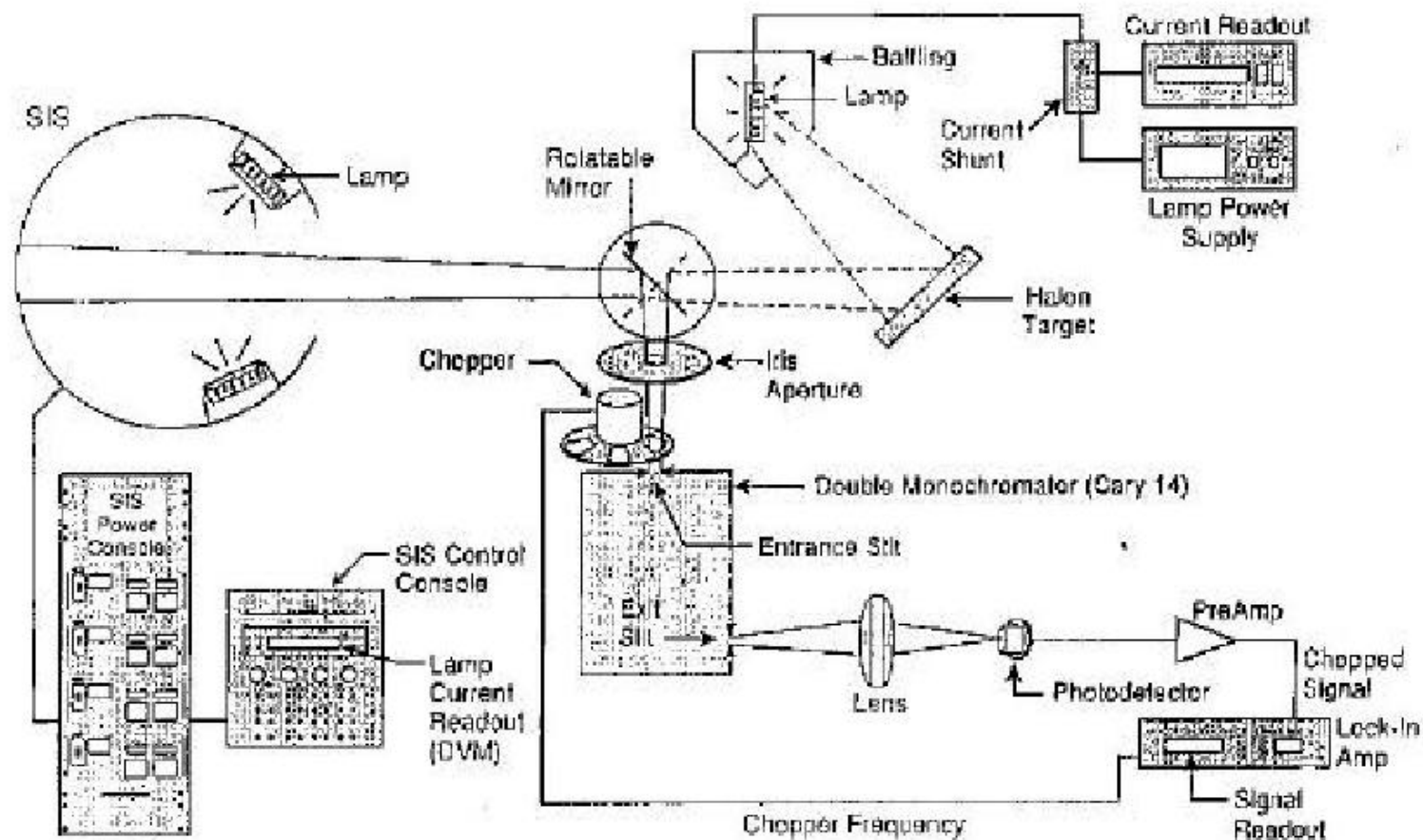


# RADIOMETRIC CHARTS

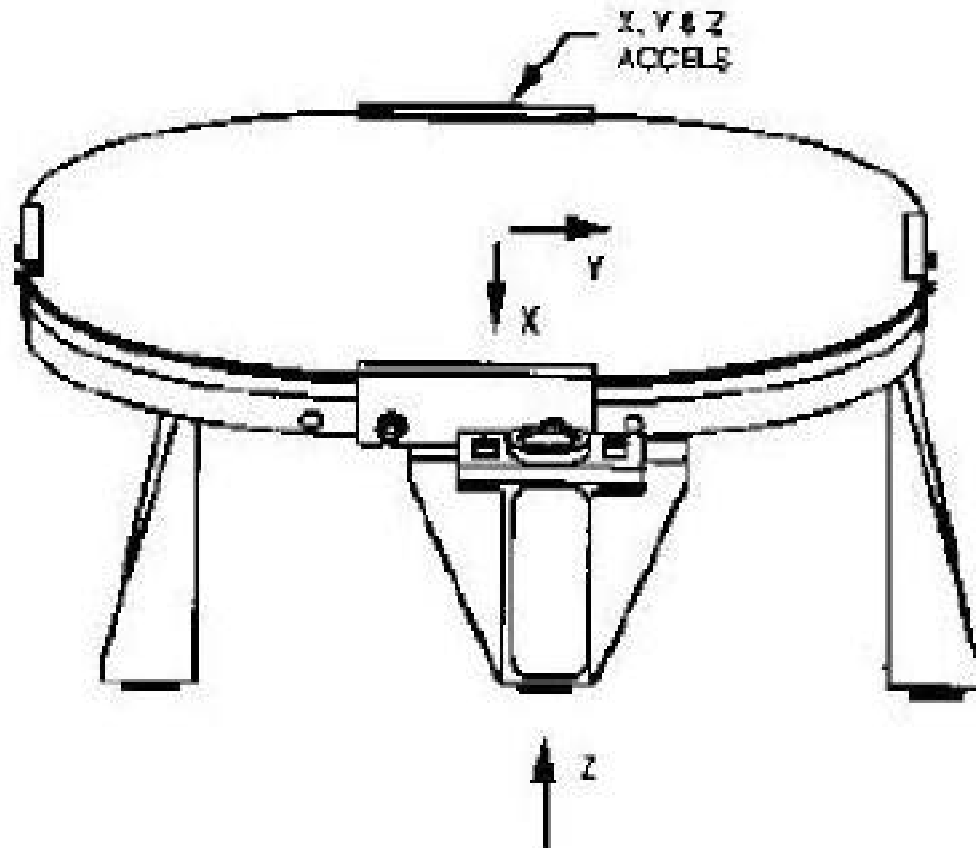
# CONFIGURATION FOR IRRADIANCE-BASED CALIBRATION OF SIS(100)



7/11/83

FIGURE 1

## Solar Diffuser (SD)



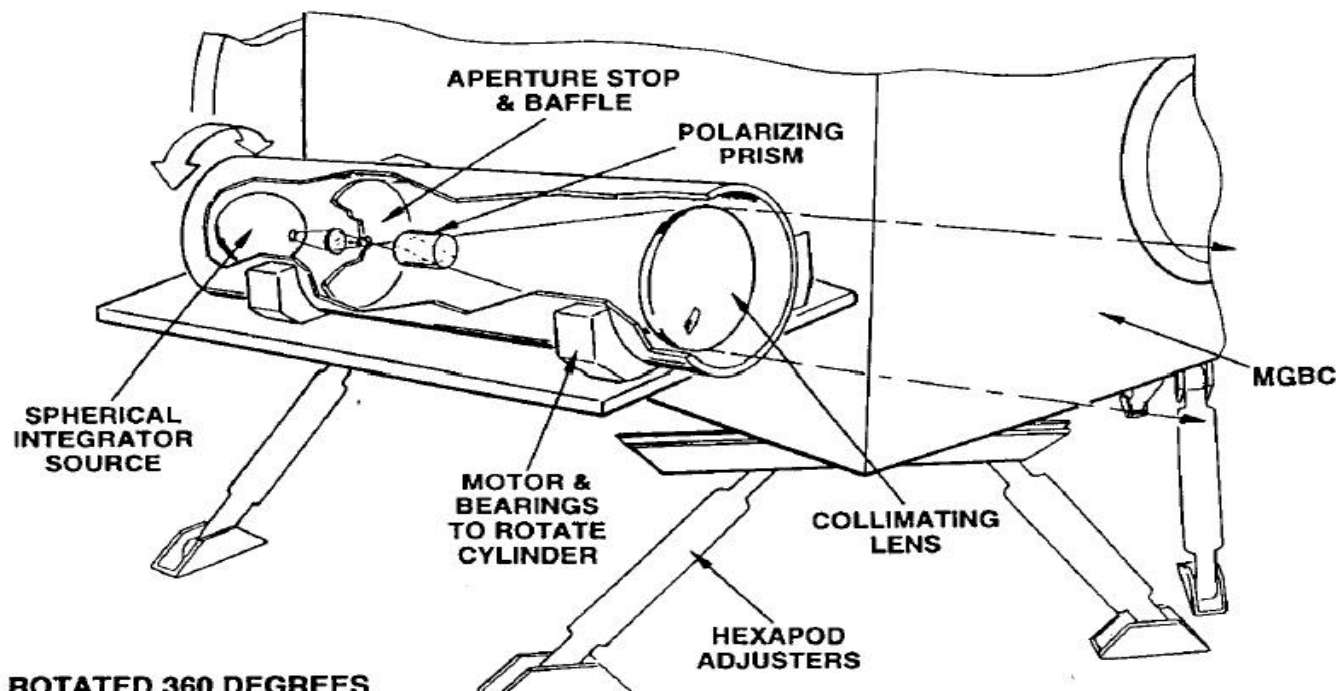
Space grade Spectralon (PTFE) by Labshpere  
11.25' by 9.32' by .25' with round corners (R3.5')



## MODIS POLARIZANCE IS CHARACTERIZED WITH A STAND-ALONE POLARIZER SUBSYSTEM

**HUGHES**

SANTA BARBARA RESEARCH CENTER  
a subsidiary



RADIOMETRIC

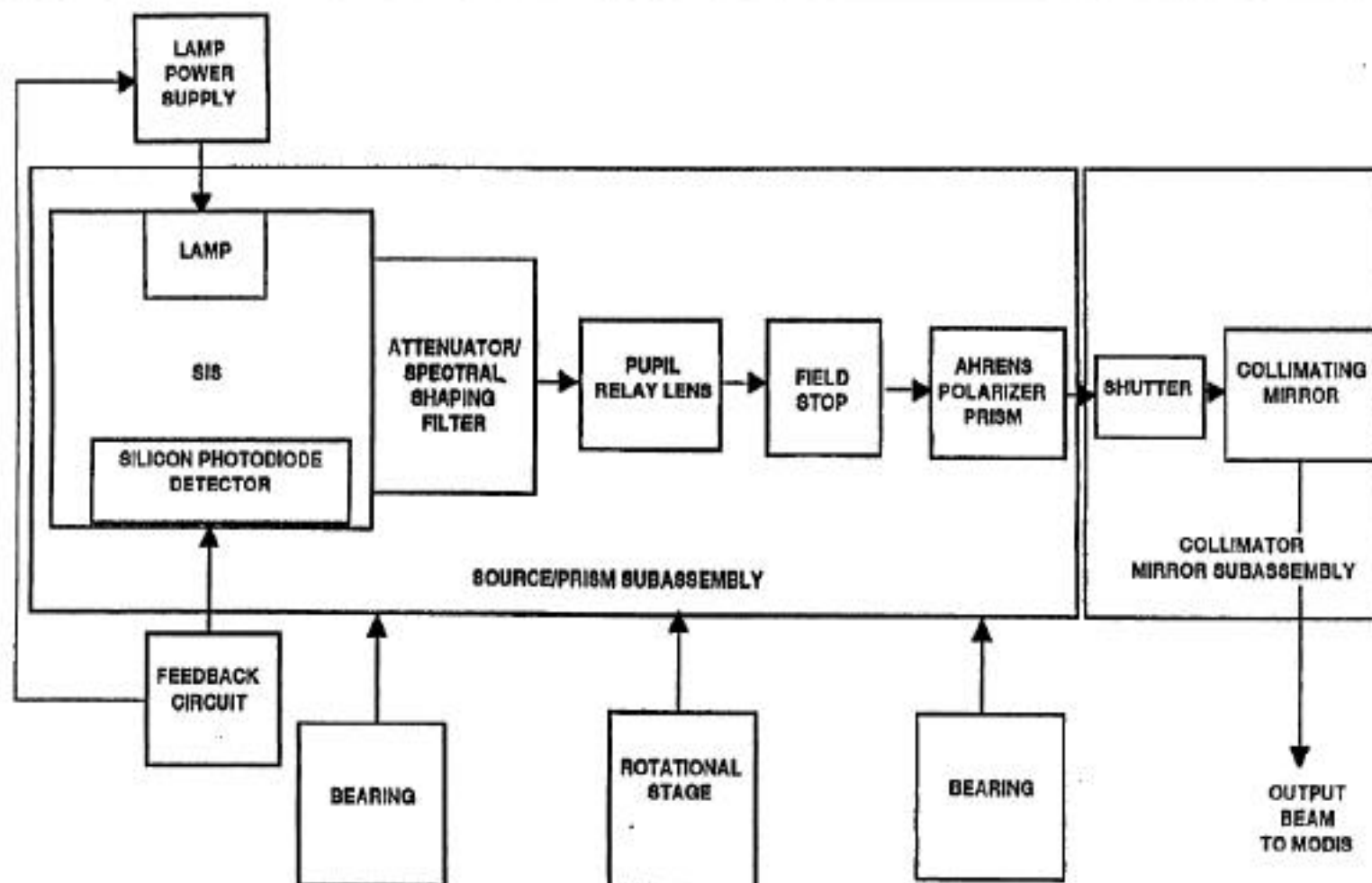


# PSA FUNCTIONAL BLOCK DIAGRAM

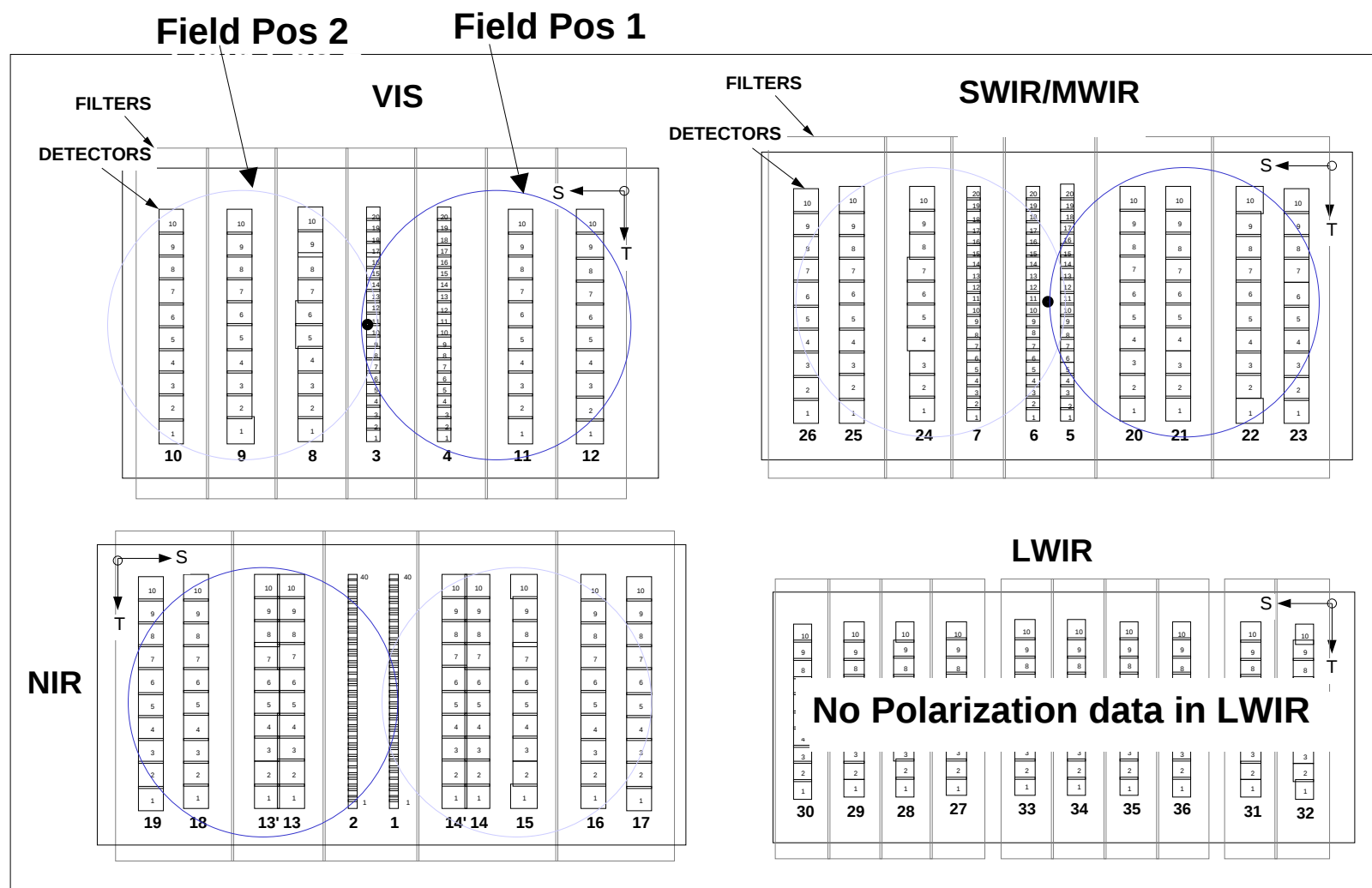
SANTA BARBARA  
RESEARCH  
CENTER



A HUGHES ELECTRONICS COMPANY



# DATA ACQUIRED IN 2 FIELD POSITIONS



RADIOMETRIC

Figure 4 illustrates the effect of polarization on the shape of band 8 relative spectral response.

