

# S-NPP/N20 VIIRS Thermal Emissive Bands On-Orbit Performance and Calibration

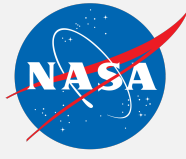
TEB group

*VIIRS Characterization Support Team, NASA/GSFC  
Acknowledgments: VIIRS SDR team and VCST members*

*MODIS and VIIRS Science Team Meeting, College Park, MD  
November 18<sup>th</sup>-22<sup>nd</sup>, 2019*



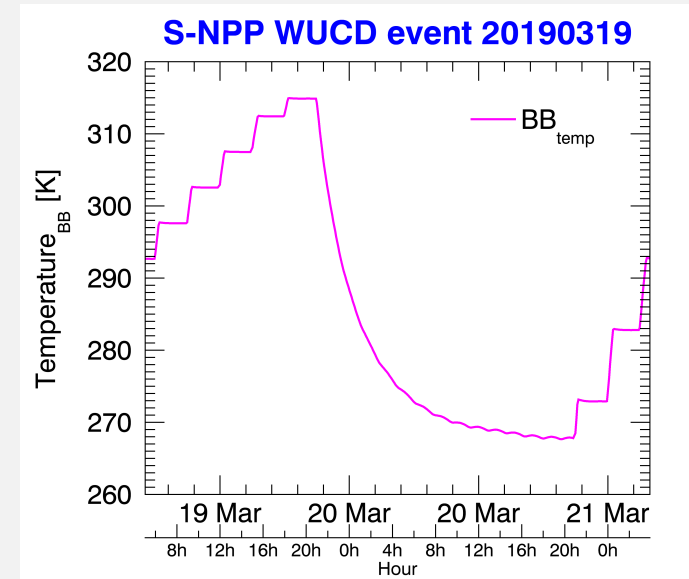
# Outline



- **Instrument overview**
  - OBC black body
  - TEBs and their calibration algorithm
- **On-orbit performance**
  - BB short- and long-term stabilities
  - F-factor short- and long-term responses
  - Noise characterization (NEdT)
  - WUCD calibration coefficients trending
- **Improvements**
- **Summary**

# OBC black body

- TEBs are calibrated using on-board calibrator (OBC) black body (BB)
- Nominal temperature maintained at 292.5 K
- Warm-up and cooldown (WUCD) events are performed to fully characterize TEBs detector response and derive the offset and non-linear terms in the calibration algorithm
- WUCD temperature range: ambient to 315 K
- Event frequency:
  - Launch → 06/2018 (quarterly)
  - 06/2018 → present (annually)
  - # of WUCD: 27 (S-NPP) and 5 (N20)

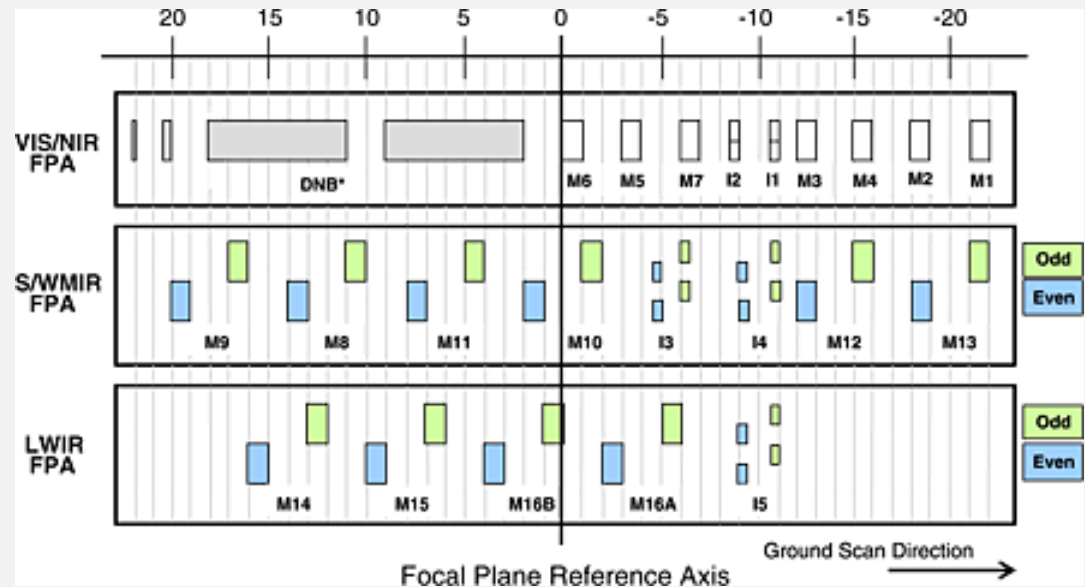


- 6 thermistors

# Thermal Emissive Bands (TEBs)

TEB	375 m		750 m				
	I4	I5	M12	M13 (D.G.)	M14	M15	M16
$\lambda$ [ $\mu\text{m}$ ]	3.74	11.45	3.70	4.05	8.55	10.76	12.01

- TEBs radiance source is BB reference
- BB has known emissivity and temperature
- Self-emission from inside the instrument must be subtracted
- Emissivities and temperatures from several instrument components are used
- Scaling factor (F-factor) is derived on a scan-by-scan basis





# TEBs calibration algorithm

VIIRS Earth view (EV) radiance is retrieved following the ATBD,

$$L_{EV}(B, \theta) = \frac{F(B) \sum_{i=0}^2 c_i(B) dn^i(B) - \Delta L_{bg}(B, \theta)}{RVS(B, \theta)},$$

B: band  
 $\theta$ : angle-of-incidence  
 dn: detector response  
 $c_i$ : calibration coeffs.

where the  $\Delta L_{bg}(B, \theta)$  is the background difference between the EV and space view (SV) path:

$$\Delta L_{bg}(B, \theta) = (RVS(B, \theta) - RVS_{SV}(B)) \left[ \frac{(1 - \rho_{RTA}(B))}{\rho_{RTA}(B)} L_{RTA} - \frac{1}{\rho_{RTA}(B)} L_{HAM} \right].$$

The F-factor is derived on a scan-by-scan basis and is band-, detector-, and HAM side-dependent:

$$F(B) = \frac{RVS_{BB}(B) L_{ap}(B) + \Delta L_{bg}(B, \theta_{BB})}{\sum_{i=0}^2 c_i dn_{BB}^i}.$$

