB\_THERMISTORS  
Range: [0, 1]  
Fill Value: (none)

---

#### 4.2.31 RVS\_TEB

This LUT holds the 2<sup>nd</sup> order polynomial coefficients used to compute the response vs. scan angle (RVS) as a function of frame number for the emissive bands. Polynomial evaluation will result in a normalized value of 1.0 at the center frame of the Solar diffuser sector (approximately 50.25 degrees).

Name: "RVS\_TEB"  
Kind: SDS Piecewise Linear LUT  
Data type: float32  
Intrinsic Rank: 4  
Dimensions: dim1: NUM\_EMISSIVE\_BANDS  
dim2: DETECTORS\_PER\_1KM\_BAND  
dim3: NUM\_MIRROR\_SIDES  
dim4: NUM\_2ND\_ORDER\_COEFFS  
Range: [-2.0E-4, 1.2]  
Fill Value: -999

#### 4.2.32 RVS\_BB\_SV\_Frame\_No

This LUT give the reference frame numbers to use when calculating the blackbody and space view RVS corrections respectively. The coefficients to use in the second order polynomial are given by the RVS\_TEB LUT (see Section 4.2.31).

Name: "RVS\_BB\_SV\_Frame\_No"  
Kind: SDS Constant LUT  
Data type: int16  
Rank: 1  
Dimension: 2  
Range: [0, 1353]  
Fill Value: (none)

---

#### 4.2.33 *Band\_21\_b1*

This lookup table holds the value of b1 to be used for each detector of MODIS Band 21. While other bands have values of b1 computed on-line, these values for Band 21 are calculated off-line.

Name: "Band\_21\_b1"  
Kind: SDS Step Function LUT  
Data type: float32  
Intrinsic Rank: 1  
Dimension: DETECTORS\_PER\_1KM\_BAND  
Range: [0.005, 0.1]  
Fill Value: (none)

---

#### 4.2.34 *Band\_21\_Uncert\_Lsat*

This lookup table holds part of an extra uncertainty term related only to Band 21. Units of this value are 1/radiance.

Name: "Band\_21\_Uncert\_Lsat"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0, 0.2]  
Fill Value: (none)

---

---

#### 4.2.35 *L\_Max*

In the Level 1B code, a calibrated radiance in the range of [L\_Min, L\_Max] is scaled to an integer in the range of [0, 32767]. This LUT contains L\_Max, the upper limit of the dynamic range, for each of the thermal emissive bands. Units are watts/m<sup>2</sup>/μm/steradians.

Name: "L\_Max"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_BANDS  
Range: [0, 110]  
(MODIS/AQUA)  
Fill Value: (none)  
[0, 100]  
(MODIS/TERRA)

---

#### 4.2.36 *L\_Min*

In the Level 1B code, a calibrated radiance in the range of [L\_Min, L\_Max] is scaled to an integer in the range of [0, 32767]. This LUT contains L\_Min, the lower limit of the dynamic range, for each of the thermal emissive bands. Units are watts/m<sup>2</sup>/μm/steradians.

Name: "L\_Min"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_BANDS  
Range: [-10., 0]  
Fill Value: (none)

---

---

#### 4.2.37 *TEB\_specified\_uncertainty*

This LUT contains the specified uncertainty factor for each thermal emissive band which is used (along with the scaling factor described by the LUT "TEB\_scaling\_factor") to convert percent uncertainty to an uncertainty index. The formula for uncertainty index (UI) in terms of uncertainty\_in\_percent, specified\_uncertainty and scaling\_factor is:

$$UI = \text{scaling\_factor} * \ln (\text{uncertainty\_in\_percent} / \text{specified\_uncertainty})$$

where "ln" is the natural logarithm. The uncertainty in percent is computed dynamically within the Level 1B code.

Name: "TEB\_specified\_uncertainty"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_BANDS  
Range: N/A  
Fill Value: (none)

---

#### 4.2.38 *TEB\_UI\_scaling\_factor*

This LUT contains the scaling factor for each thermal emissive band for computing the uncertainty index from uncertainty in percent. (see comments above for "TEB\_specified\_uncertainty", Section 4.2.37).