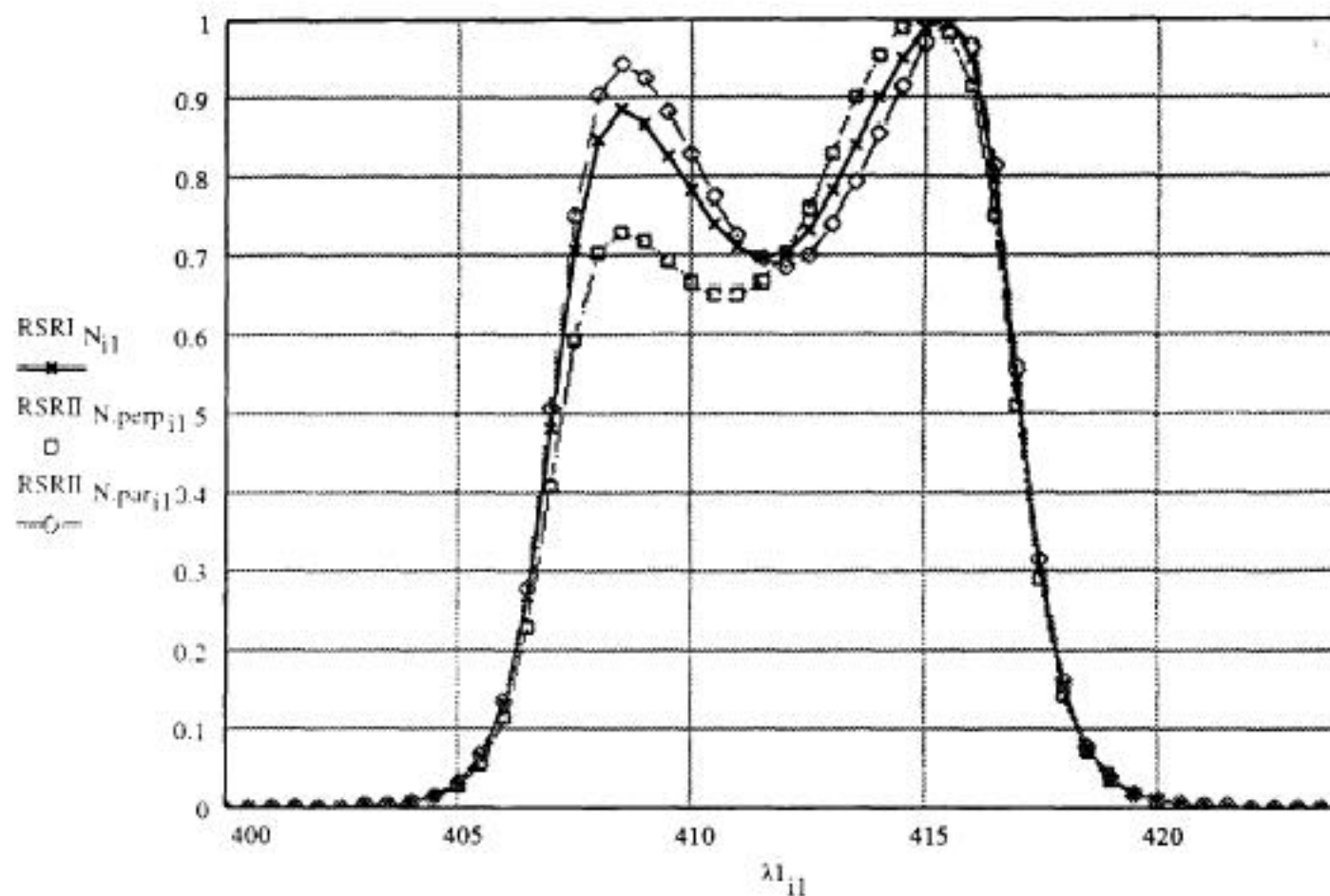
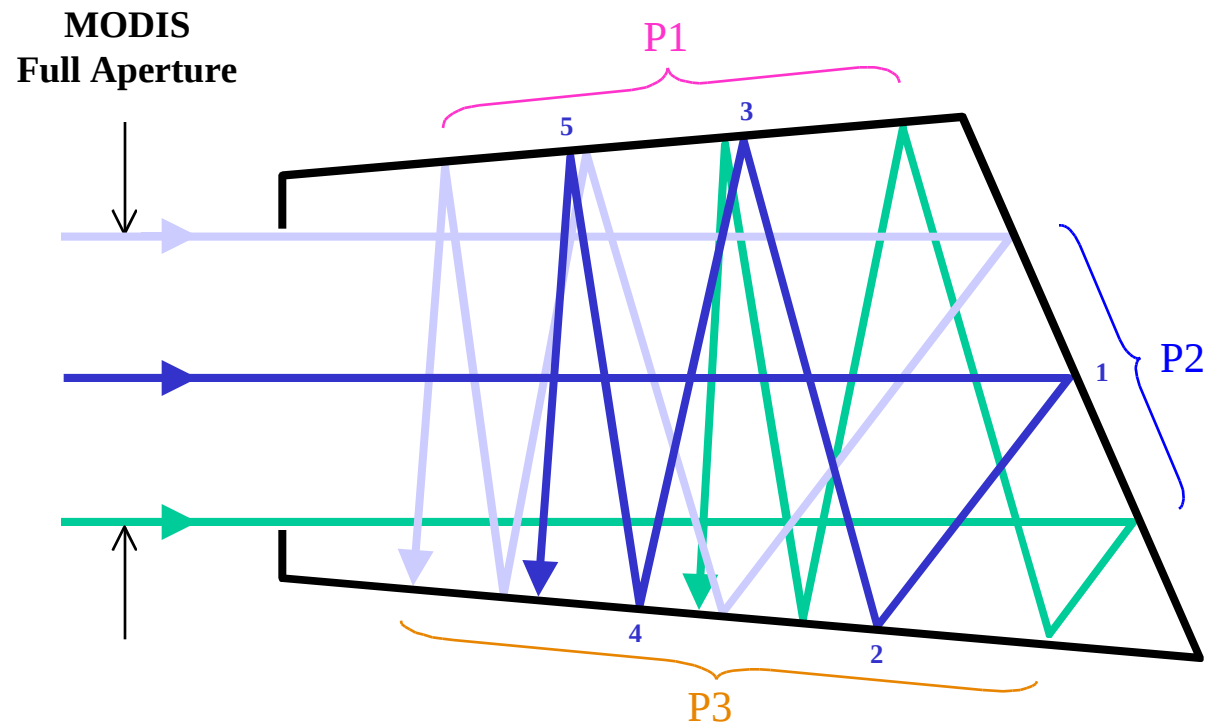


Figure 4. The simulated relative spectral response functions for case I (unpolarized) and case II (polarized).



# Blackbody Calibration Source (BCS)

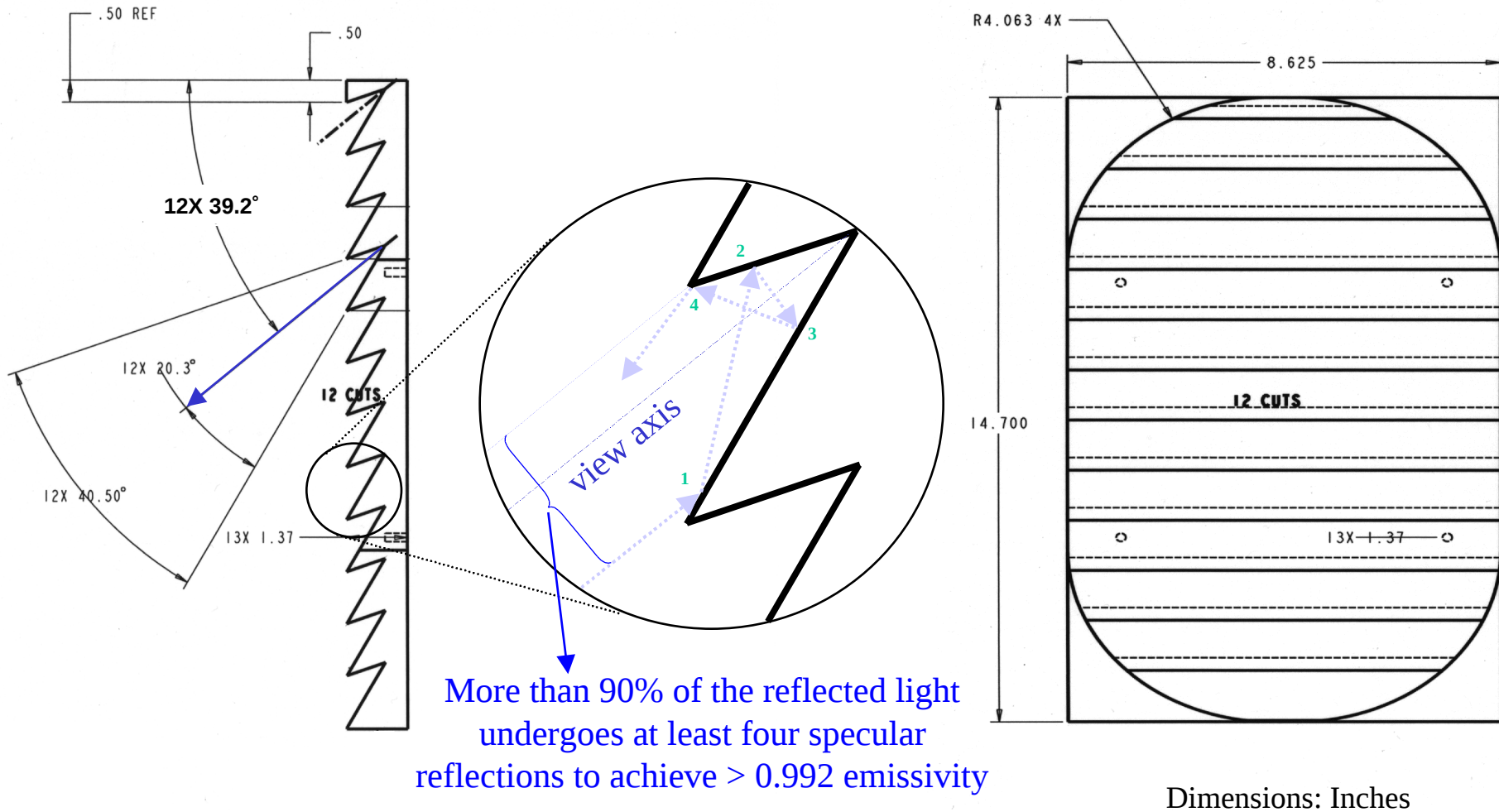


$$L_{\text{BCS}} = W_{\text{P1}} \cdot L(l, T_{\text{P1}}) + W_{\text{P2}} \cdot L(l, T_{\text{P2}}) + W_{\text{P3}} \cdot L(l, T_{\text{P3}})$$

BCS Trapezoid Configuration Achieves **> 0.9998** Emissivity (From SBRS)  
RADIOMETRIC



# The MODIS On-Board Calibrator (OBC) Blackbody



RADIOMETRIC

CAL-9

## DATA FORMAT ISSUE

- THE FIRST ARE LAST and  
THE LAST ARE FIRST
  - SBRS detector numbering convention not consistent with COTS Earth-mapping conventions
  - L1A rennumbers to COTS-friendly (pixel) order
  - MCST characterization work done in SBRS (detector) numbering conventions
  - L-7 IR band (some channels) saturated for ~ half year due to this issue

NOTE: All the MCST Look-up Tables (LUTs) are published in the SBRS-detector numbering convention and this detector numbering convention is inverted from the pixel convention used in the MODIS Level-1 products. To investigate an “anomaly” in pixel 7 of Band 10 in L1 product, must look at characterization data for detector 3 of that Band.

# Detector Number Convention for MODIS Sensor

|                                                          |  |                    |    |    |    |    |    |    |        |
|----------------------------------------------------------|--|--------------------|----|----|----|----|----|----|--------|
|                                                          |  | LWIR (abbreviated) |    |    |    |    |    |    |        |
| Band                                                     |  | 30                 | 29 | 28 | 27 | 33 | 34 | 35 |        |
| Scan 1                                                   |  | 10                 | 10 | 10 | 10 | 10 | 10 | 10 | ↓<br>T |
|                                                          |  | 9                  | 9  | 9  | 9  | 9  | 9  | 9  |        |
|                                                          |  | 8                  | 8  | 8  | 8  | 8  | 8  | 8  |        |
|                                                          |  | 7                  | 7  | 7  | 7  | 7  | 7  | 7  |        |
|                                                          |  | 6                  | 6  | 6  | 6  | 6  | 6  | 6  |        |
|                                                          |  | 5                  | 5  | 5  | 5  | 5  | 5  | 5  |        |
|                                                          |  | 4                  | 4  | 4  | 4  | 4  | 4  | 4  |        |
|                                                          |  | 3                  | 3  | 3  | 3  | 3  | 3  | 3  |        |
|                                                          |  | 2                  | 2  | 2  | 2  | 2  | 2  | 2  |        |
|                                                          |  | 1                  | 1  | 1  | 1  | 1  | 1  | 1  |        |
| Scan 2                                                   |  | 20                 | 20 | 20 | 20 | 20 | 20 | 20 | ↓<br>T |
|                                                          |  | 19                 | 19 | 19 | 19 | 19 | 19 | 19 |        |
|                                                          |  | 18                 | 18 | 18 | 18 | 18 | 18 | 18 |        |
|                                                          |  | 17                 | 17 | 17 | 17 | 17 | 17 | 17 |        |
|                                                          |  | 16                 | 16 | 16 | 16 | 16 | 16 | 16 |        |
|                                                          |  | 15                 | 15 | 15 | 15 | 15 | 15 | 15 |        |
|                                                          |  | 14                 | 14 | 14 | 14 | 14 | 14 | 14 |        |
|                                                          |  | 13                 | 13 | 13 | 13 | 13 | 13 | 13 |        |
|                                                          |  | 12                 | 12 | 12 | 12 | 12 | 12 | 12 |        |
|                                                          |  | 11                 | 11 | 11 | 11 | 11 | 11 | 11 |        |
| Above numbering convention is SBRS "detector" convention |  |                    |    |    |    |    |    |    |        |

The arrow shows the track direction, which is the direction of satellite motion.

# Pixel Numbering Convention for L1B Product

| Below numbering convention is L1 Product "pixel" numbering convention |    |    |    |    |    |    |    |        |  |
|-----------------------------------------------------------------------|----|----|----|----|----|----|----|--------|--|
| LWIR (abbreviated)                                                    |    |    |    |    |    |    |    |        |  |
| Band                                                                  | 30 | 29 | 28 | 27 | 33 | 34 | 35 |        |  |
| Scan 1                                                                | 1  | 1  | 1  | 1  | 1  | 1  | 1  | ↓<br>T |  |
|                                                                       | 2  | 2  | 2  | 2  | 2  | 2  | 2  |        |  |
|                                                                       | 3  | 3  | 3  | 3  | 3  | 3  | 3  |        |  |
|                                                                       | 4  | 4  | 4  | 4  | 4  | 4  | 4  |        |  |
|                                                                       | 5  | 5  | 5  | 5  | 5  | 5  | 5  |        |  |
|                                                                       | 6  | 6  | 6  | 6  | 6  | 6  | 6  |        |  |
|                                                                       | 7  | 7  | 7  | 7  | 7  | 7  | 7  |        |  |
|                                                                       | 8  | 8  | 8  | 8  | 8  | 8  | 8  |        |  |
|                                                                       | 9  | 9  | 9  | 9  | 9  | 9  | 9  |        |  |
|                                                                       | 10 | 10 | 10 | 10 | 10 | 10 | 10 |        |  |
| Scan 2                                                                | 11 | 11 | 11 | 11 | 11 | 11 | 11 | ↓<br>T |  |
|                                                                       | 12 | 12 | 12 | 12 | 12 | 12 | 12 |        |  |
|                                                                       | 13 | 13 | 13 | 13 | 13 | 13 | 13 |        |  |
|                                                                       | 14 | 14 | 14 | 14 | 14 | 14 | 14 |        |  |
|                                                                       | 15 | 15 | 15 | 15 | 15 | 15 | 15 |        |  |
|                                                                       | 16 | 16 | 16 | 16 | 16 | 16 | 16 |        |  |
|                                                                       | 17 | 17 | 17 | 17 | 17 | 17 | 17 |        |  |
|                                                                       | 18 | 18 | 18 | 18 | 18 | 18 | 18 |        |  |
|                                                                       | 19 | 19 | 19 | 19 | 19 | 19 | 19 |        |  |
|                                                                       | 20 | 20 | 20 | 20 | 20 | 20 | 20 |        |  |

Charts showing the SBRs detector (previous chart) and L1 Product pixel (this chart) numbering conventions for consecutive scans at a specific frame or location across scan line, e.g. at nadir. The benefit of the L1 Product pixel convention is the logical progression of samples for mapping software. Sample X from Scan Y is labeled as " $X+10*(Y-1)$ "

The arrow shows the track direction, which is the direction of satellite motion.