→ **Estimated**  
 → **Retrieved**

$\rho$ : reflectance

$L_{ap}$ : aperture radiance

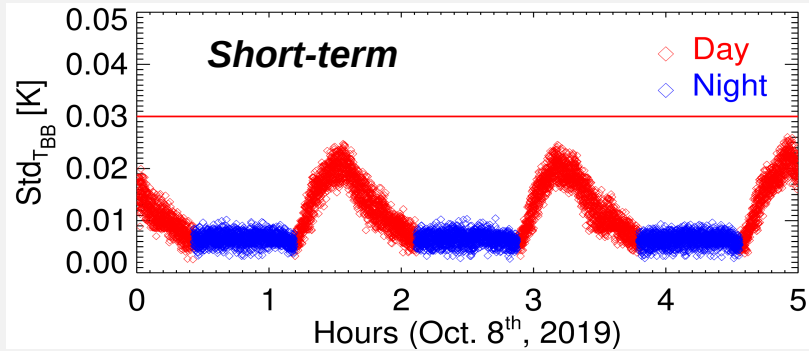
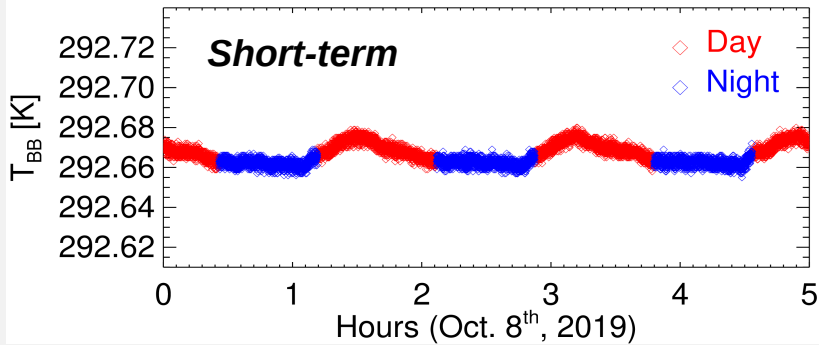
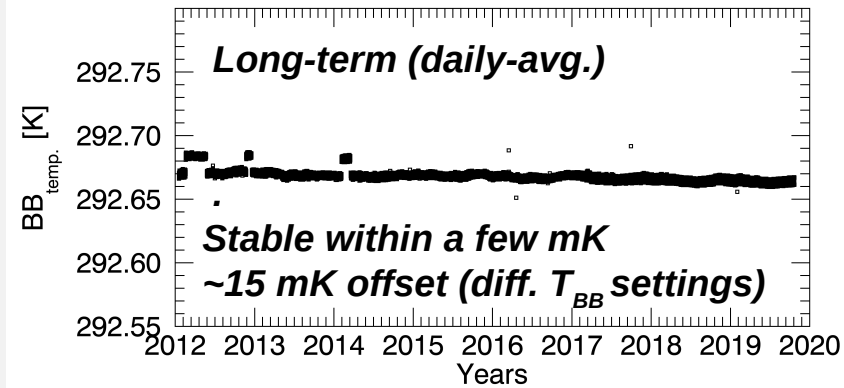
RTA: rotating telescope assembly

HAM: half-angle mirror

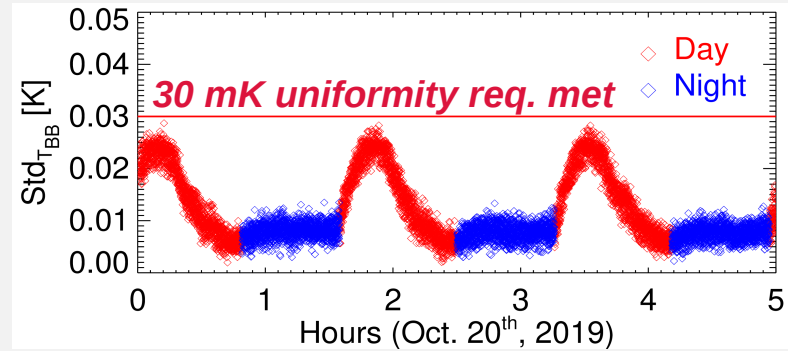
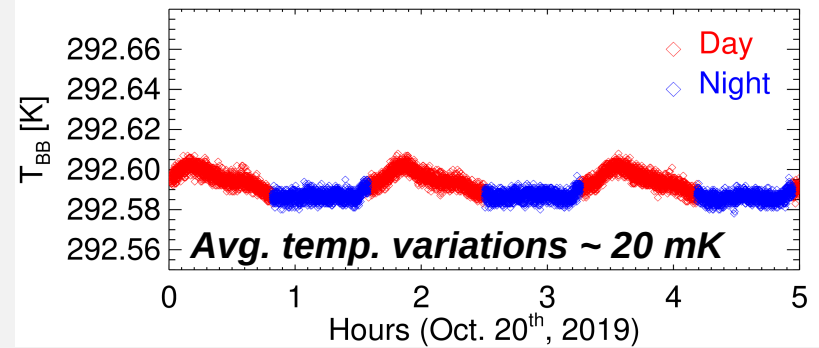
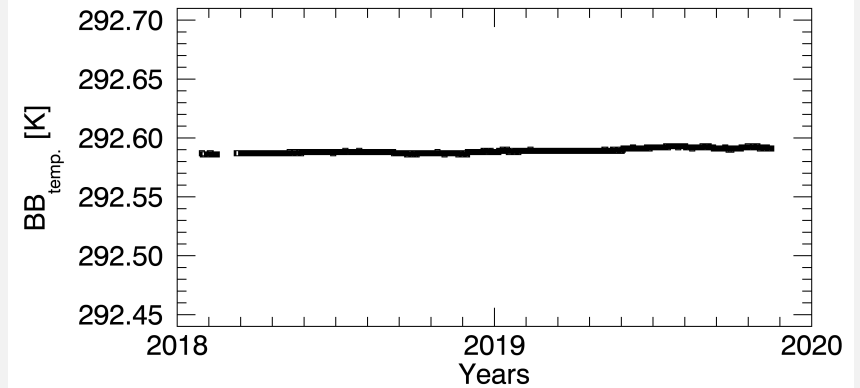
RVS: response-versus-scan-angle