Name: "TEB\_UI\_scaling\_factor"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_BANDS  
Range: N/A  
Fill Value: (none)

---

#### 4.2.39 PC\_XT

This LUT holds coefficients used in the PC bands electronic cross-talk correction algorithm. The algorithm is used to correct dn\_bb (effective digital number for the black body) and dn\_ev (effective Earth-view digital number).

Name: "PC\_XT"  
 Kind: SDS Constant LUT  
 Data type: float32  
 Rank: 3  
 Dimensions: dim1: NUM\_PC\_XT\_BANDS  
 dim2: DETECTORS\_PER\_1KM\_BAND  
 dim3: NUM\_PC\_XT\_PARAMETERS  
 Range: [-15, 15]  
 Fill Value: (none)

For dimension 3, the meanings of the elements are:

XT[:, :][0]  
 Percentage amount of cross-talk from Band 31 to Band i (quantity "Xtalk<sub>Band(31)-Band(i)</sub>")

XT[:, :][1]  
 EV frame offset of Band 31 to Band i (not used for dn\_bb) (quantity "FO<sub>31-i</sub>").

XT[:, :][2]  
 quantity "PCX\_correction\_switch"

This LUT defines whether the PC bands electronic cross-talk correction to DN is applied. The PCX correction enables the "PC\_XT" LUT described earlier (see Section 4.2.39).

Name: "PCX\_correction\_switch"  
 Kind: SDS Constant LUT  
 Data type: int8  
 Rank: 1  
 Dimension: 1  
 Range: [0, 1]  
 Fill Value: (none)

---

4.2.41 *BB\_T\_sat\_switch\_aqua (MODIS/AQUA only)*

For use with MODIS/Aqua data only, this LUT defines whether the blackbody warmup temperature saturation correction is applied. If this correction switch is set to 0 (OFF), then the correction will not be applied regardless of the values in the LUTs relating to this correction (see *BB\_T\_sat\_aqua*, Section 4.2.42, and *BB\_T\_sat\_default\_b1\_aqua*, Section 4.2.43).

Name: "BB\_T\_sat\_switch\_aqua"  
Kind: SDS Constant LUT  
Data type: int8  
Rank: 1  
Dimension: 1  
Range: [0, 1]  
Fill Value: (none)

---

4.2.42 *BB\_T\_sat\_aqua (MODIS/AQUA only)*

For use with MODIS/Aqua data only. For each emissive band affected, this LUT gives a blackbody temperature in degrees Kelvin at which its detectors are considered saturated.

Name: "BB\_T\_sat\_aqua"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_AQUA\_BB\_SAT\_BANDS  
Range: [270.0, 315.0]  
Fill Value: (none)

---

---

4.2.43 *BB\_T\_sat\_default\_b1\_aqua (MODIS/AQUA only)*

For use with MODIS/Aqua data only. For each emissive band/detector combination affected, this LUT gives a default b1 value to use when the blackbody temperature is over the saturation threshold defined by "BB\_T\_sat\_aqua" (see Section 4.2.42).

Name: "BB\_T\_sat\_default\_b1\_aqua"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 3  
Dimensions: dim1: NUM\_AQUA\_BB\_SAT\_BANDS  
dim2: DETECTORS\_PER\_1KM\_BAND  
dim3: NUM\_MIRROR\_SIDES  
Range: [0.001, 0.01]  
Fill Value: (none)