## What is Electronic Cross-talk



- Readout integrating capacitor fails to return to baseline voltage following signal reset
- First realized nature of problem from RSR testing
  - Actually “seen” as negative dns & sub-frame differences, B5-7
  - Nearly impossible to detect when observing flat scenes; these calibrations in error, but cannot detect that
- Structured scene analysis in 1998 showed effects of 5-10% or more in SWIR bands & several percent in 4- $\mu$ m SST bands
- Observed in both SW/MWIR and PV LWIR bands (5-7, 20-30)
- Focus here on SW/MWIR bands (5-7, 20-26)

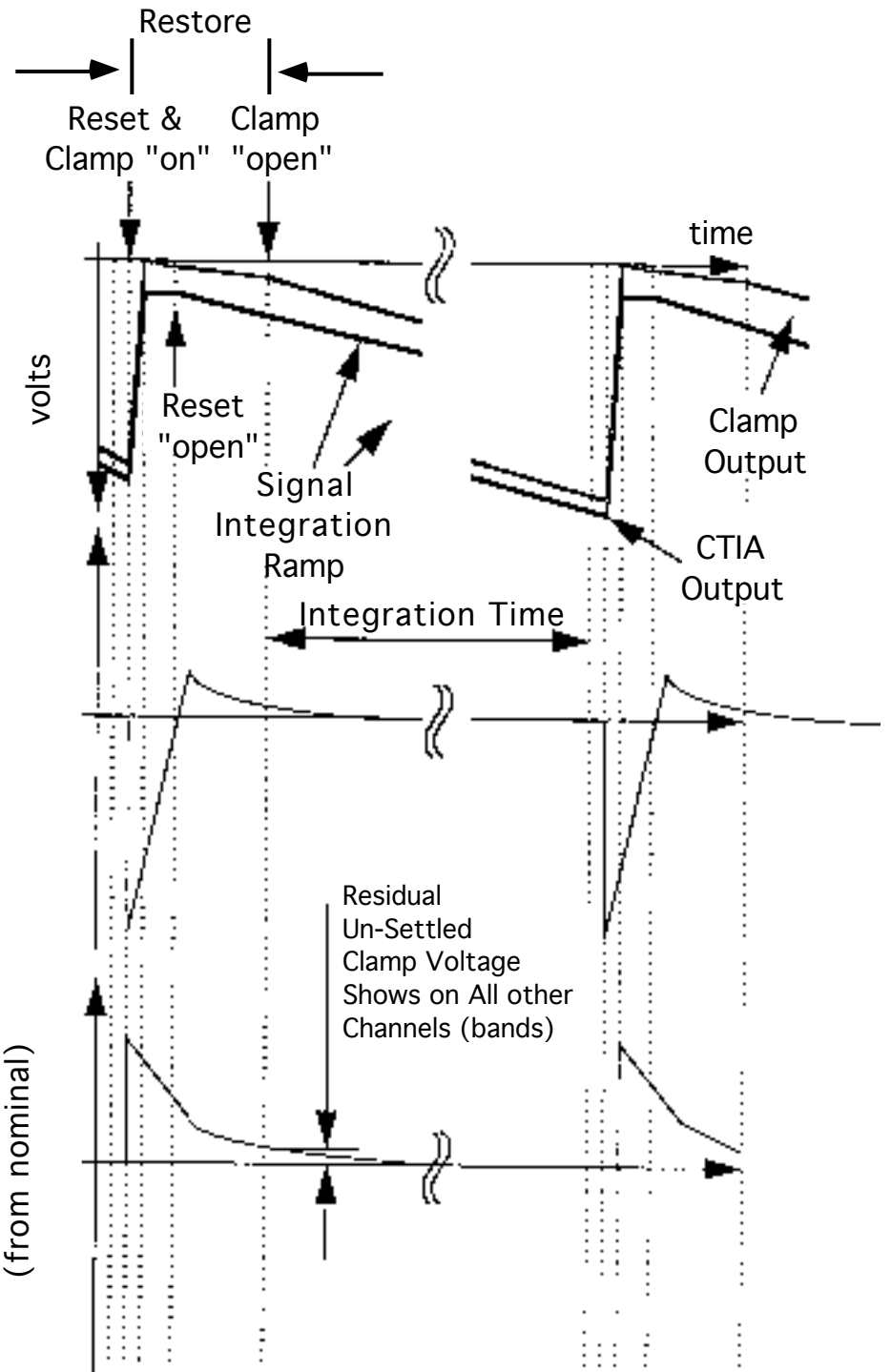


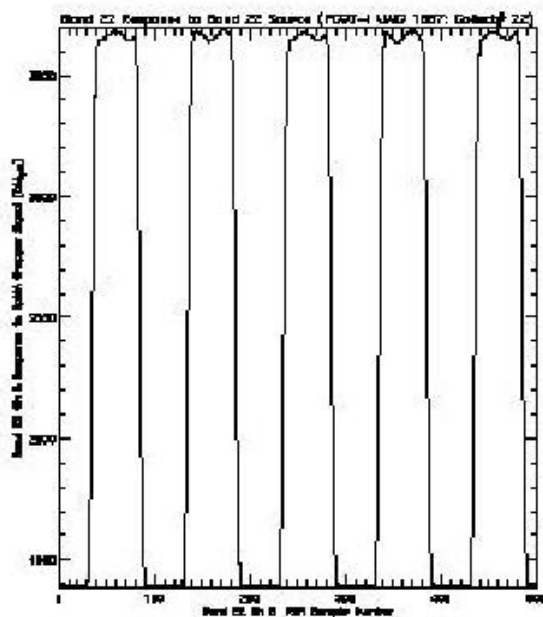
To buffer &  
Sample & Hold

mp

Clamp Current  
(from  $V_{\text{clamp}}$ )

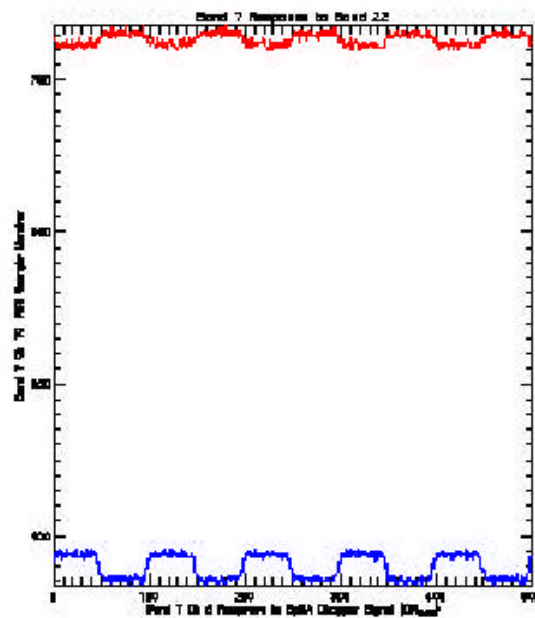
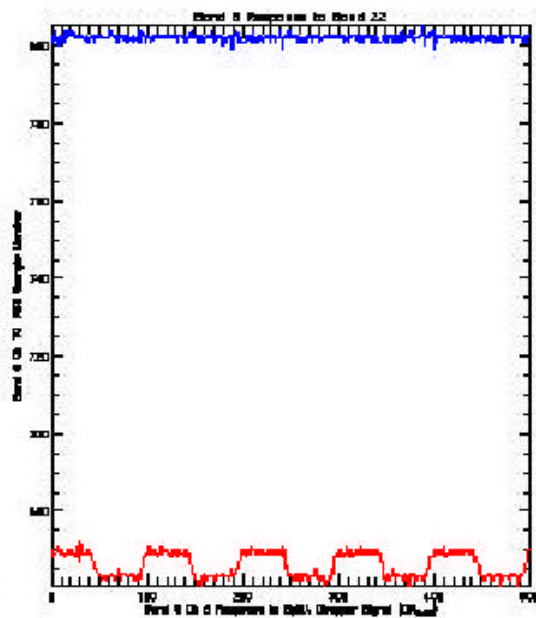
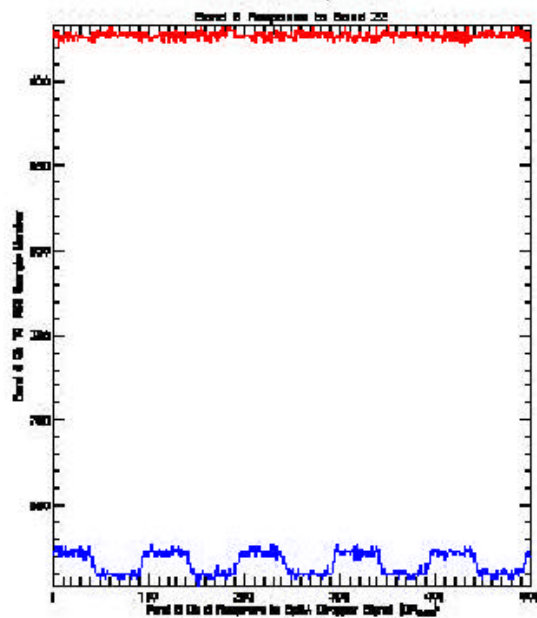
$V_{\text{clamp}}$   
(from nominal)

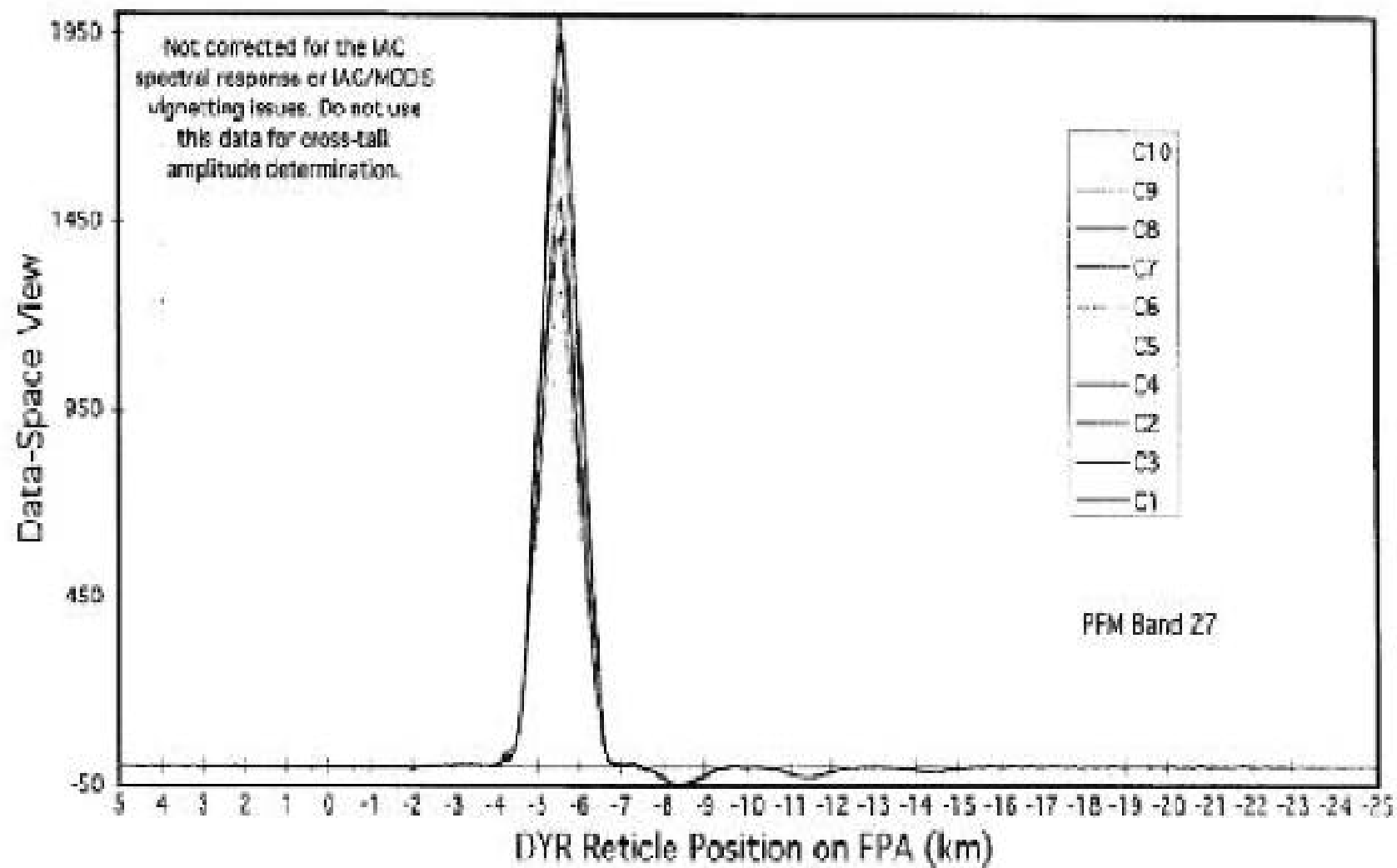


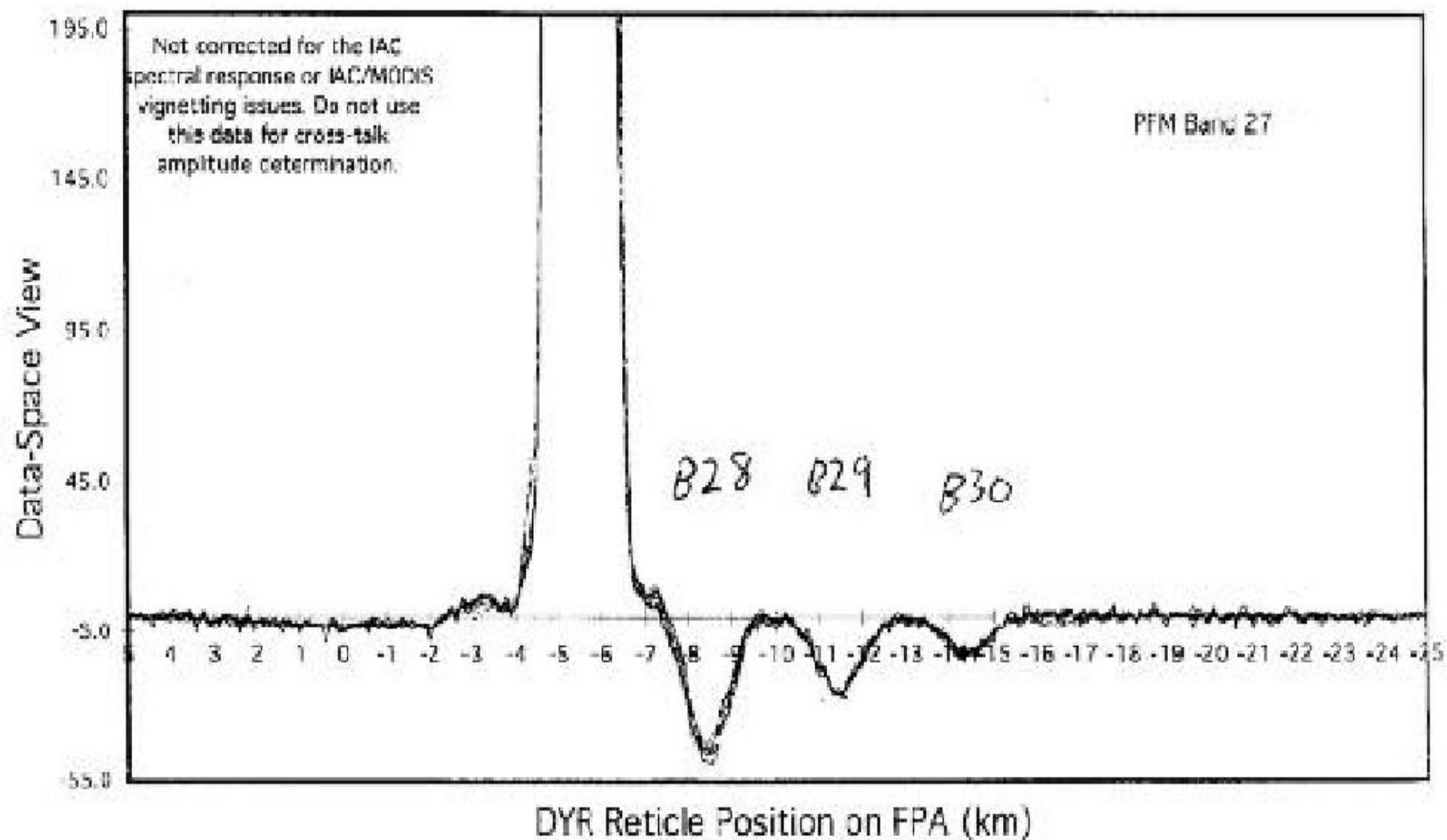


—— SF1  
 ---- SF2

$$\begin{aligned}
 L_{\text{SpMk}} &= 1.25 \text{ Watts/m}^2/\text{sr}/\mu\text{m} \\
 &= 0.66 L_{\text{noise}} \\
 &= 1.86 L_{\text{typo}} \\
 &= 1.88 L_{\text{cls udr}}
 \end{aligned}$$