# BB performance

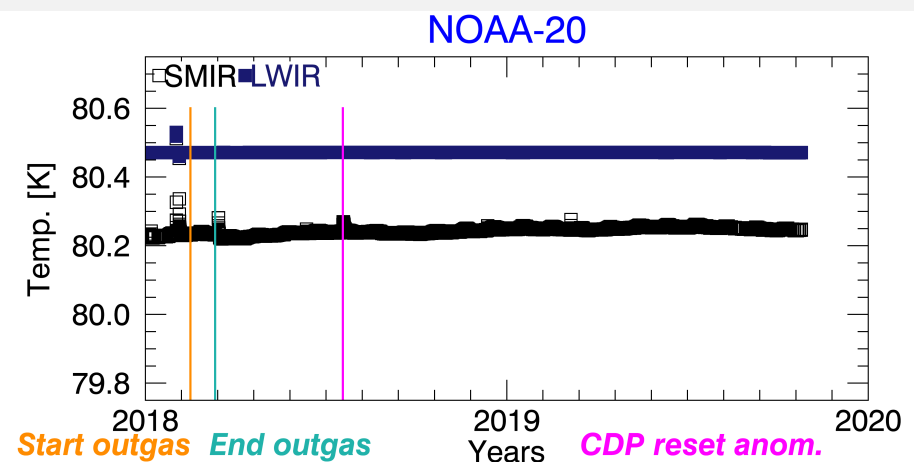
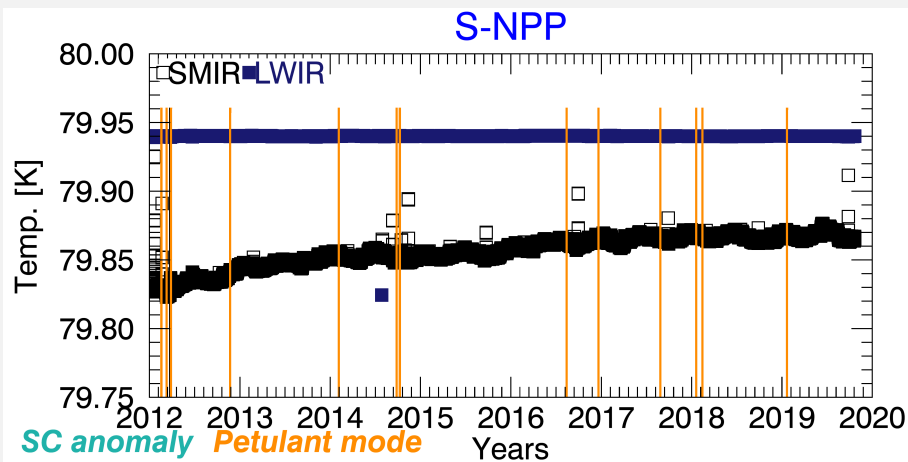
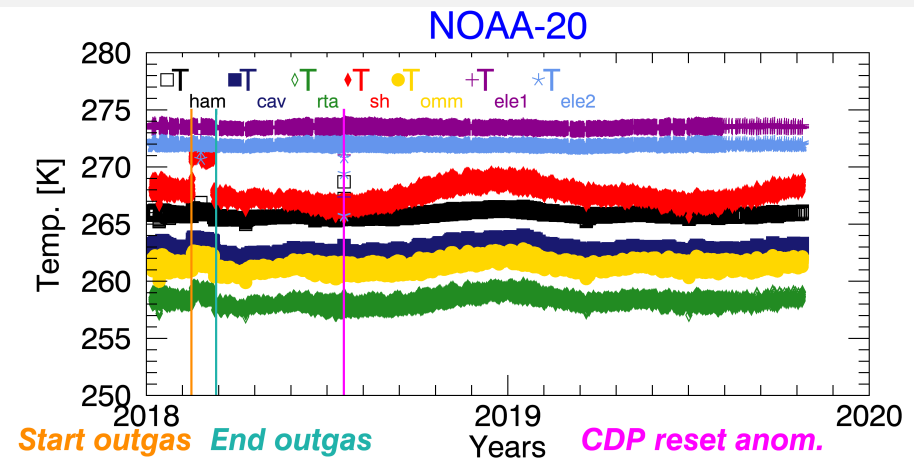
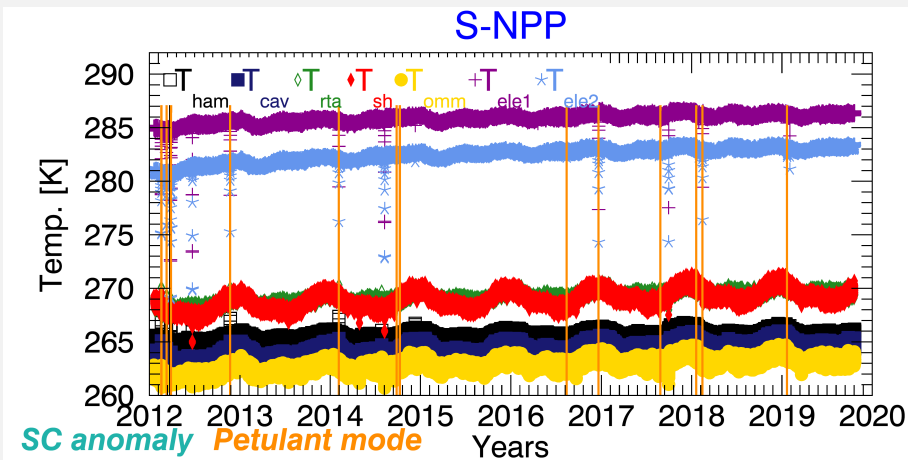
S-NPP