---

## 5 QUALITY ASSURANCE (QA) LUTS

### 5.1 Summary of Quality Assurance LUTs and their Dimensions

Table 5.1 Summary of QA Lookup Tables

LUT name	Section	Meaning
QA serial number	5.2.1	Version of the science content of the QA LUTs
PGE_Version	5.2.2	3-digit PGE version number
MCST_Version	5.2.3	4-digit ALGORITHMPACKAGEVERSION
ASSOCIATEDPLATFORM SHORTNAME	5.2.4	Identifies the platform (e.g. "Terra").
ALGORITHMPACKAGE ACCEPTANCEDATE	5.2.5	Identifies the algorithm package date, written to ECS archive metadata.
ALGORITHMPACKAGE MATURITYCODE	5.2.6	Identifies the algorithm package maturity code, written to ECS archive metadata.
mission phase	5.2.7	Identifies the mission phase.
Control options	5.2.8	Miscellaneous code switches
Detector Quality Flag Values	5.2.9	Array of integers identifying noisy, dead and anomalous detectors.
a1	5.2.10	Pre-launch values of the average MODIS linear response term for each emissive detector.
NedL	5.2.11	Pre-launch noise equivalent difference in radiance for each emissive detector.
T_BB_Variance	5.2.12	Pre-launch variance of each of the 12 BB temperatures.
BB Average Temperature Variance	5.2.13	Pre-launch variance of the average BB temperature.
visual FPA base variance	5.2.14	pre-launch value of the variance of the VIS FPA temperature
NIR FPA base variance	5.2.15	pre-launch value of the variance of the NIR FPA temperature
LWIR FPA Temperature Variance	5.2.16	Pre-launch variance of the LWIR FPA temperature.
MWIR FPA Temperature Variance	5.2.17	Pre-launch variance of the MWIR FPA temperature.
MirrorSide 1 Temperature Variance	5.2.18	Pre-launch variance of the mirror-side 1 temperature.
MirrorSide 2 Temperature Variance	5.2.19	Pre-launch variance of the mirror-side 2 temperature.
Mirror Average Temperature Variance	5.2.20	Pre-launch variance of the average mirror-side temperature.
Instrument Temperature Variance	5.2.21	Pre-launch variance of the instrument temperature.
Cavity Temperature Variance	5.2.22	Pre-launch variance of the cavity temperature.
Moon Offset Limits	5.2.23	Defines the limits of the "Keep-out" box relative to center of the SVP. (This is not strictly a QA LUT because it is used in processing)
Spacecraft Roll Threshold Angle	5.2.24	Defines the upper limit of the absolute value in degrees by which the spacecraft roll angle may deviate from nominal before the spacecraft maneuver flag is set.
Spacecraft Pitch Threshold Angle	5.2.25	Defines the upper limit of the absolute value in degrees by which the spacecraft pitch angle may deviate from nominal before the spacecraft maneuver flag is set.
Spacecraft Yaw Threshold Angle	5.2.26	Defines the upper limit of the absolute value in degrees by which the spacecraft yaw angle may deviate from nominal before the spacecraft maneuver flag is set.

**Table 5.2 Dimensions used in QA LUTs**

<b>Dimension macro</b>	<b>value</b>	<b>meaning</b>
<b>MAX_ALGORITHMPACKAGEACCEPTANCEDATE_BUFFER</b>	<b>11</b>	<b>Largest allowable string length</b>
<b>MAX_ALGORITHMPACKAGEMATURITYCODE_BUFFER</b>	<b>11</b>	<b>Largest allowable string length</b>
<b>MAX_ASSOCIATEDPLATFORMSHORTNAME_BUFFER</b>	<b>21</b>	<b>Largest allowable string length</b>
<b>MAX_MCST_VERSION_BUFFER</b>	<b>21</b>	<b>Largest allowable string length</b>
<b>MAX_MISSION_PHASE_BUFFER</b>	<b>11</b>	<b>Largest allowable string length</b>
<b>MAX_PGE_VERSION_BUFFER</b>	<b>11</b>	<b>Largest allowable string length</b>
<b>MAX_SERIAL_NUMBER_BUFFER</b>	<b>31</b>	<b>Largest allowable string length</b>
<b>NUM_BANDS</b>	<b>38</b>	<b>Number of MODIS Bands</b>
<b>NUM_BB_THERMISTORS</b>	<b>12</b>	<b>Number of blackbody thermistors</b>
<b>NUM_BITS_IN_UINT8</b>	<b>8</b>	<b>Number of bits in unsigned integer</b>
<b>NUM_CONTROL_OPTIONS</b>	<b>2</b>	<b>Number of control options</b>
<b>NUM_DETECTORS</b>	<b>490</b>	<b>Number of MODIS detectors</b>
<b>NUM_EMISSIVE_DETECTORS</b>	<b>160</b>	<b>Number of MODIS Emissive band detectors</b>
<b>NUM_MOON_OFFSET_LIMITS</b>	<b>4</b>	<b>Number of moon offset limits</b>