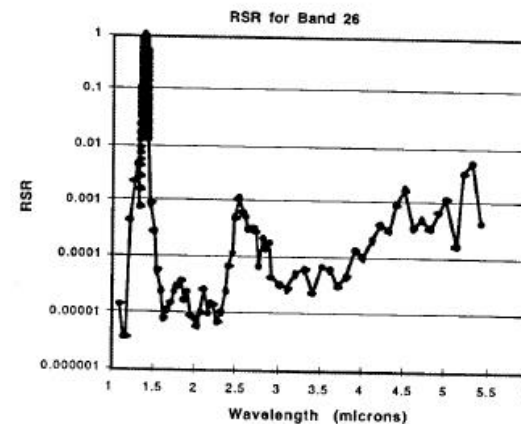
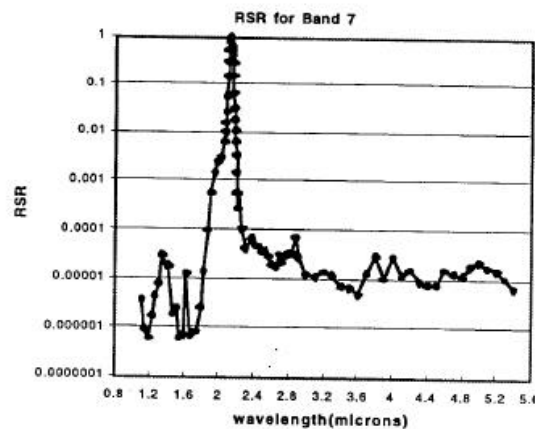
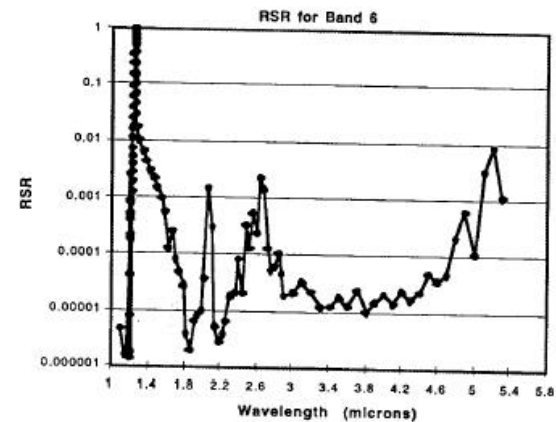
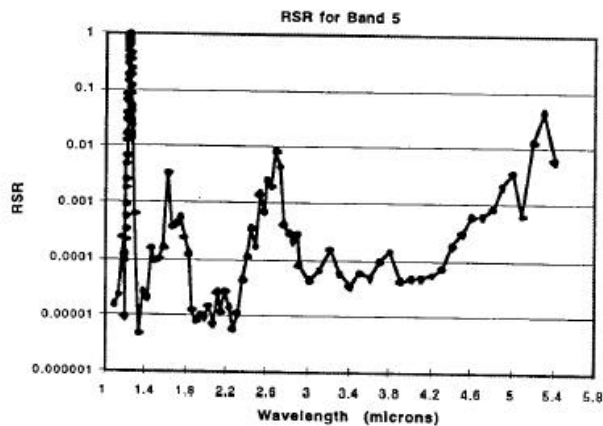








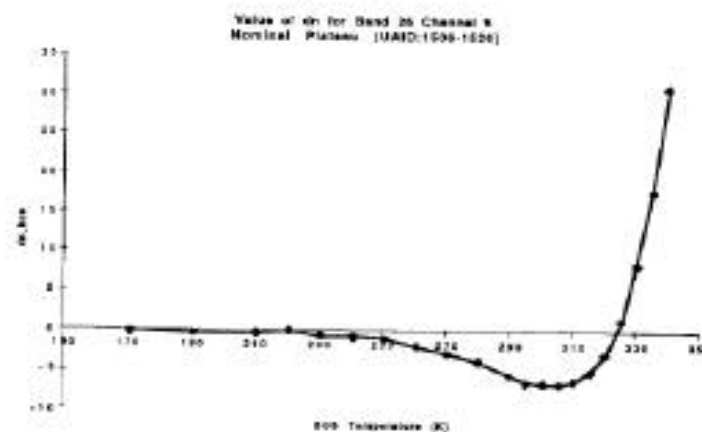
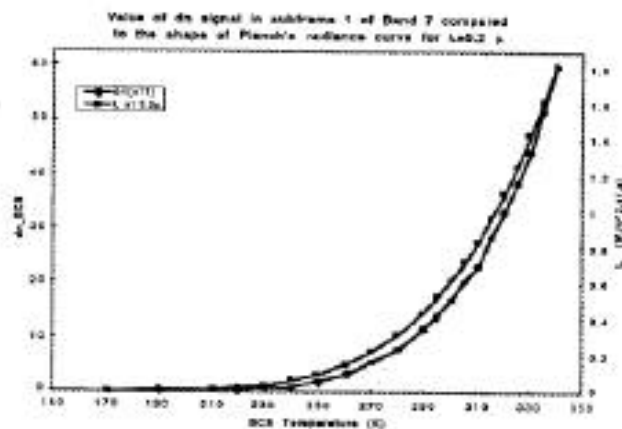
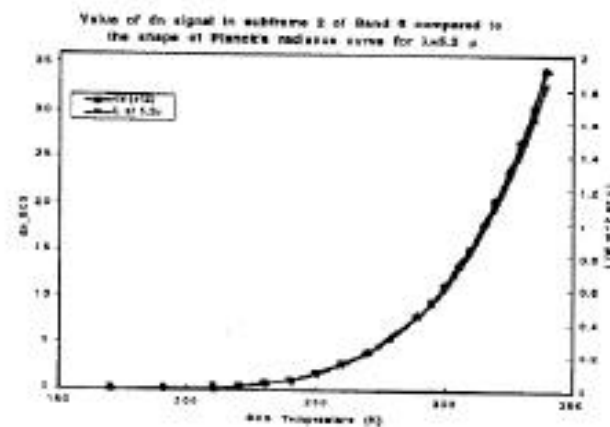
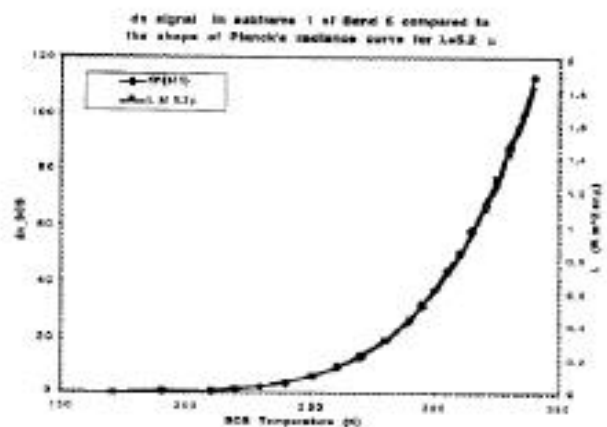
# Relative Spectral Response (RSR)



SWIR-5



# Behavior of SWIR Signal with BCS Temperature



SWIR-6

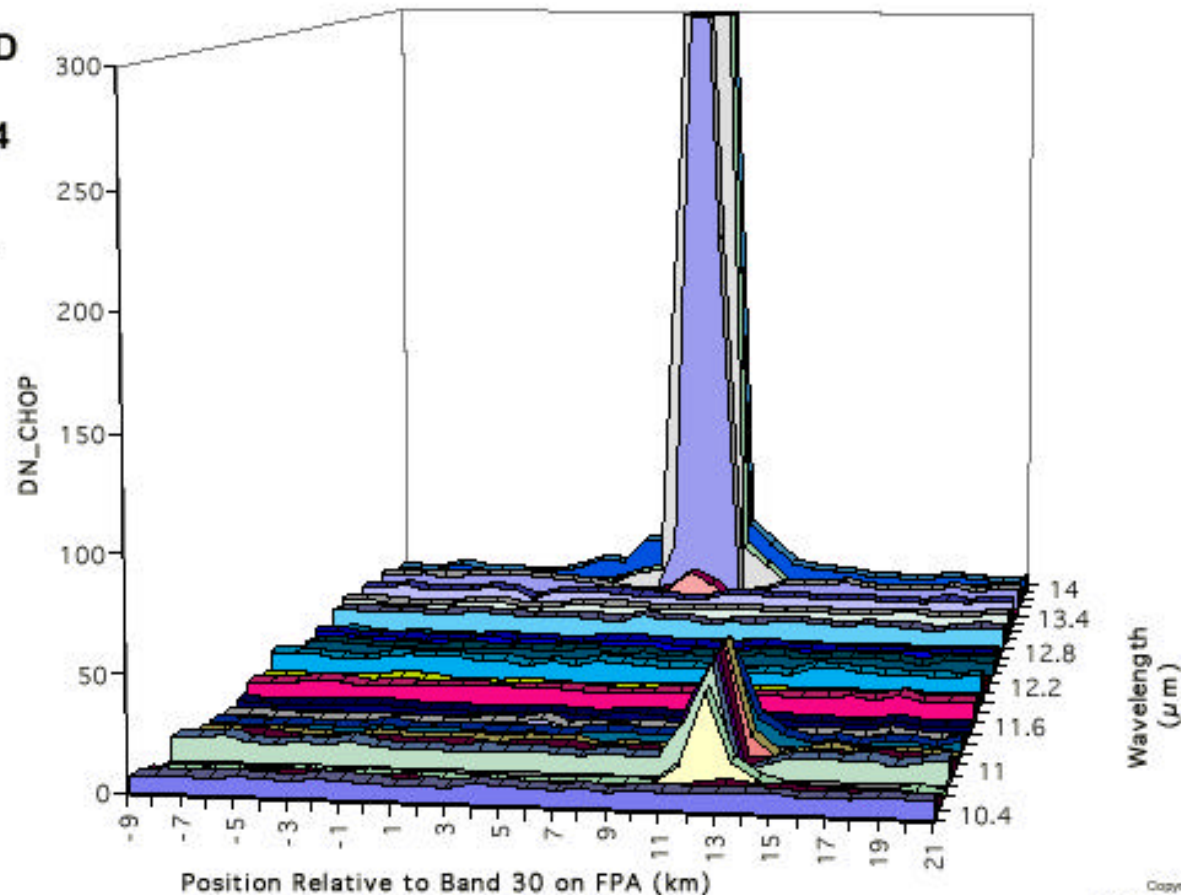


# Pre-launch Crosstalk Observation (Spatial/Spectral OOB Response Testing)



BAND 35 SPECTRAL AND SPATIAL RESPONSE

- DATA ACQUIRED DURING OOB SCAN FROM 10.4 TO 14  $\mu\text{m}$
- RESPONSE NOT VISIBLE IN BANDS 31-33

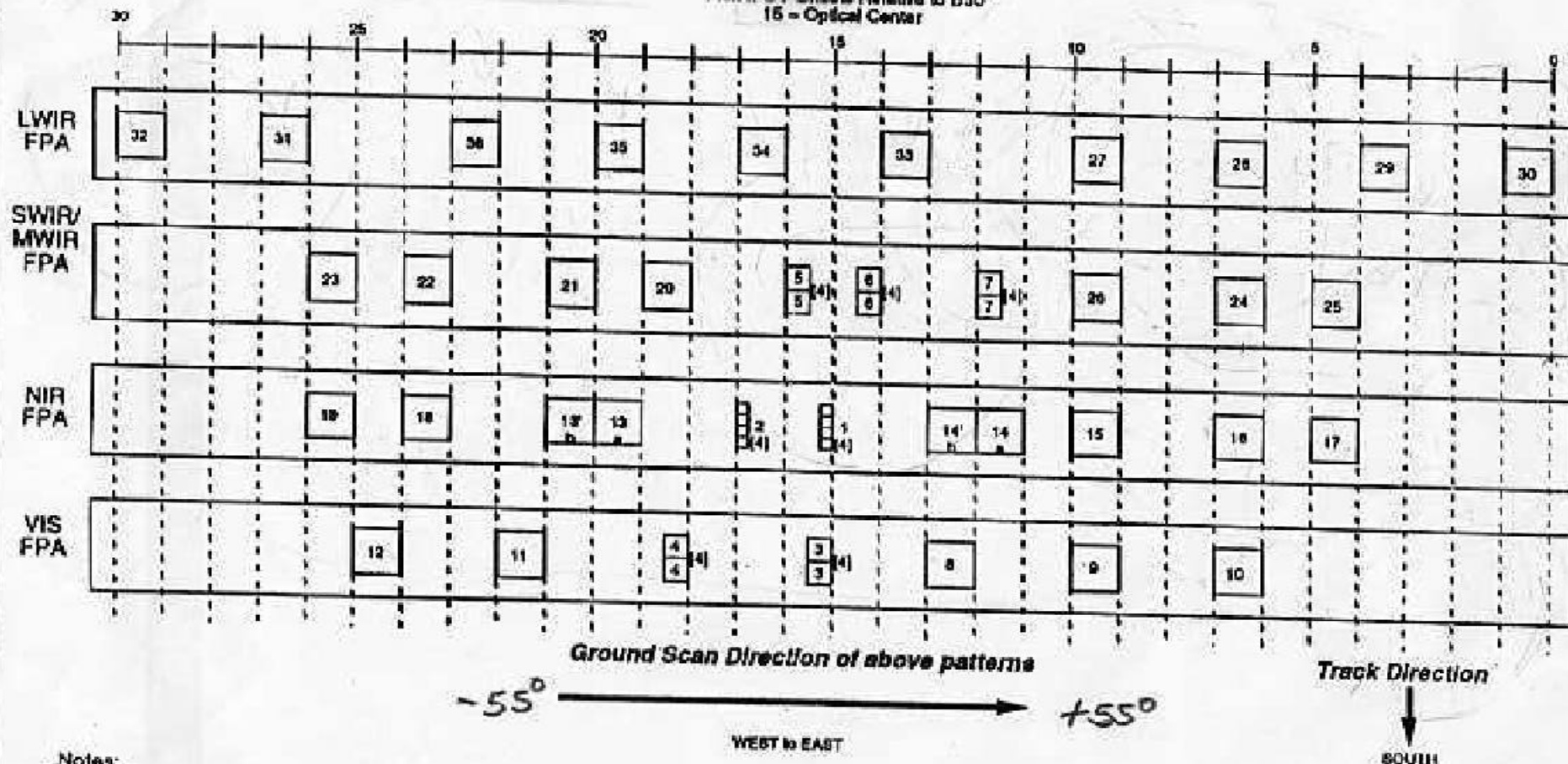


10/16/96

22

Copyright © 1996  
Hughes Aircraft Company  
Santa Barbara Remote Sensing  
An unpublished work. All rights reserved.