NOAA-20

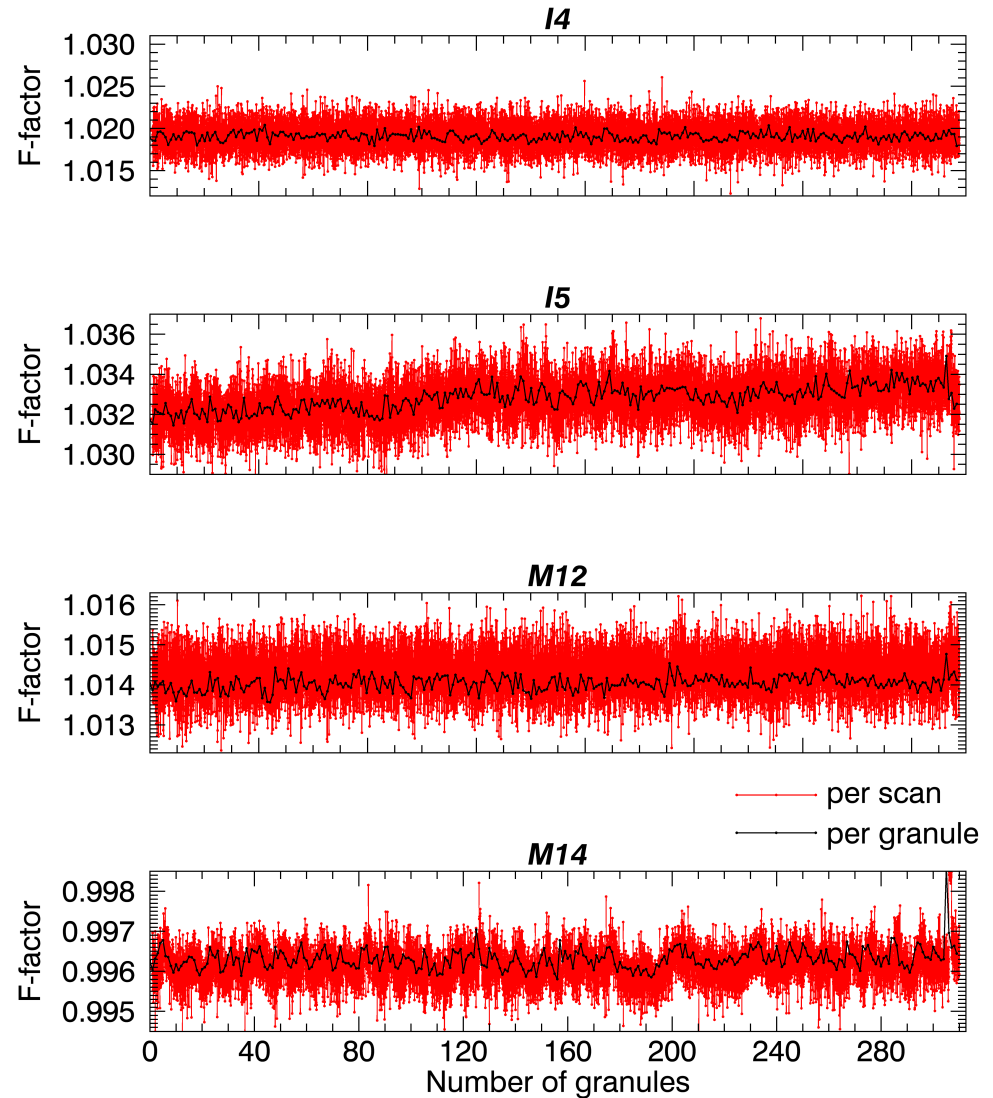


# Telemetry temperatures (Daily-avg.)



- Out-of-character features coincide with Earth passage through perihelion
- Small SMWIR focal plane temp. increasing trend that can be seen in MWIR TEBs (i.e. I4, M12, M13)

- Detector responses show small orbital variations
- ◆  $\pm 0.2$  % or less on a scan-by-scan basis
- ◆  $\pm 0.1$  % or less on a per granule basis
- F-factor fluctuations can be correlated to  $T_{BB}$  and instrument temperature variations
- Orbits 41375-41394

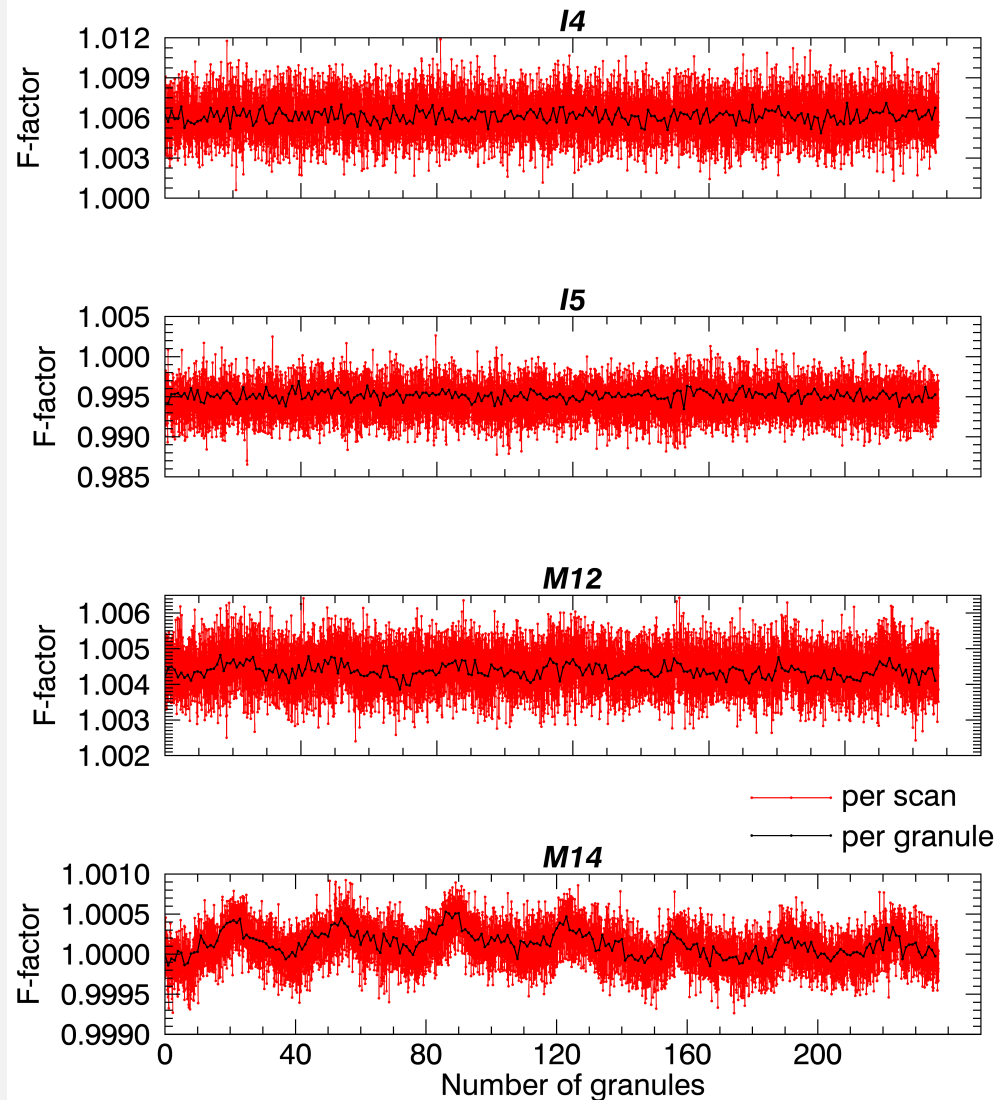


➤ Detector responses show small orbital variations

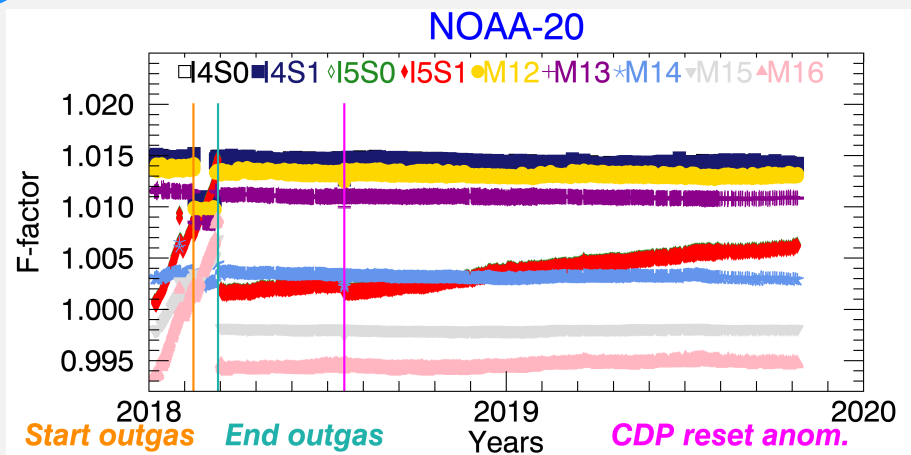
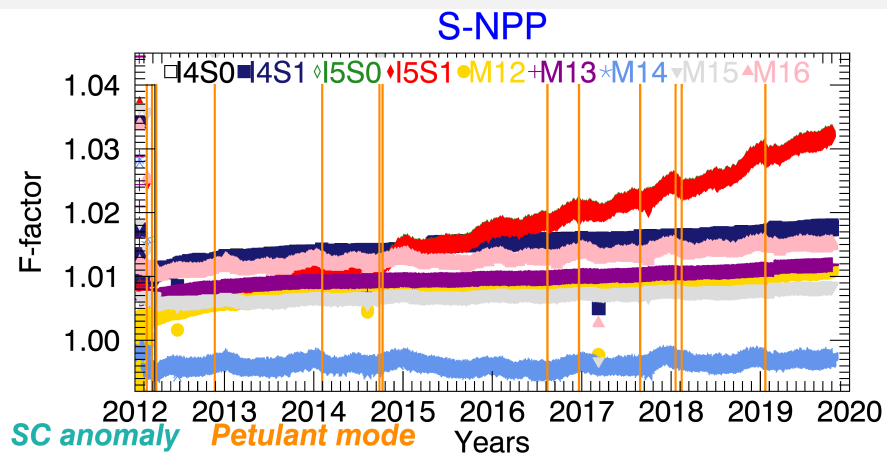
◆  $\pm 0.2$  % or less on a scan-by-scan basis