## 5.2 QA LUT Listing

---

### 5.2.1 QA serial number

This serial number serves to identify the "science content" version of the QA lookup tables. It is stored as a string and has the form: "Qvvv yyyy:MM:dd:hh:mm", where "Q" is for quality assurance, "vvv" is an integer version number that gets incremented each time values of any LUT other than PGE\_Version and MCST\_Version are changed, and "yyyy:MM:dd:hh:mm" is the date and time of the last change to any LUT other than PGE\_Version and MCST\_Version. This serial number is updated as lookup tables are modified during operations. It is placed in each of the L1B Earth-view products.

Name: "QA serial number"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_SERIAL\_NUMBER\_BUFFER  
Range: N/A  
Fill Value: (none)

---

### 5.2.2 PGE Version

This LUT contains the value of PGEVERSION, written to the ECS core metadata in each of the L1B products (see amplified description in Section 3.3.2).

Name: "PGE Version LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_PGE\_VERSION\_BUFFER  
Range: N/A  
Fill Value: (none)

---

### 5.2.3 MCST Version

This LUT contains the MCST version of Level 1B algorithms, code and LUTs (see amplified description in Section 3.3.3).

Name: "MCST Version LUT"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_MCST\_VERSION\_BUFFER  
Range: N/A  
Fill Value: (none)

---

---

#### 5.2.4 ASSOCIATEDPLATFORMSHORTNAME

This LUT contains the value of ASSOCIATEDPLATFORMSHORTNAME written to the ECS archive metadata in each of the L1B products. This value identifies the satellite platform.

Name: "ASSOCIATEDPLATFORMSHORTNAME"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_ASSOCIATEDPLATFORMSHORTNAME\_BUFFER  
Range: N/A  
Fill Value: (none)  
Example: "Aqua"

---

#### 5.2.5 ALGORITHMPACKAGEACCEPTANCEDATE

This LUT contains the value of ALGORITHMPACKAGEACCEPTANCEDATE, written to the ECS archive metadata in each of the L1B products. This value should be the date that the algorithm passed AI&T procedures and were accepted as ECS standard algorithm. In practice, we will set this value to the date that a major code delivery is made. Upon notification from the DAAC that the code has been accepted, we will update the date to that value. The format of the date, YYYY-MM-DD, is indicated in the example below.

Name: "ALGORITHMPACKAGEACCEPTANCEDATE"  
Kind: Global Attribute  
Data type: String  
Rank: 1  
Dimension: MAX\_ALGORITHMPACKAGEACCEPTANCEDATE\_BUFFER  
Range: N/A  
Fill Value: (none)  
Example: "1999-10-31"

---

---

### 5.2.6 ALGORITHMPACKAGEMATURITYCODE

This LUT contains the value of ALGORITHMPACKAGEMATURITYCODE, written to the ECS archive metadata in each of the L1B products. This value specifies the maturity of the algorithm as a whole. Values may come from the following set: "pre-launch", "PREL" (preliminary), "OPL" (operational), "stable", "final".

Name: "ALGORITHMPACKAGEMATURITYCODE"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_ALGORITHMPACKAGEMATURITYCODE\_BUFFER  
Range: N/A  
Fill Value: (none)  
Example: "pre-launch"

---

### 5.2.7 mission phase

This LUT defines the current mission phase.

Name: "mission phase"  
Kind: Global Attribute  
Data type: string  
Rank: 1  
Dimension: MAX\_MISSION\_PHASE\_BUFFER  
Range: N/A  
Fill Value: (none)  
Value: Set to "A&E" at launch, updated as necessary thereafter.

---

### 5.2.8 *Control options*

This LUT contains miscellaneous L1B control options. Currently, there are two:

- Split scans — control whether or not to treat as missing if a split scan is detected. There are only two values, ON or OFF.
- Bad scan quality — control whether or not to treat as missing if an invalid value of Scan quality array is detected. There are only two values, ON or OFF.