1 km IFOV Offsets Relative to B30  
15 = Optical Center

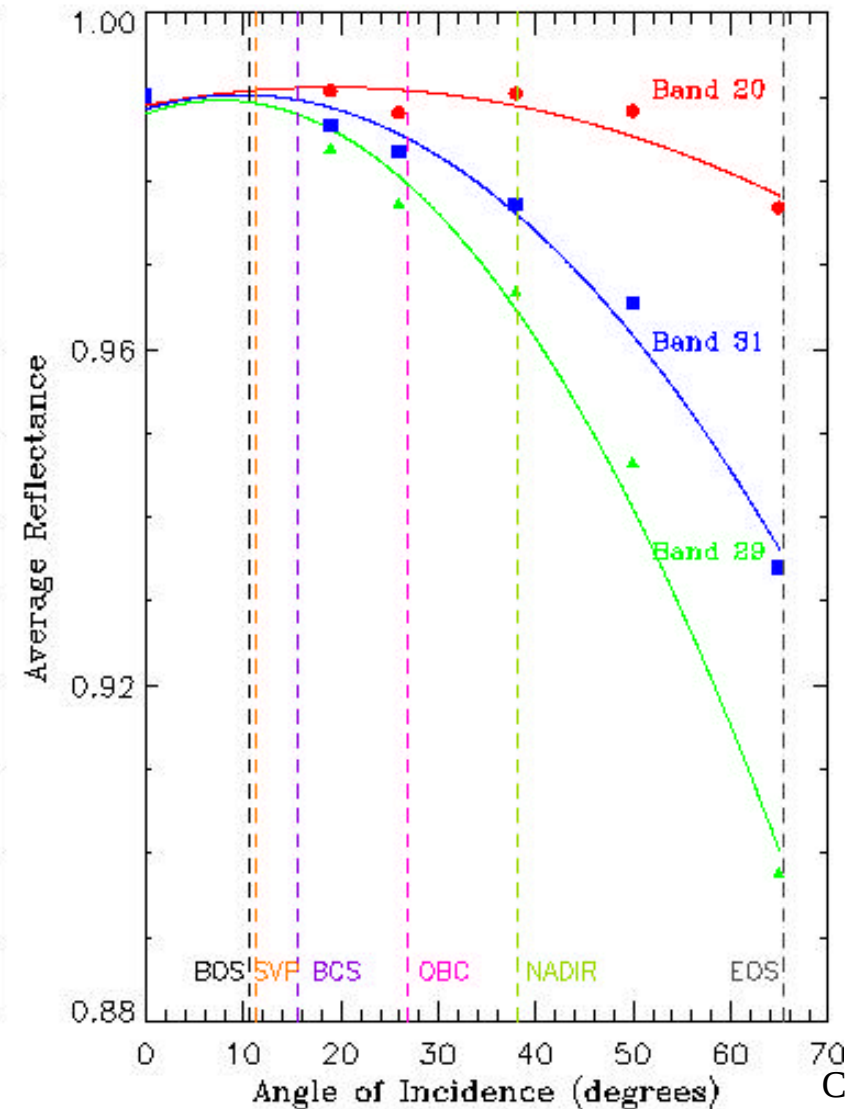
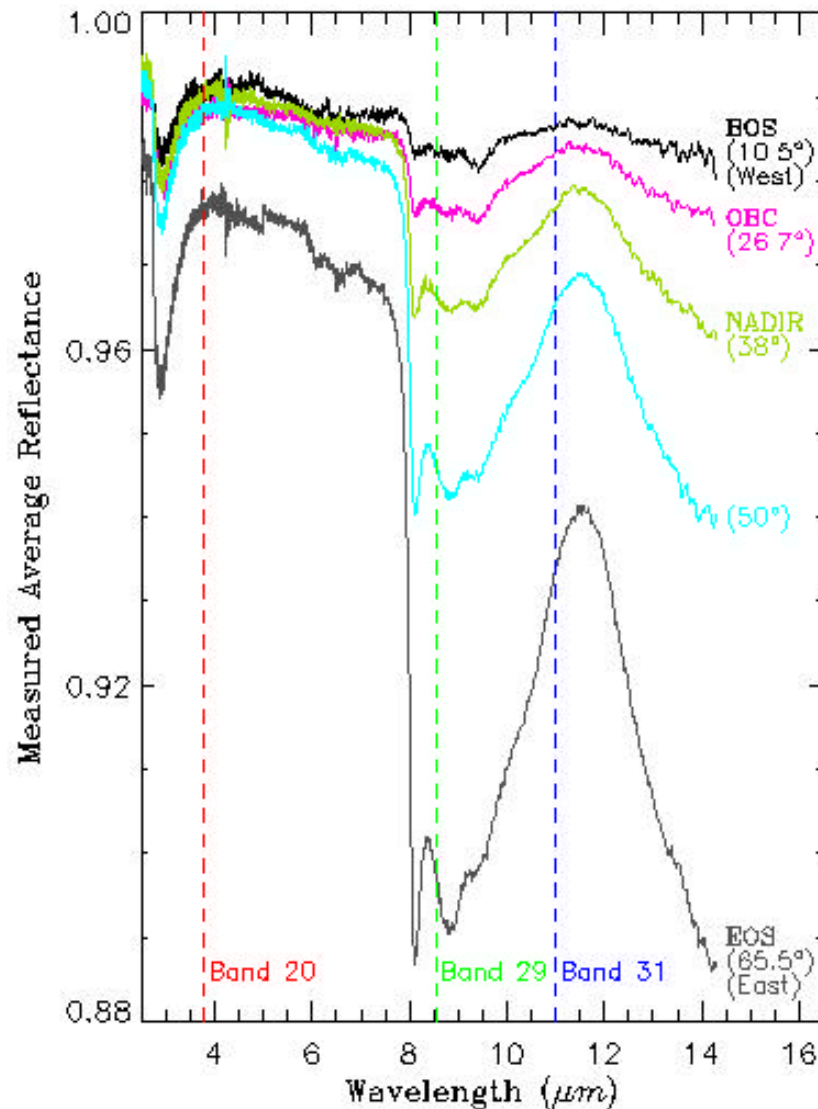
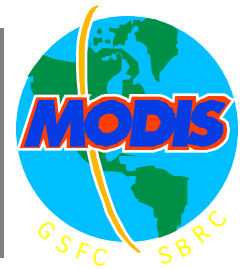


#### Notes:

1. Band single detector locations are shown in terms of 1 km IFOV offsets from Band 30, which is first band to see a scene. FPA optical center is at IFOV 15.
2. Ground scan direction of projected pattern is to the right (west to east). FPA's have false vertical separation to clarify individual FPA extent.
3. During one 1km sample, 250 m Bands 1 & 2 produce 16 samples/each and 500 m Bands 3-7 produce 4 samples/each.
4. The formatter assembles a science packet with bands aligned by the depicted offsets to view a common ground sample. This requires the collection of 30 frames of data before all bands have observed the same ground sample.
5. Science packet locations for TDI Bands 13, 13' and 14, 14' will be at first sample location of 13 and 14. High and low gain data is provided for Band 13 and Band 14. Bands 27-30 are sampled 4X, but single sample is sent to ground.

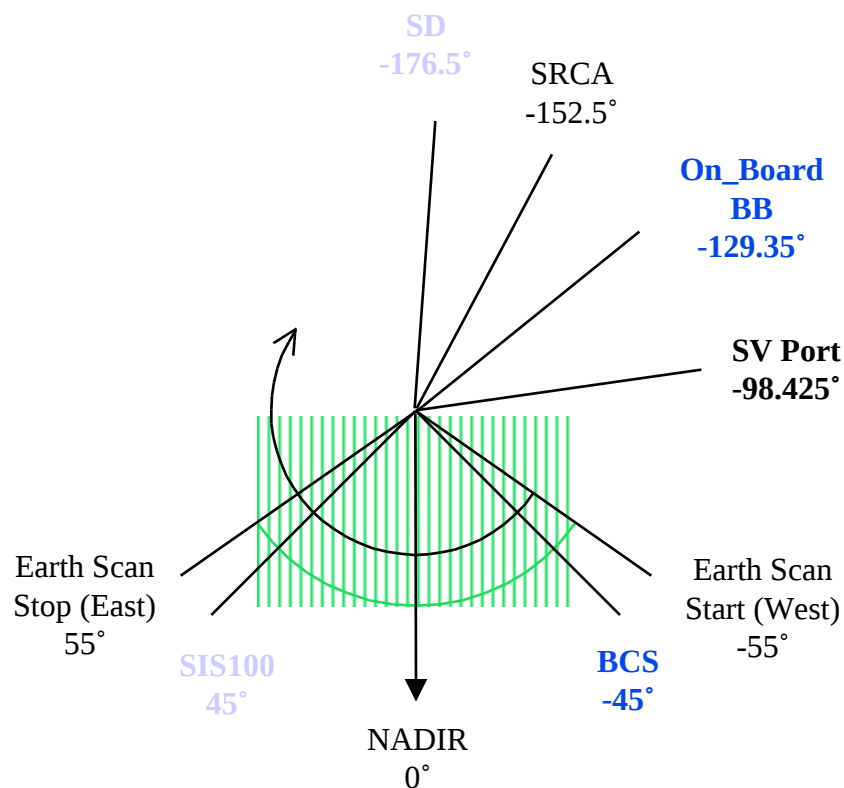
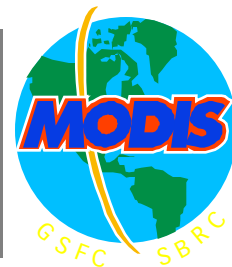


# Lincoln Laboratory Measured Scan Mirror Average Reflectance

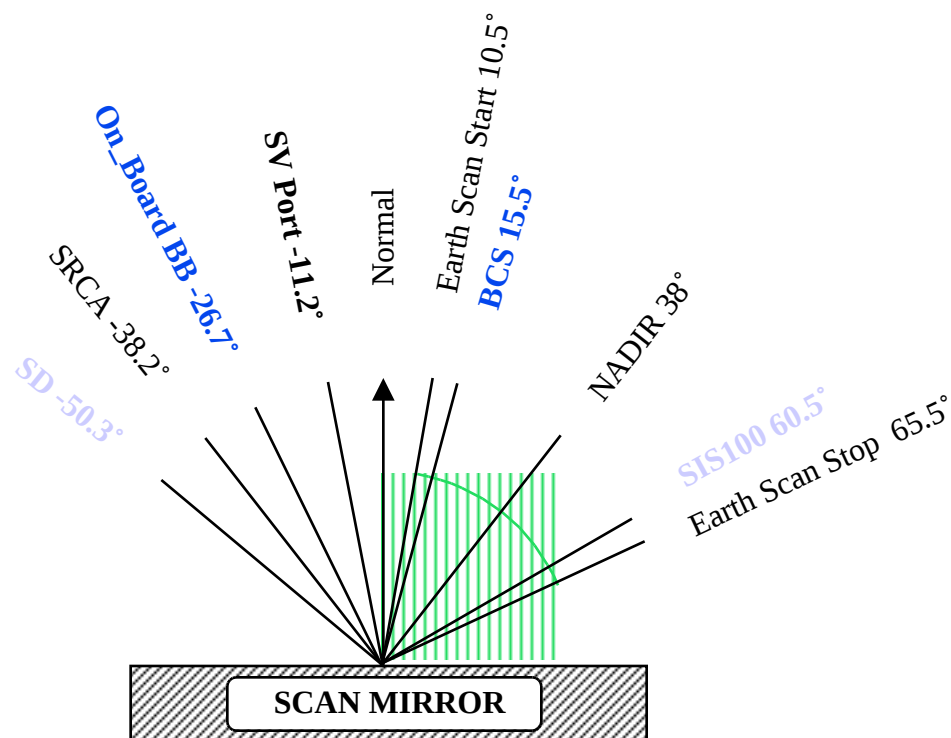




# Principal Scan Angles Mapped to Scan Mirror Angles of Incidence



**Principal Scan Angles**  
(Earth View: -55° to 55°)



**Angles of Incidence**  
(Earth View: 10.5° to 65.5°)

RADIOMETRIC

# MCST/Sun De-stripping Algorithm

- PHYSICAL ASSUMPTION: Image should be smooth across channels in-track, at each AOI, for each mirror side
- APPROACH: Uses Least Mean Square Method to determine improved linear gains for each channel, and improved RVS at each AOI, each MS, starts with Scaled Integer from product; best results when applied to scenes that appear to have little to no geophysical structure
- CONSTRAINT: Improved relative linear gains and RVS must be physically realistic, and best to be within measurement uncertainty
- Table on following page contains the perturbation values for measured RVS and linear gains

Table 1.  $f_{UD}(P)$  at  $F = 15, 955$ , and  $1325$  for band 8.

| D  | M=1    |        |        | M=2    |        |        |
|----|--------|--------|--------|--------|--------|--------|
|    | F=15   | F=955  | F=1325 | F=15   | F=955  | F=1325 |
| 1  | 1.0000 | 1.0000 | 1.0000 | 1.0059 | 1.0043 | 1.0058 |
| 2  | 1.0059 | 1.0043 | 1.0058 | 1.0071 | 1.0045 | 1.0067 |
| 3  | 1.0071 | 1.0045 | 1.0067 | 1.0083 | 1.0047 | 1.0074 |
| 4  | 1.0083 | 1.0047 | 1.0074 | 1.0087 | 1.0048 | 1.0074 |
| 5  | 1.0087 | 1.0048 | 1.0074 | 1.0119 | 1.0075 | 1.0099 |
| 6  | 1.0119 | 1.0075 | 1.0099 | 1.0105 | 1.0061 | 1.0078 |
| 7  | 1.0105 | 1.0061 | 1.0078 | 1.0128 | 1.0080 | 1.0094 |
| 8  | 1.0128 | 1.0080 | 1.0094 | 1.0131 | 1.0084 | 1.0093 |
| 9  | 1.0131 | 1.0084 | 1.0093 | 1.0138 | 1.0095 | 1.0097 |
| 10 | 1.0138 | 1.0095 | 1.0097 | 1.0271 | 1.0203 | 1.0132 |

Table 2. Coefficients  $c_{UD}^0$ ,  $c_{UD}^1$ , and  $c_{UD}^2$  for band 8.

| D  | M=1     |          |         | M=2     |          |          |
|----|---------|----------|---------|---------|----------|----------|
|    | c0      | c1       | c2      | c0      | c1       | c2       |
| 1  | 1.00000 | 0.00000  | 0.00000 | 1.00602 | -0.00815 | 0.00810  |
| 2  | 1.00602 | -0.00815 | 0.00810 | 1.00728 | -0.01259 | 0.01226  |
| 3  | 1.00728 | -0.01259 | 0.01226 | 1.00854 | -0.01649 | 0.01566  |
| 4  | 1.00854 | -0.01649 | 0.01566 | 1.00895 | -0.01698 | 0.01573  |
| 5  | 1.00895 | -0.01698 | 0.01573 | 1.01216 | -0.01769 | 0.01572  |
| 6  | 1.01216 | -0.01769 | 0.01572 | 1.01073 | -0.01578 | 0.01307  |
| 7  | 1.01073 | -0.01578 | 0.01307 | 1.01303 | -0.01597 | 0.01253  |
| 8  | 1.01303 | -0.01597 | 0.01253 | 1.01331 | -0.01435 | 0.01048  |
| 9  | 1.01331 | -0.01435 | 0.01048 | 1.01397 | -0.01143 | 0.00722  |
| 10 | 1.01397 | -0.01143 | 0.00722 | 1.02707 | 0.00221  | -0.01674 |

Table 3.  $\mathcal{M}_{UD}^{(w)}$  for band 8.

| D  | M=1     | M=2     |
|----|---------|---------|
| 1  | 0.98589 | 1.00552 |
| 2  | 0.99019 | 1.00801 |
| 3  | 0.99041 | 1.00777 |
| 4  | 0.99062 | 1.00908 |
| 5  | 0.99072 | 1.00902 |
| 6  | 0.99337 | 1.01026 |
| 7  | 0.99195 | 1.00900 |
| 8  | 0.99381 | 1.00823 |
| 9  | 0.99418 | 1.00836 |
| 10 | 0.99523 | 1.00838 |

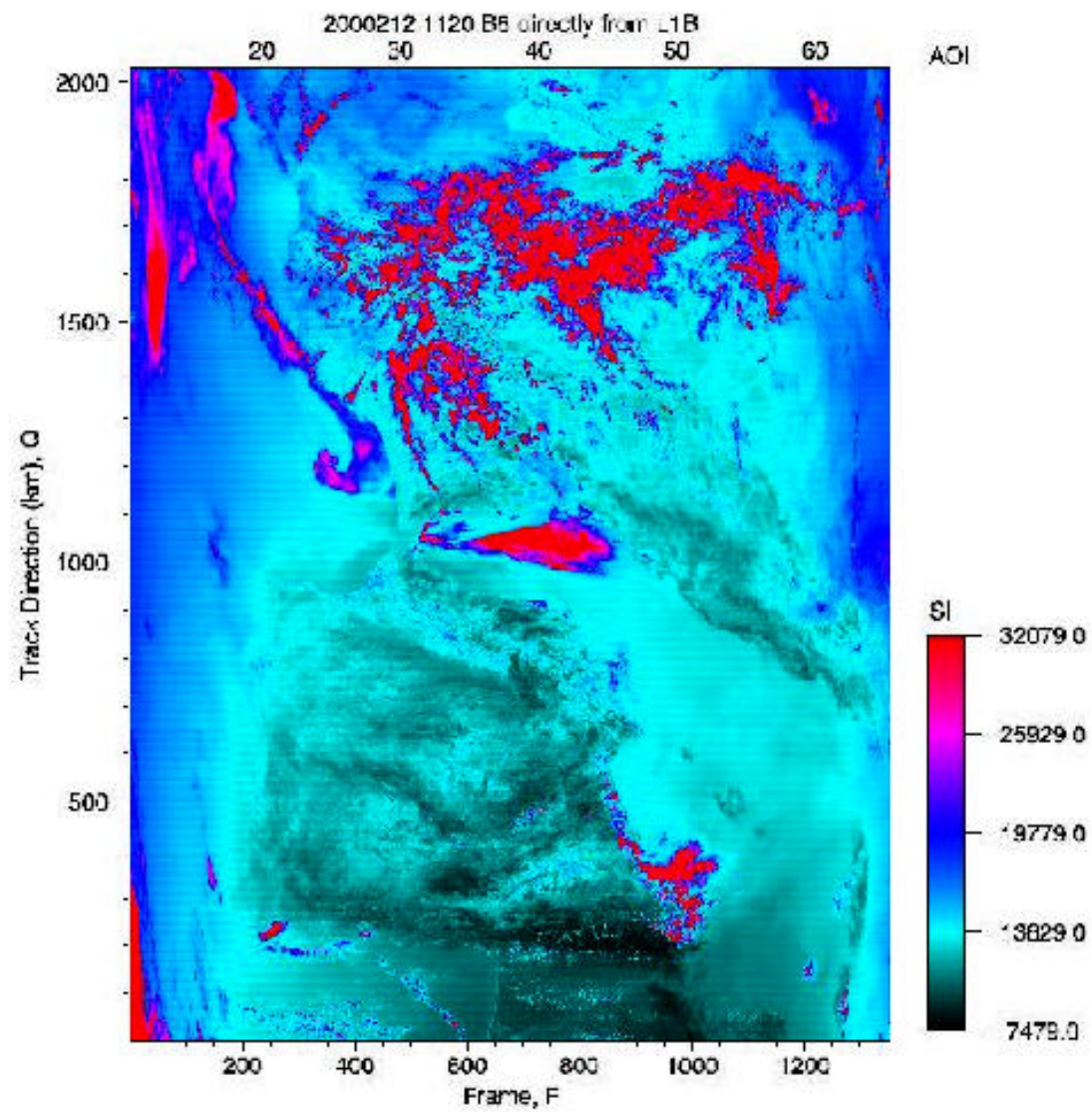


Figure1. Image of original SI from L1B for Band 8.