◆  $\pm 0.1$  % or less on a per granule basis

➤ Orbits 10000-10012



# Long-term detector response

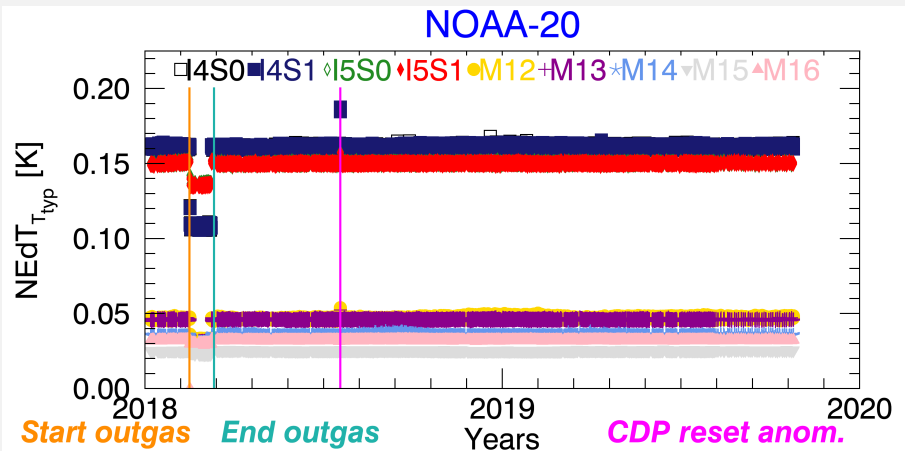
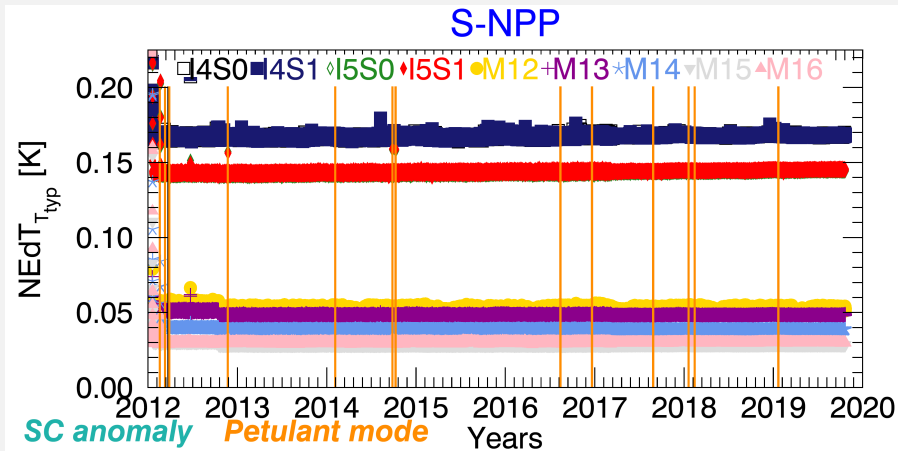


Satellite	TEB	I4	I5	M12	M13	M14	M15	M16
S-NPP	Avg. F-factor 03-26-2012	1.0105	1.0040	1.0035	1.0070	0.9946	1.0056	1.0101
	Avg. F-factor 10-22-2019	1.0176	1.0323	1.0110	1.0119	0.9972	1.0084	1.0152
N20	Avg. F-factor 04-17-2018	1.0145	1.0018	1.0134	1.0112	1.0035	0.9981	0.9942
	Avg. F-factor 10-25-2019	1.0139	1.0062	1.0130	1.0109	1.0029	0.9981	0.9948

**TEB I5 upward trend is approximately 0.35 % / yr for both VIIRS instruments. All other TEBs exhibit relatively stable trends.**

★ Band- and daily-averaged

# Noise characterization (NEdT)



$$NEdT = \frac{NEdL}{\partial L / \partial T}$$

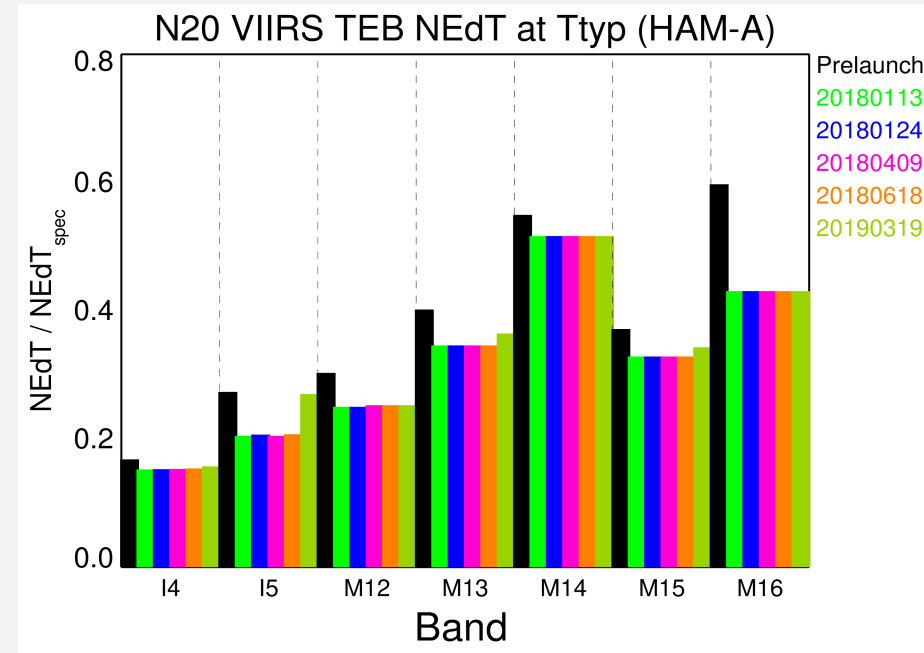
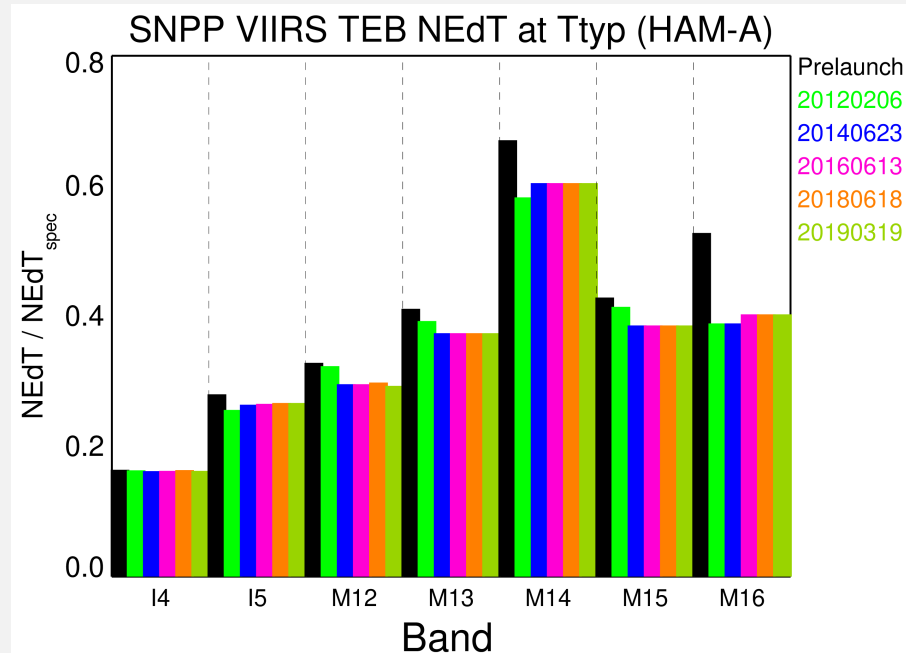
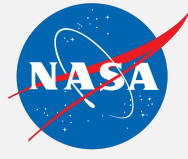
- NEdT trended routinely at 292.5 K
- Stable for both instruments
- Band-averaged values
  - - I-bands: ~ 0.2 K
  - - M-bands: ~ 0.07 K