Name:                "Control options"  
Kind:                SDS                    Constant LUT  
Data type:          uint8  
Rank:                1  
Dimension:          NUM\_CONTROL\_OPTIONS  
Range:               [0, 1]  
Fill Value:          (none)

### 5.2.9 Detector Quality Flag Values

This LUT contains the array used to fill the L1B global attribute "Detector Quality Flag". Each array element of this LUT (having a value of 0 or 1) sets one bit of the corresponding L1B attribute. The second dimension of this LUT cycles through the 8 bits of each word of the attribute, with array element [w][0] corresponding to the least significant bit on word w and array element [w][7] corresponding to the most significant bit.

Name: "Detector Quality Flag Values"  
 Kind: SDS Step Function LUT  
 Data type: uint8  
 Intrinsic Rank: 2  
 Dimensions: dim1: NUM\_DETECTORS  
 dim2: NUM\_BITS\_IN\_UINT8  
 Range: [0, 1]  
 Fill Value: (none)

The following table maps the meaning of the bits in the Detector Quality Flag to the array elements:

Detector Quality Flag bit (0-based)	LUT 2nd dimension array element index (0-based)	Meaning
bit 0 (LSB)	0	Noisy Detector
bit 1	1	Dead Detector
bit 2	2	Out-of-Family Gain
bit 3	3	Dynamic Range
bit 4	4	RSB: Detector DN saturates on illuminated, unscreened Solar Diffuser TEB: Detector DN saturates on blackbody when $T_{BB} = 300$ K
bit 5	5	High calibration fit residuals
bit 6	6	Electrical Crosstalk
bit 7 (MSB)	7	(TBD)

---

### 5.2.10 *a1*

This LUT holds pre-launch values of the average MODIS linear response term,  $b_{1BB}$ , for each of the emissive detectors. It is used in a ratio with current values of the average (over scans) of  $b_{1BB}$  to track changes during the mission.

Name: "a1"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [0.0002, 0.1]  
Fill Value: (none)

---

### 5.2.11 *NEdL*

This LUT contains pre-launch values of the Noise Equivalent Difference in Radiance for all emissive detectors.

Name: "NEdL"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_EMISSIVE\_DETECTORS  
Range: [0.0002, 0.1]  
Fill Value: (none)

---

---

### 5.2.12 *T\_BB\_Variance*

This LUT holds pre-launch values of the variance of each of the 12 blackbody (BB) temperatures. The code forms the ratio of the current variances (across scans within the granule) with these values to track the change during the mission.

Name: "T\_BB\_Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: NUM\_BB\_THERMISTORS  
Range: [0.0002, 0.1]  
(MODIS/AQUA)  
Fill Value: (none)  
[0.001, 0.1]  
(MODIS/TERRA)

---

### 5.2.13 *BB Average Temperature Variance*

This LUT holds a pre-launch value of the variance of the average blackbody (BB) temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "BB Average Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.00003, 0.1]  
(MODIS/AQUA)  
Fill Value: (none)  
[0.0001, 0.1]  
(MODIS/TERRA)

---

---

*5.2.14 visual FPA base variance*

This LUT holds a pre-launch value of the variance of the VIS FPA temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "visual FPA base variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: N/A  
Fill Value: (none)

---

*5.2.15 NIR FPA base variance*

This LUT holds a pre-launch value of the variance of the NIR FPA temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "NIR FPA base variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: N/A  
Fill Value: (none)

---

---

### 5.2.16 LWIR FPA Temperature Variance

This LUT holds a pre-launch value of the variance of the LWIR FPA temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "LWIR FPA Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.000001, 0.1]  
Fill Value: (none)

---

### 5.2.17 MWIR FPA Temperature Variance

This LUT holds a pre-launch value of the variance of the MWIR FPA temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "MWIR FPA Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0005, 0.1]  
(MODIS/AQUA)  
Fill Value: (none)  
[0.001, 0.1]  
(MODIS/TERRA)