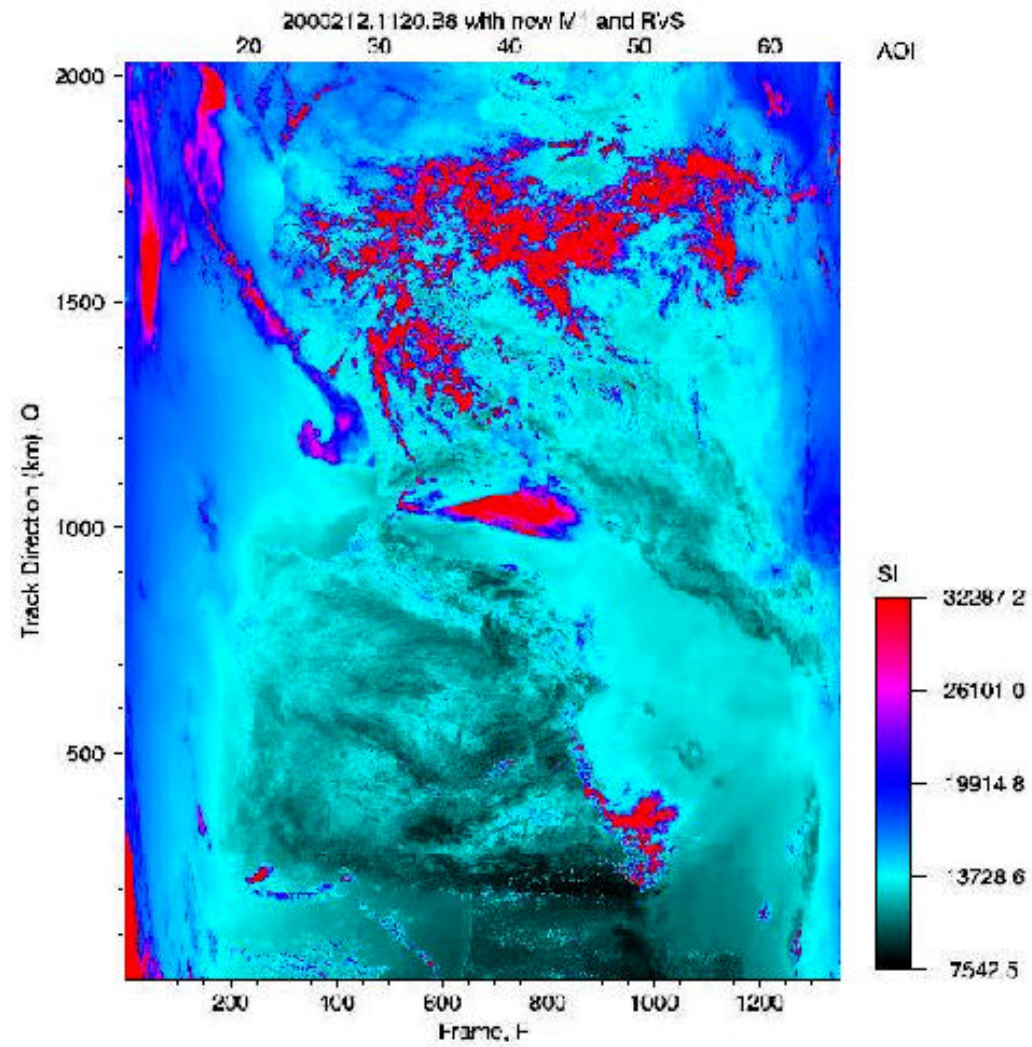


Figure 2. Image of SI reproduced with new  $M^1$  and  $RVS$  from the same data presented in Figure 1.

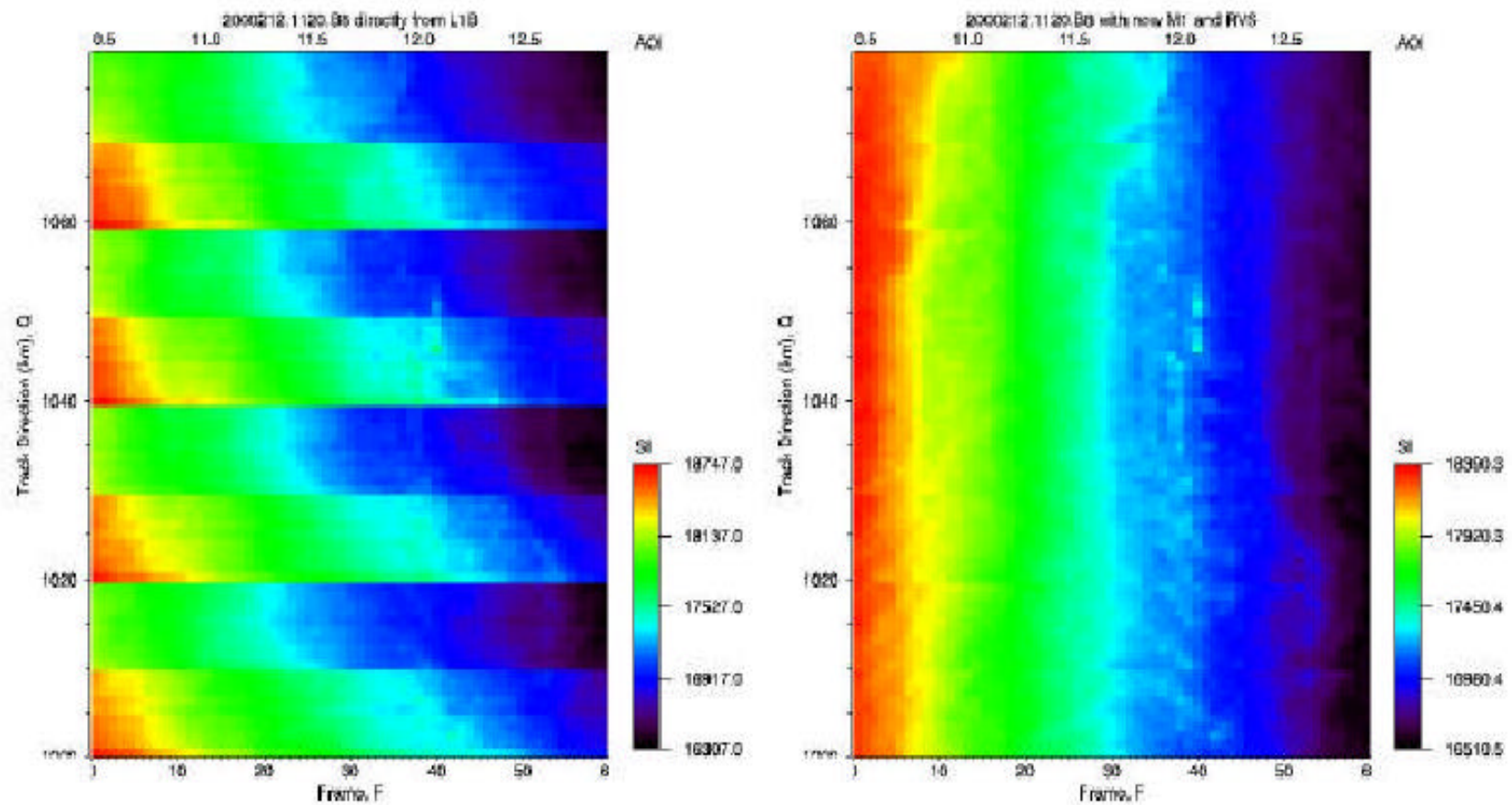


Figure 3. Comparison of the original SI directly from L1B and that reproduced with the new *MI* and *RVS*. The images are taken from left side of Figure 1 and Figure 2.

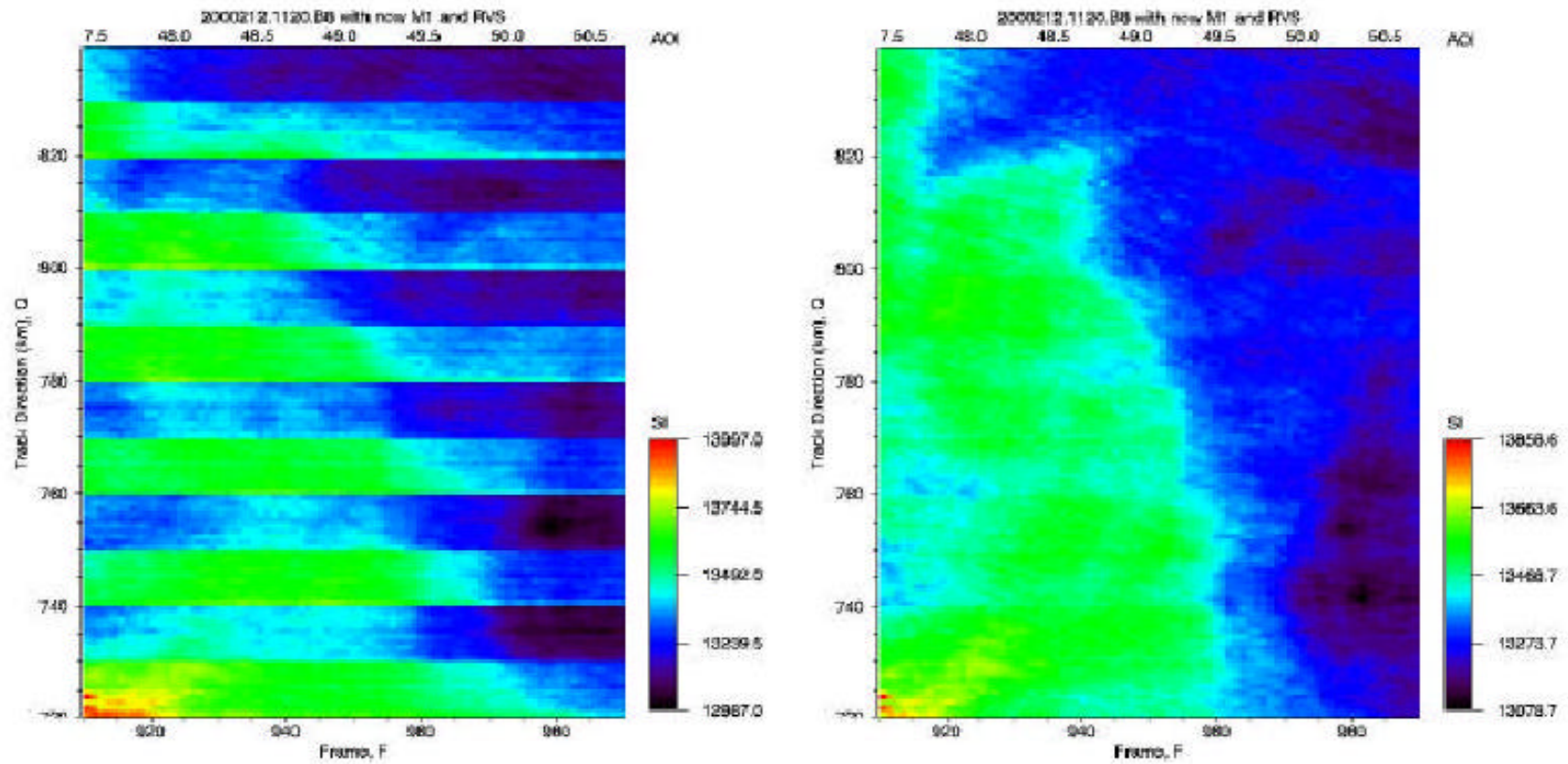


Figure 4. Comparison of the original SI directly from L1B and that reproduced with the new *MI* and *RVS*. The images are taken from middle of Figure 1 and Figure 2.

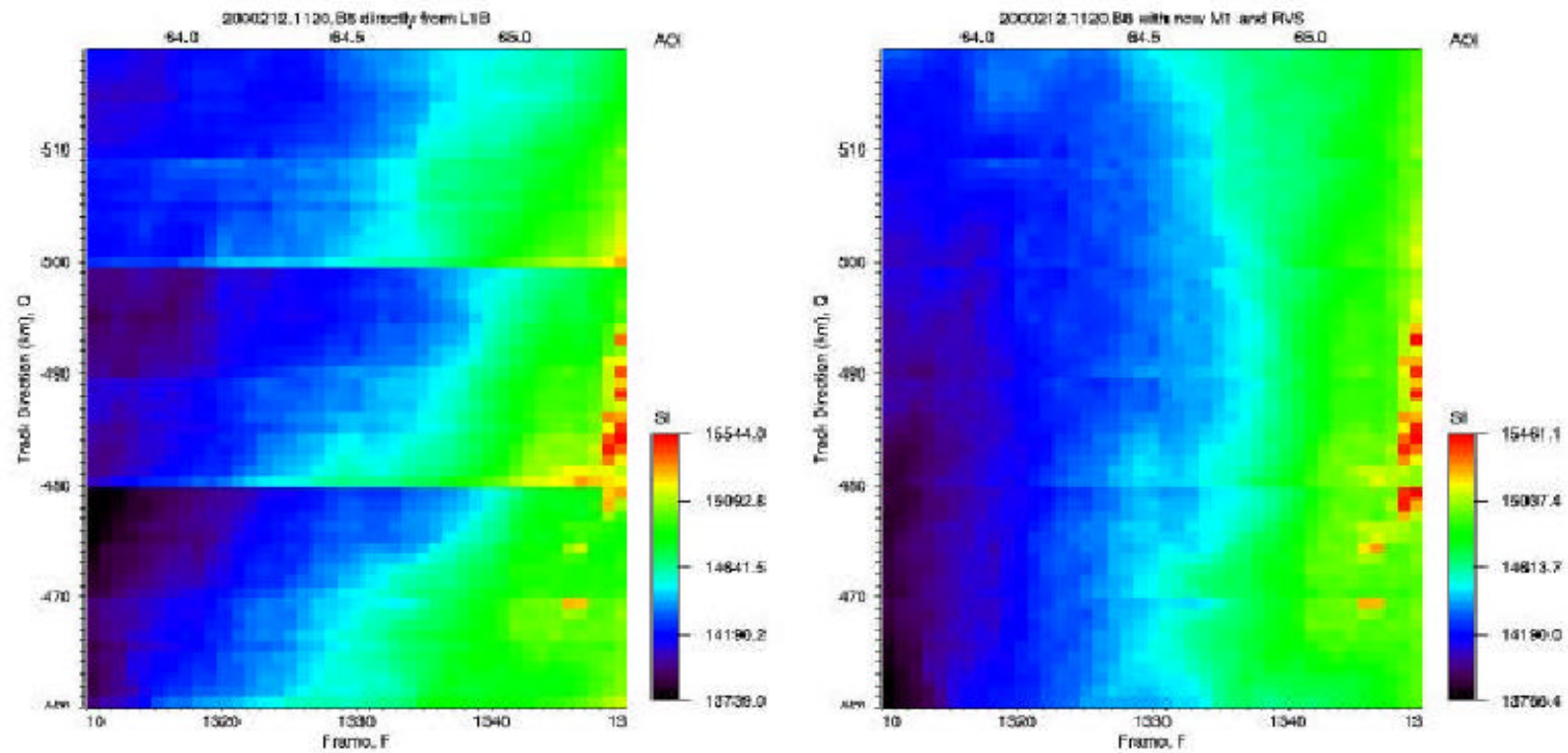


Figure 5. Comparison of the original SI directly from LIB and that reproduced with the new *MI* and *RVS*. The images are taken from right side of Figure 1 and Figure 2.