★ Band- and daily-averaged

TEB	T <sub>typ</sub> (K)	Spec. (K)
<b>I4</b>	<b>270</b>	<b>2.5</b>
<b>I5</b>	<b>210</b>	<b>1.5</b>
<b>M12</b>	<b>270</b>	<b>0.396</b>
<b>M13</b>	<b>300</b>	<b>0.107</b>
<b>M14</b>	<b>270</b>	<b>0.091</b>
<b>M15</b>	<b>300</b>	<b>0.070</b>
<b>M16</b>	<b>300</b>	<b>0.072</b>



# Noise characterization (NEdT)

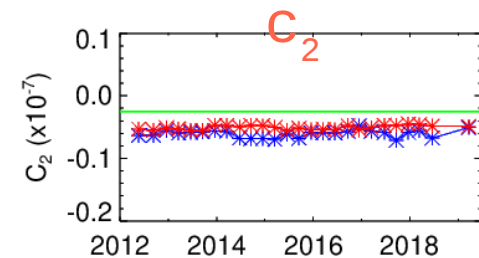
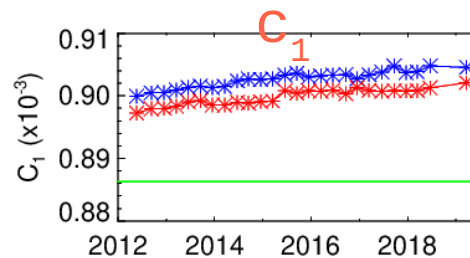
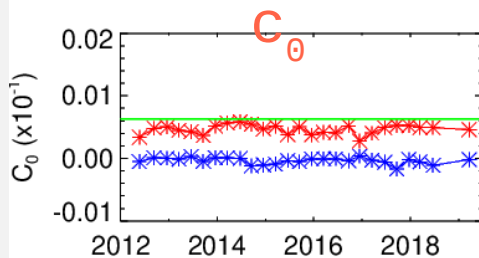


→  $NEdT_{T_{typ}}$  derived from BB CD data – TEBs meet sensor design req.

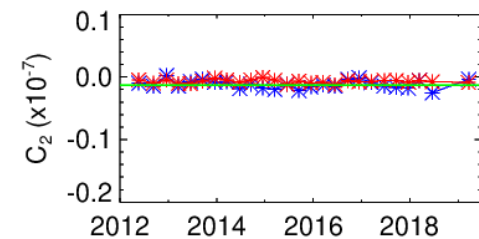
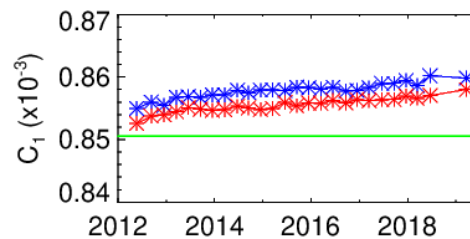
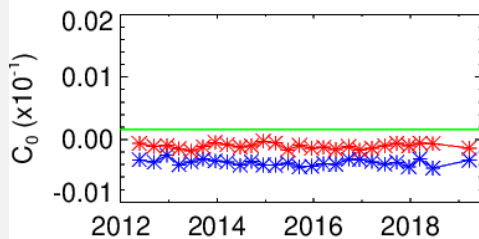


# Calibration coefficients (S-NPP)

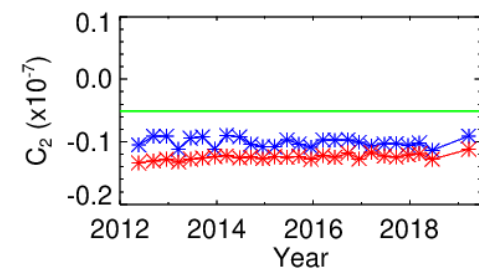
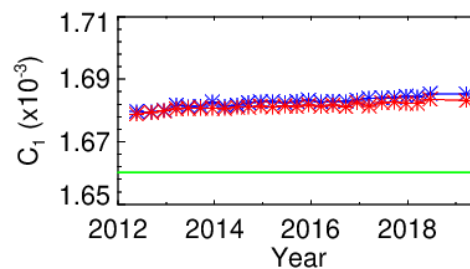
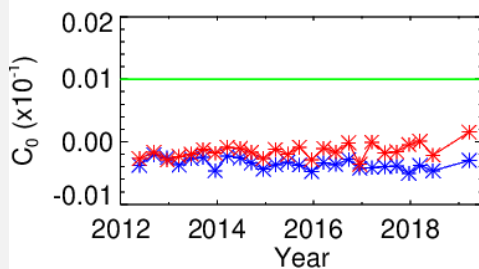
I4