---

---

### 5.2.18 *MirrorSide 1 Temperature Variance*

This LUT holds a pre-launch value of the variance of the mirror (side 1) temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "MirrorSide 1 Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0001, 0.1]  
Fill Value: (none)

---

### 5.2.19 *MirrorSide 2 Temperature Variance*

This LUT holds a pre-launch value of the variance of the mirror (side 2) temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "MirrorSide 2 Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0001, 0.1]  
Fill Value: (none)

---

---

### 5.2.20 *Mirror Average Temperature Variance*

This LUT holds a pre-launch value of the variance of the average mirror temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "Mirror Average Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.001, 0.1]  
Fill Value: (none)

---

### 5.2.21 *Instrument Temperature Variance*

This LUT holds a pre-launch value of the variance of the instrument temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "Instrument Temperature Variance"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.001, 0.1]  
Fill Value: (none)

---

### 5.2.22 Cavity Temperature Variance

This LUT holds a pre-launch value of the variance of the cavity temperature. The code forms the ratio of the current variance (across scans within the granule) with this value to track the change during the mission.

Name: "Cavity Temperature Variance"  
 Kind: SDS Constant LUT  
 Data type: float32  
 Rank: 1  
 Dimension: 1  
 Range: [0.01, 0.1]  
 Fill Value: (none)

### 5.2.23 Moon Offset Limits

This LUT holds scan and track offsets (in terms of 1km pixel increments) from the center of the space view (SV) port. These are used to define a "box" in the scan-track coordinate system about the center of the SV port. If the vector to the moon, after converting to these coordinates, lies within the box, then the moon is considered to be within the SV port.

Name: "Moon Offset Limits"  
 Kind: SDS Constant LUT  
 Data type: float32  
 Rank: 2  
 Dimensions: dim1: NUM\_BANDS  
 dim2: NUM\_MOON\_OFFSET\_LIMITS  
 Range: [-200, 200]  
 Fill Value: (none)

The second dimension of this LUT is used to extract the four different offsets:

index	macro	meaning
0	TRK_UPPER	Upper limit offset in the along track direction.
1	TRK_LOWER	Lower limit offset in the along track direction.
2	SCN_UPPER	Upper limit offset in the along scan direction.
3	SCN_LOWER	Lower limit offset in the along scan direction.

For both MODIS/Terra and MODIS/Aqua, values for TRK\_UPPER and TRK\_LOWER for all bands are presently +15 and -15 and values for SCN\_UPPER and SCN\_LOWER for all bands are +55 and -55.

---

#### 5.2.24 *Spacecraft Roll Threshold Angle*

This LUT defines the upper limit of the absolute deviation from nominal allowed in the spacecraft roll angle, measured in degrees. If the spacecraft deviates from nominal by more than this amount, the maneuver flag is set.

Name: "Spacecraft\_Roll\_Threshold\_Angle"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0, 360.0]  
Fill Value: (none)  
Value: Currently set to 1.0

---

#### 5.2.25 *Spacecraft Pitch Threshold Angle*

This LUT defines the upper limit of the absolute deviation from nominal allowed in the spacecraft pitch angle, measured in degrees. If the spacecraft deviates from nominal by more than this amount, the maneuver flag is set.

Name: "Spacecraft\_Pitch\_Threshold\_Angle"  
Kind: SDS Constant LUT  
Data type: float32  
Rank: 1  
Dimension: 1  
Range: [0.0, 360.0]  
Fill Value: (none)  
Value: Currently set to 1.0

---

### 5.2.26 *Spacecraft Yaw Threshold Angle*

This LUT defines the upper limit of the absolute deviation from nominal allowed in the spacecraft yaw angle, measured in degrees. If the spacecraft deviates from nominal by more than this amount, the maneuver flag is set.

Name:                "Spacecraft\_Yaw\_Threshold\_Angle"  
Kind:                SDS                    Constant LUT  
Data type:          float32  
Rank:                1  
Dimension:          1  
Range:               [0.0, 360.0]  
Fill Value:          (none)  
Value:                Currently set to 1.0

---