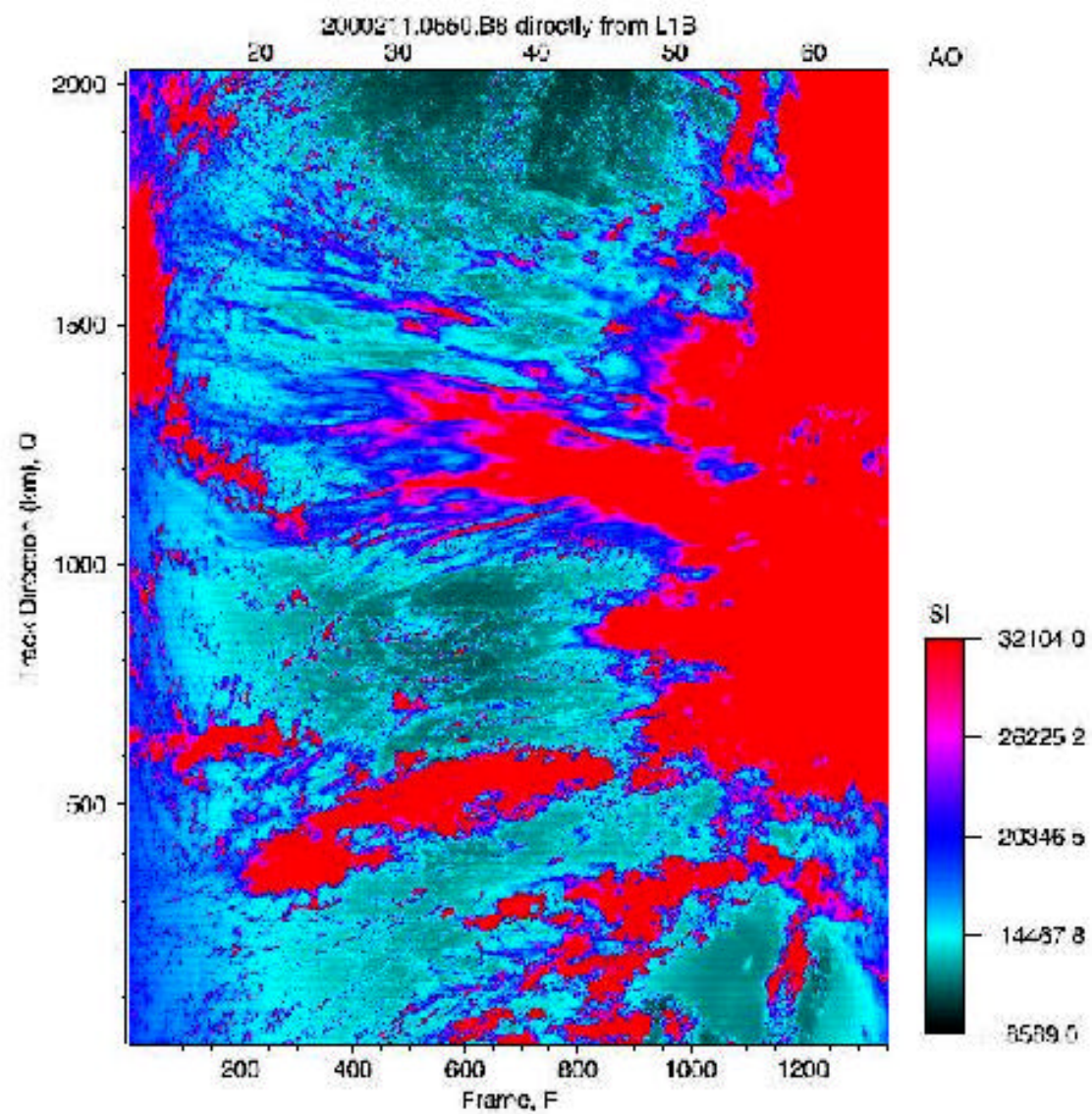


Figure 6. Image of original SI from L1B for Band 8.

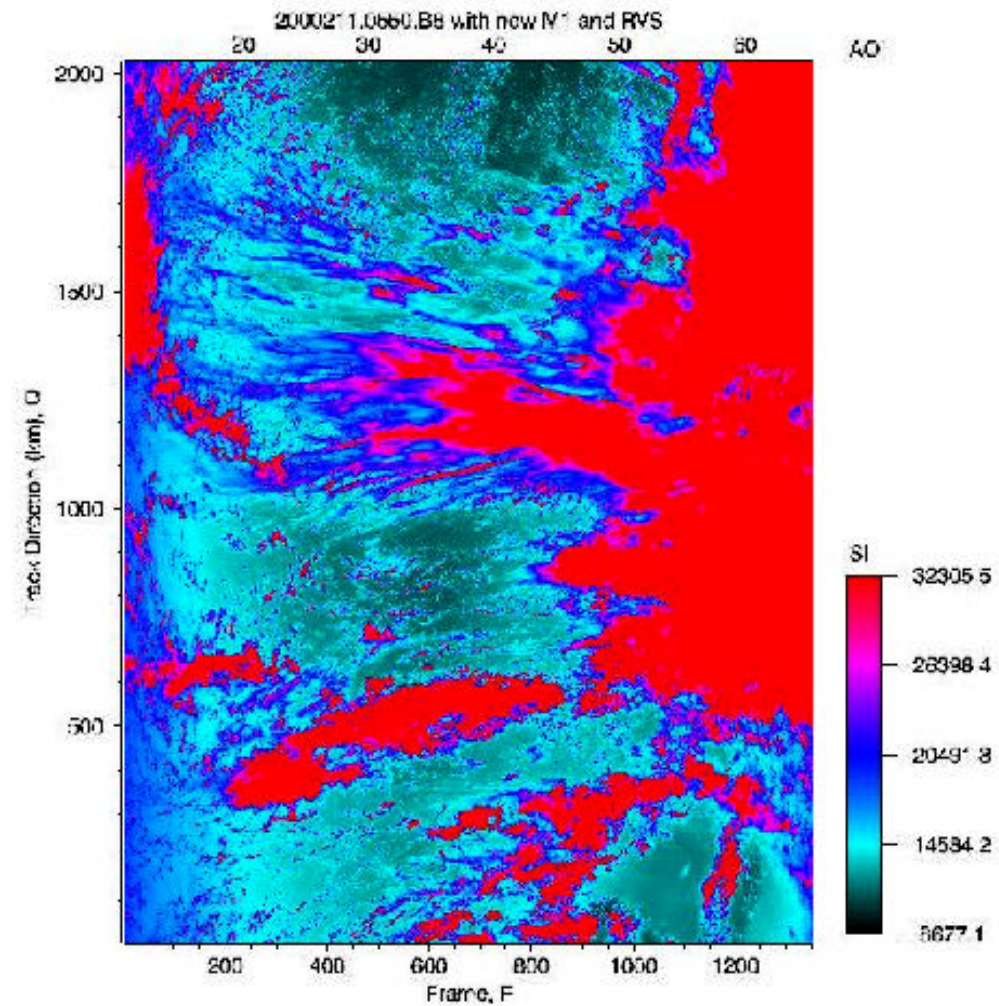


Figure 7. Image of SI reproduced with new *MI* and *RVS* from the data plotted in Figure 6.

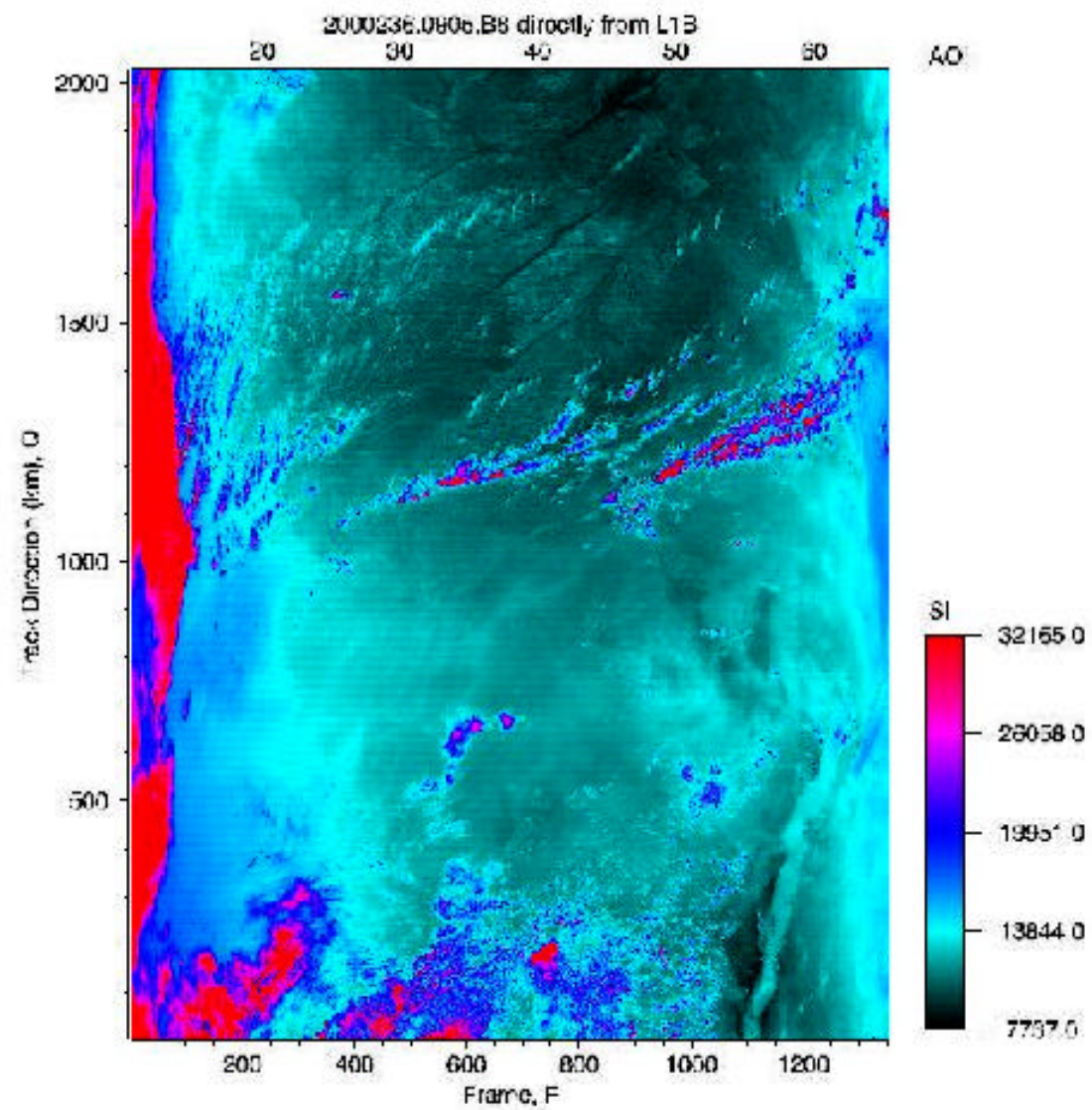


Figure 8. Image of original SI from LIB for Band 8.

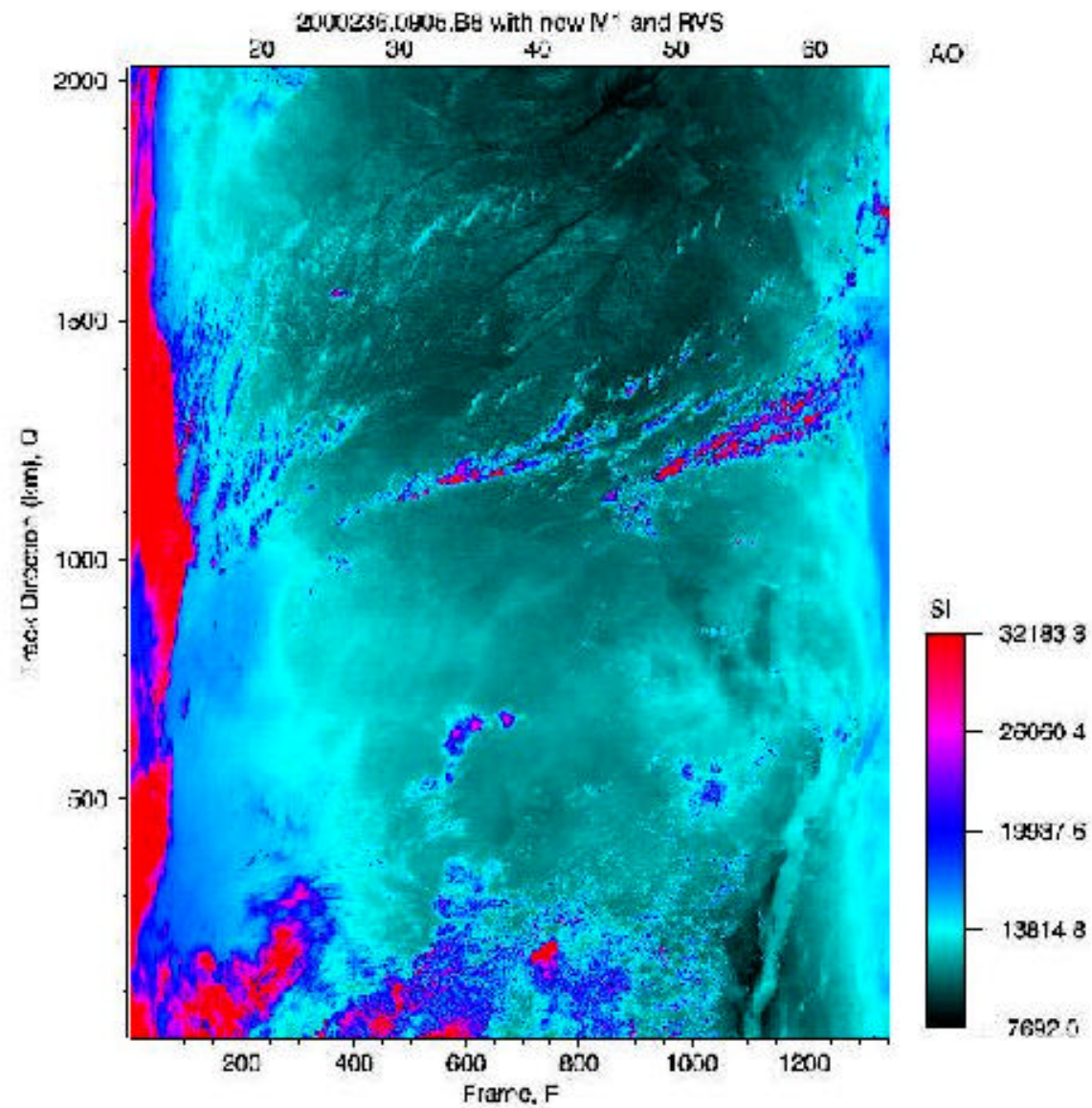


Figure 9. Image of SI reproduced with the new *MI* and *RVS* from the data presented in Figure 8.

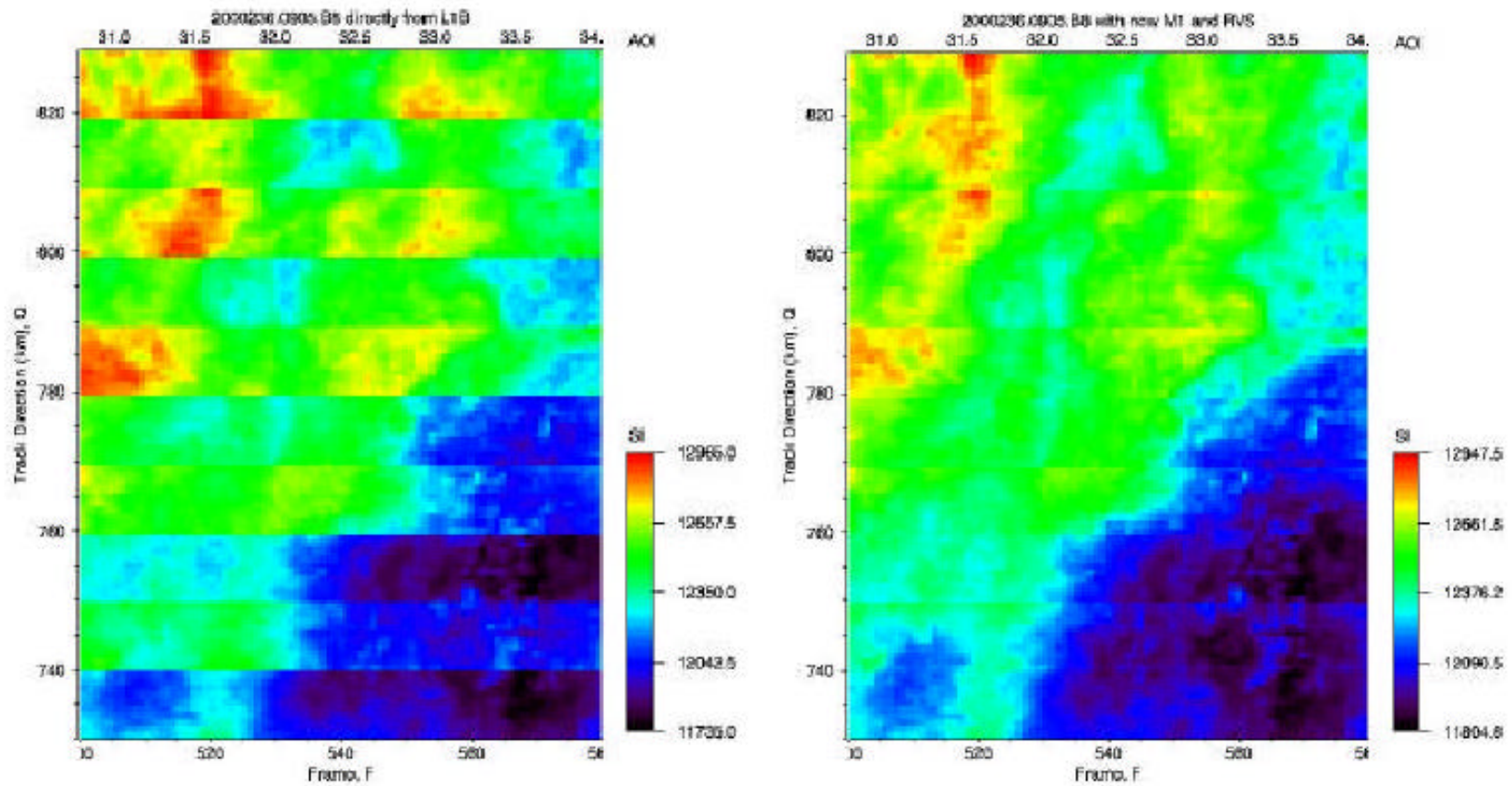


Figure 10. Comparison of the original SI directly from L1B and that reproduced with the new *MI* and *RVS*. The images are taken from right side of Figure 8 and Figure 9.