M12



M13



LUT

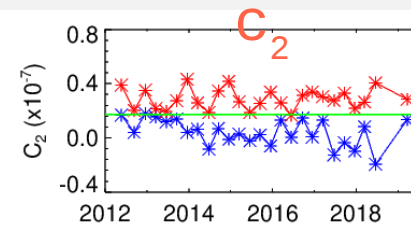
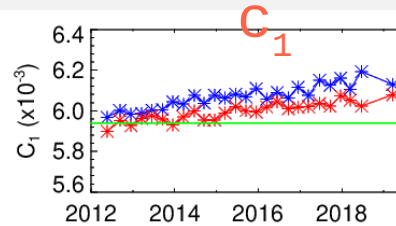
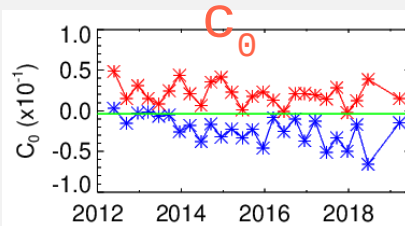
cool-down

warm-up

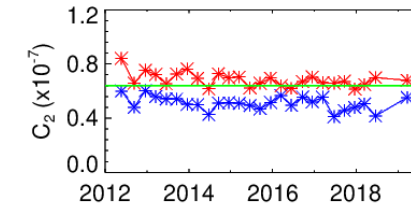
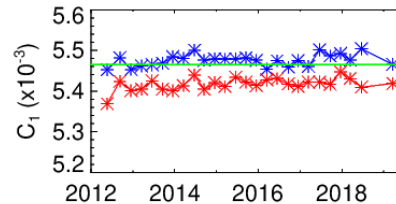
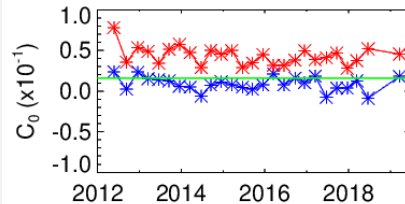
- Band-avg.  $c_0$ ,  $c_1$ , and  $c_2$  calibration coeffs. Derived from 27 WUCD through November 2019.
- Band-avg.  $c_1$  coeffs. are within 1.9 % from pre-launch

# Calibration coefficients (S-NPP)

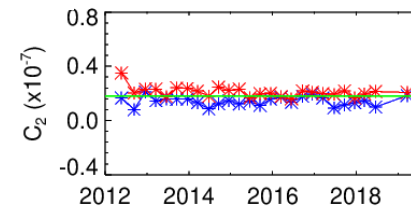
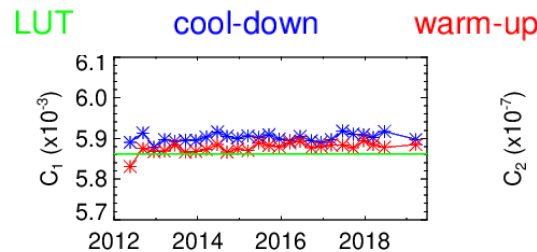
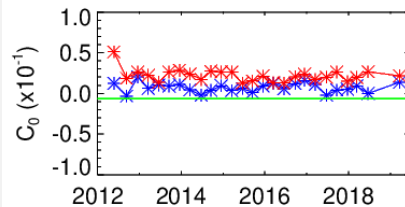
I5



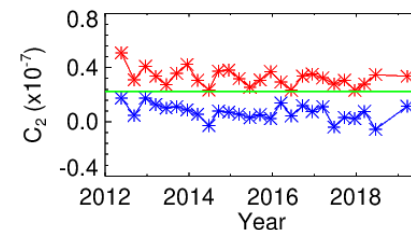
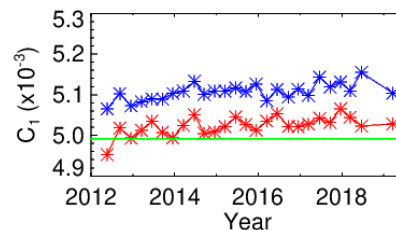
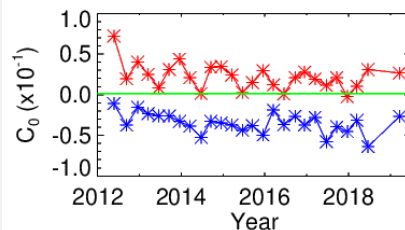
M14



M15



M16

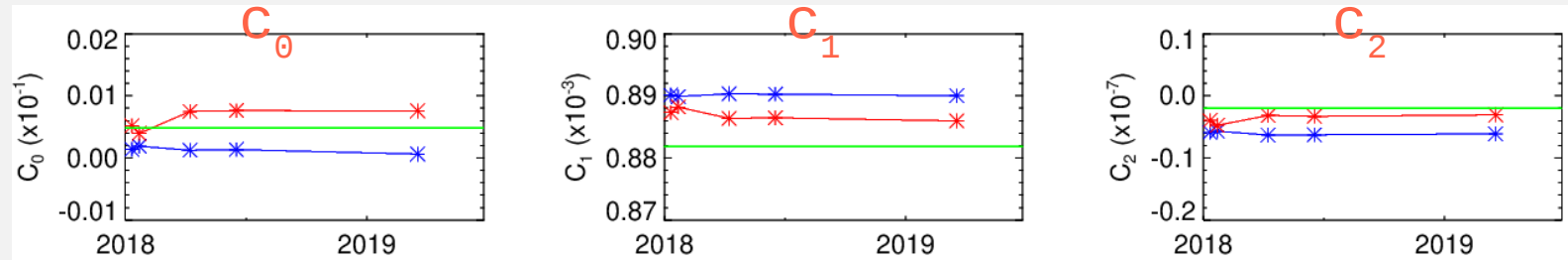


- $C_{0LUT} \pm 0.002$  (MWIR),  $\pm 0.01$  (LWIR)

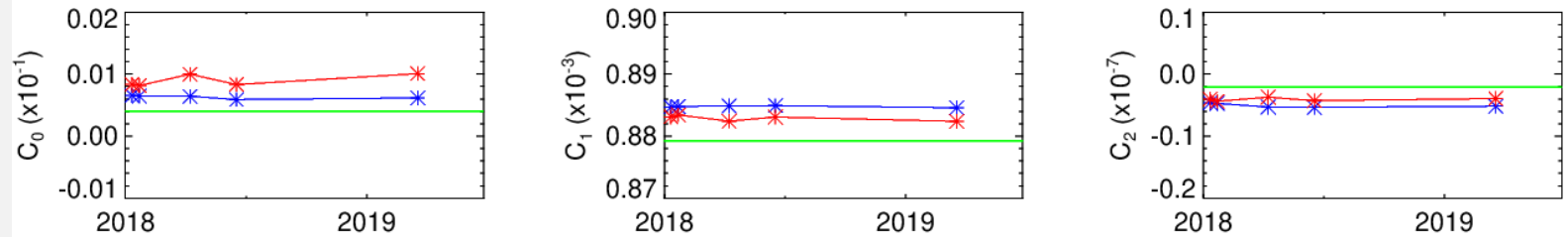
$$C_{2LUT} \pm 3 \cdot C_{2LUT}$$

# Calibration coefficients (N20)

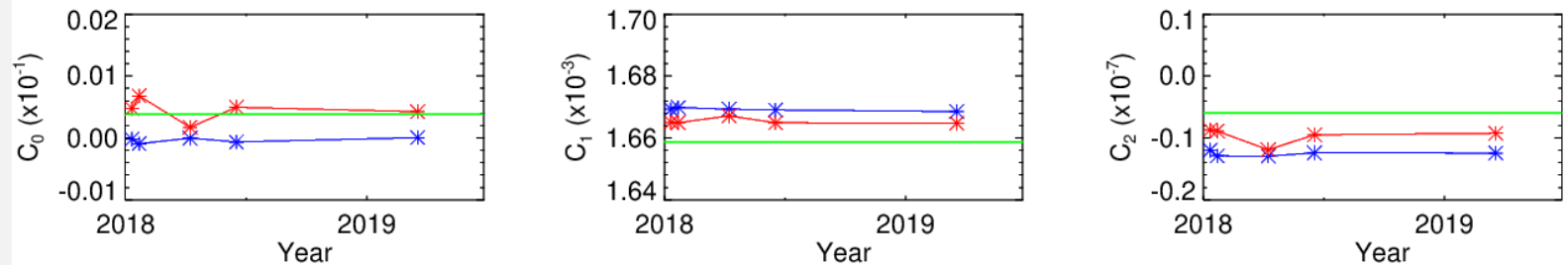
I4



M12



M13



LUT

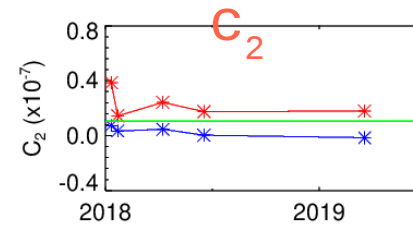
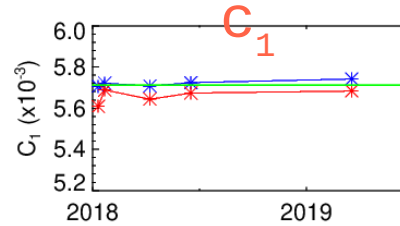
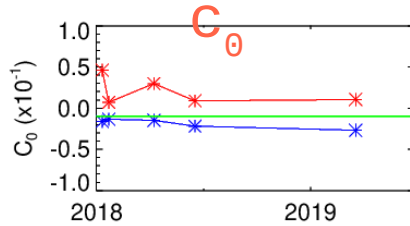
cool-down

warm-up

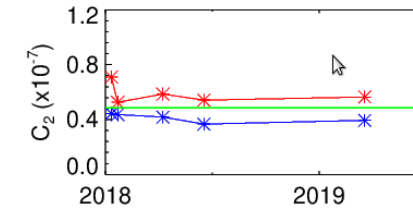
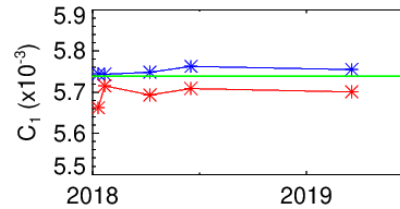
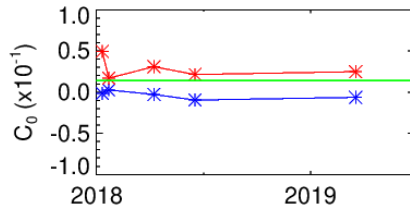
- Band-avg.  $c_0$ ,  $c_1$ , and  $c_2$  calibration coeffs. Derived from 5 WUCD through November 2019.
- Band-avg.  $c_1$  coeffs. are within 3.0 % from pre-launch

# Calibration coefficients (N20)

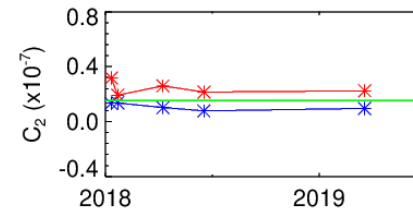
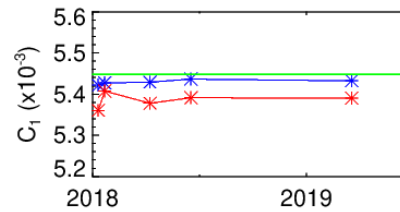
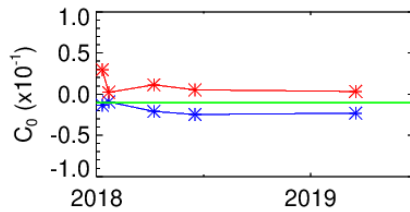
I5



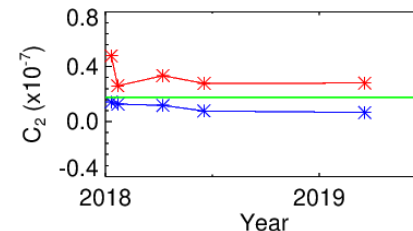
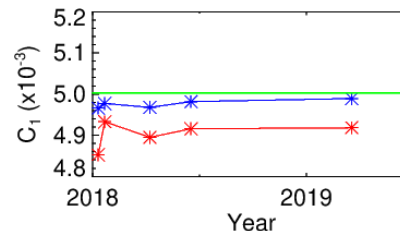
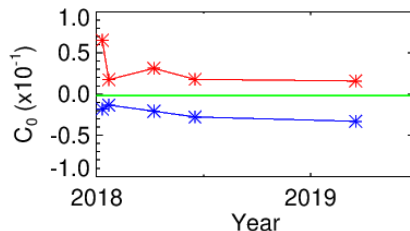
M14



M15



M16



- $C_{0LUT} \pm 0.002$  (MWIR),  $\pm 0.01$  (LWIR)

$$C_{2LUT} \pm 5 \cdot C_{2LUT}$$



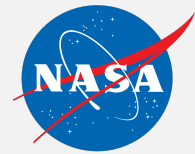
# NASA L1B TEBs improvements



- **NASA L1B TEBs improvements**
  - ◆ **Modified scaled integer table for M13 to increase radiometric resolution at low radiance / BT**
  - ◆ **Updated S-NPP VIIRS Delta-C LUT with correct temperature dependence to reprocess**
  
- **NASA L1B TEBs improvements in progress**
  - ◆ **Uncertainty algorithm in the L1B products**
  - ◆ **N20 pre-launch RVS LUT update**



# Summary



- ◆ On-orbit BB long-term performances for S-NPP (~8 years) and JPSS-1 (~2 years) VIIRS are quite stable. Short-term (orbital) temperature variations are present, but generally within the uniformity requirement of 30 mK.
- ◆ Detector response (F-factor) trending is stable for both S-NPP and JPSS-1 VIIRS. S-NPP VIIRS TEB I5 shows the maximum band-average trend of 2.8 %, followed by M12 and I4. JPSS-1 VIIRS TEB I5 shows a maximum trend of 0.5 %. Small orbital variations ( $\pm 0.05 - 0.1$  %) are present for both instruments.
- ◆ The TEBs detector noise characteristics are stable for both instruments. The NEdT at  $T_{\text{TYP}}$  is compliant with the requirements.