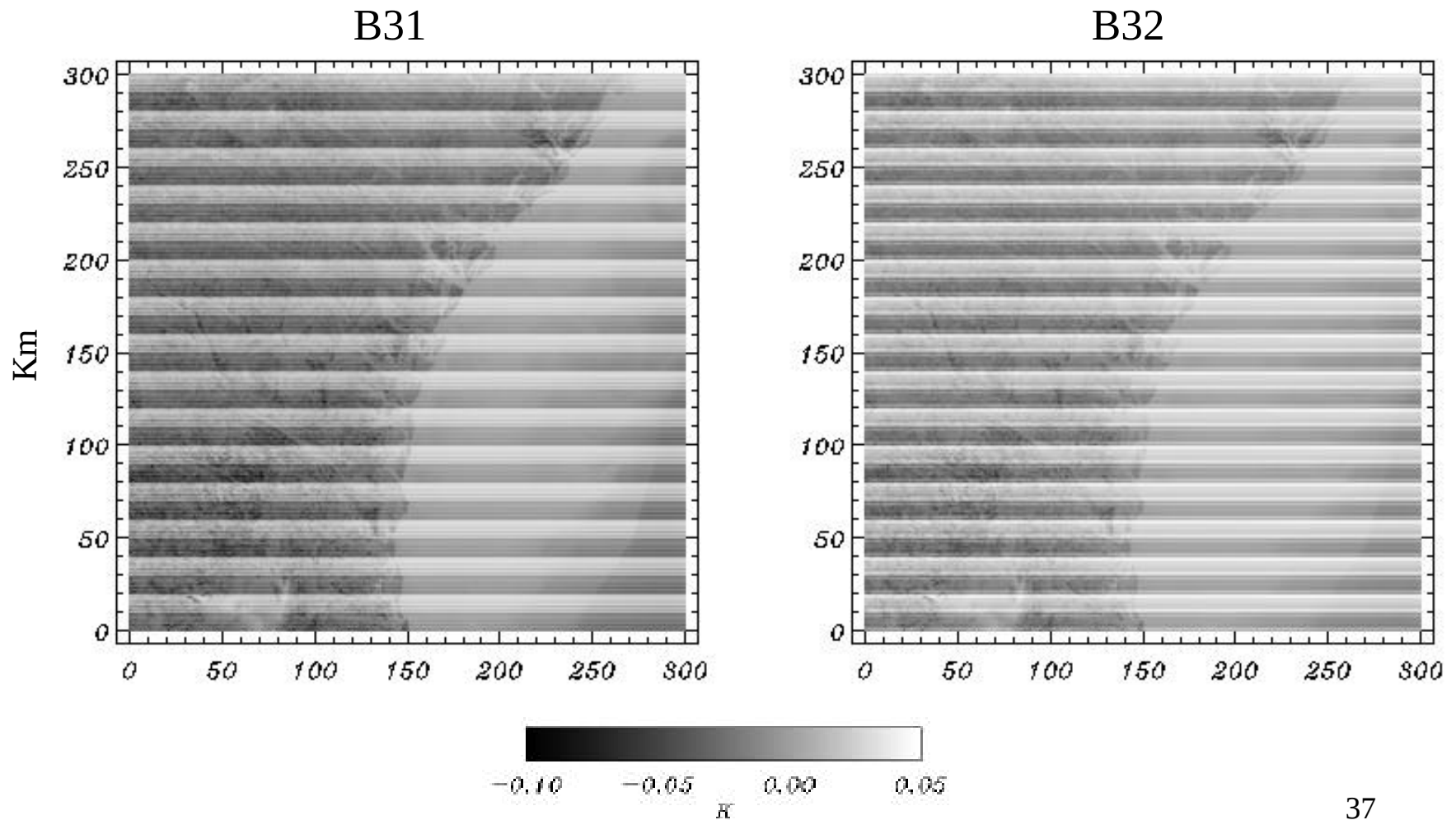
## Conclusions: MCST/Sun Algorithm

- Shows promise, but algorithm still in experimental stage
- Applied only to Band 8 data so far
- Does much better at beginning of scan than end of scan
  - Such a result may be expected from geophysical basis
- Do not fully understand yet details of how to implement this operationally in L1B, but would need to navigate scenes by hand to determine improved RVS and linear gains
- MCST and users (Miami!) must work closely together to sort out this approach as well as others; hopefully one will be good enough to apply operationally in L1B product

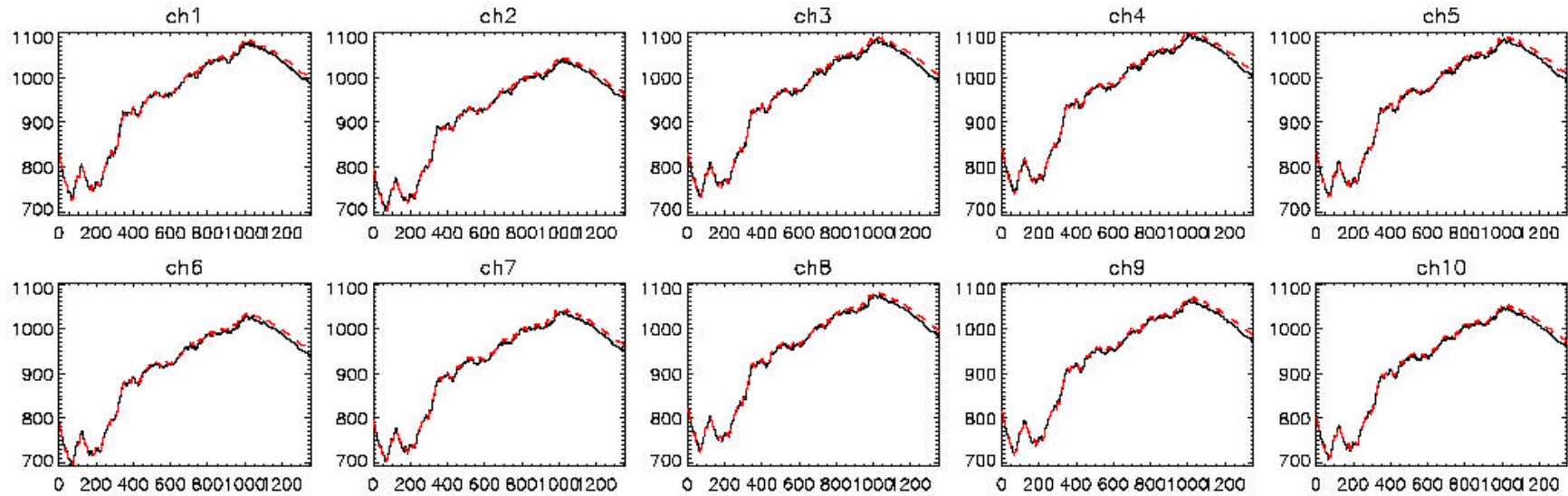
# TOA B31 & B32 Temperature Retrieval Difference Between a0, a2 (Pre-launch) and b0, b2 (On-orbit)

$$DT = T_{EV}(b_0, b_2) - T_{EV}(a_0, a_2)$$

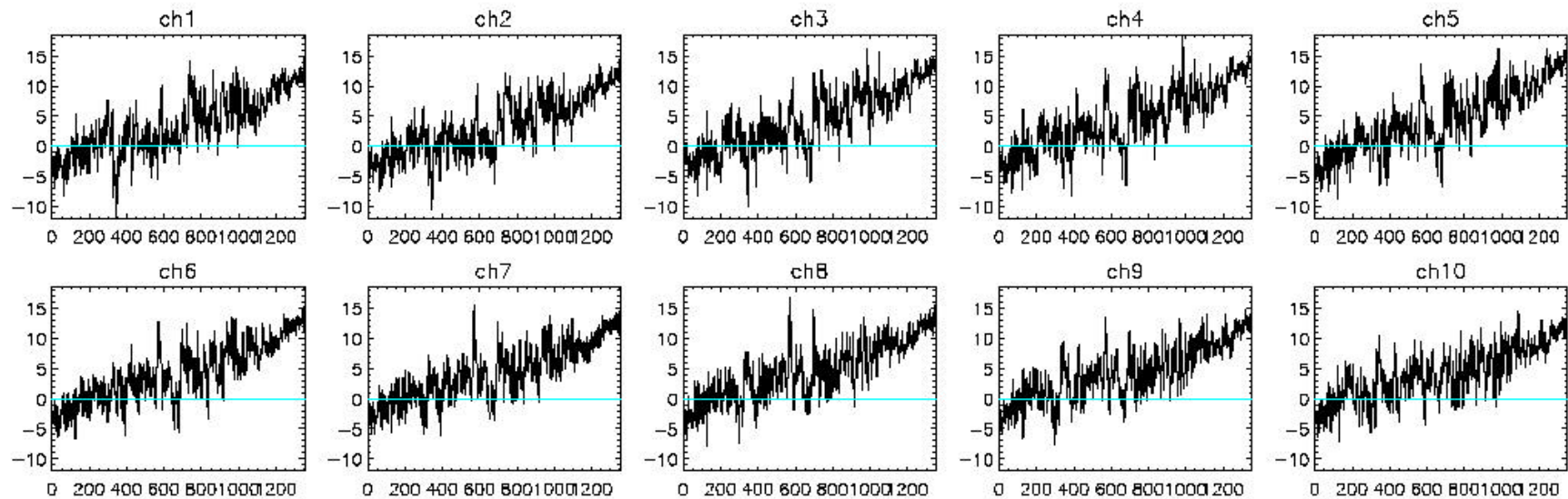
Images begin with mirror side 1



# 2000088.0350 Band 35 dn



## 2000088.0350 Band 35 Mirror Side dn Difference (ms2-ms1)

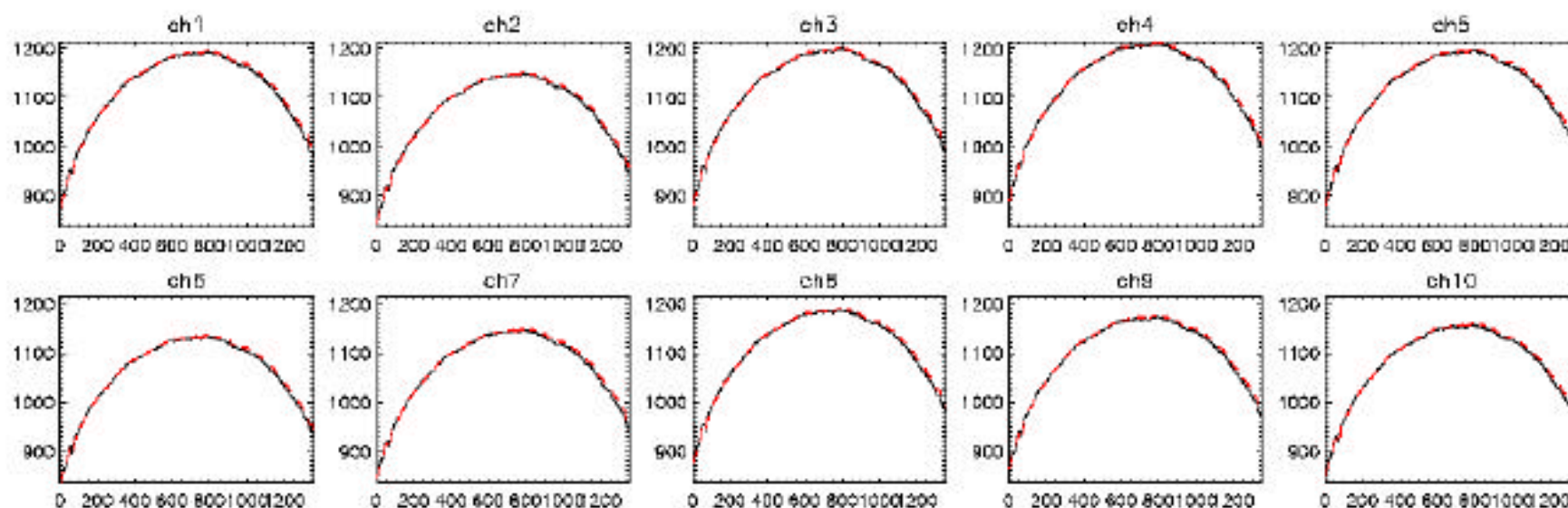


Side 1 —

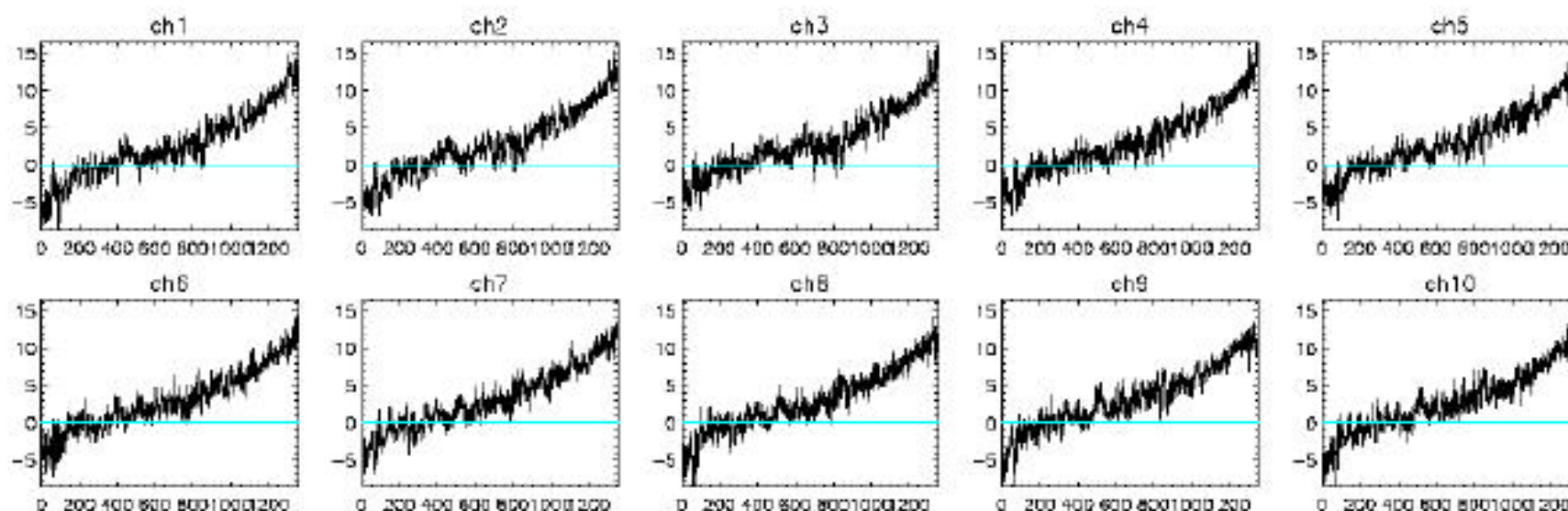
Side 2 ---

dn\_ev difference

# 2000103.1830 Band 35 dn



## 2000103.1830 Band 35 Mirror Side dn Difference ( $ms2 - ms1$ )

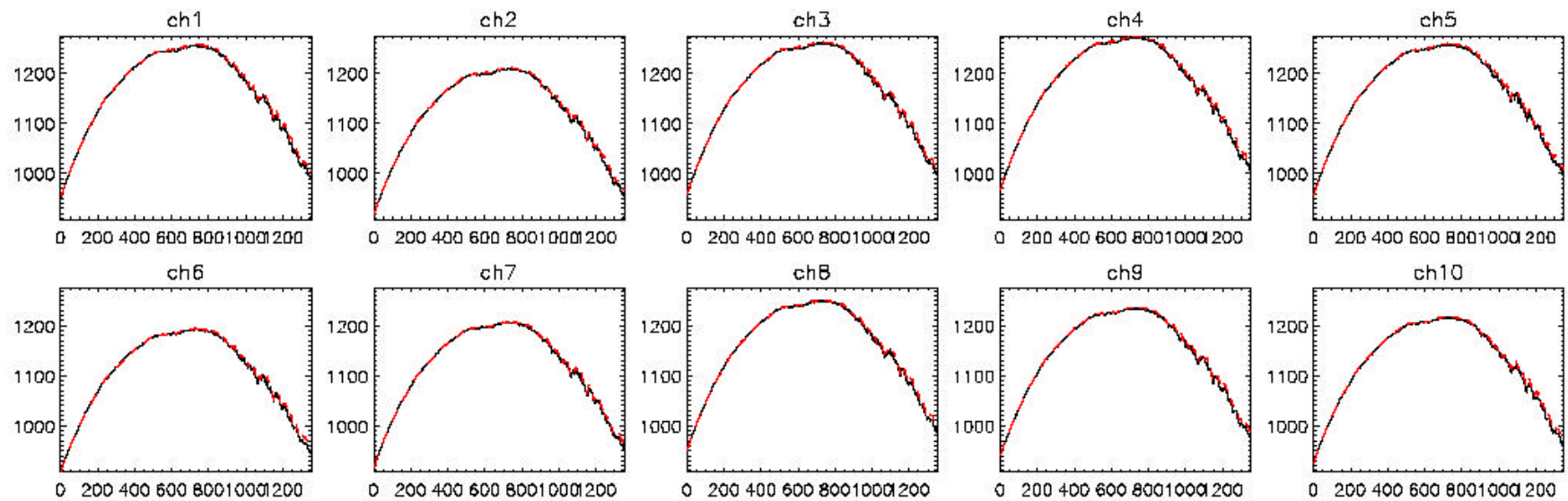


Side 1 —

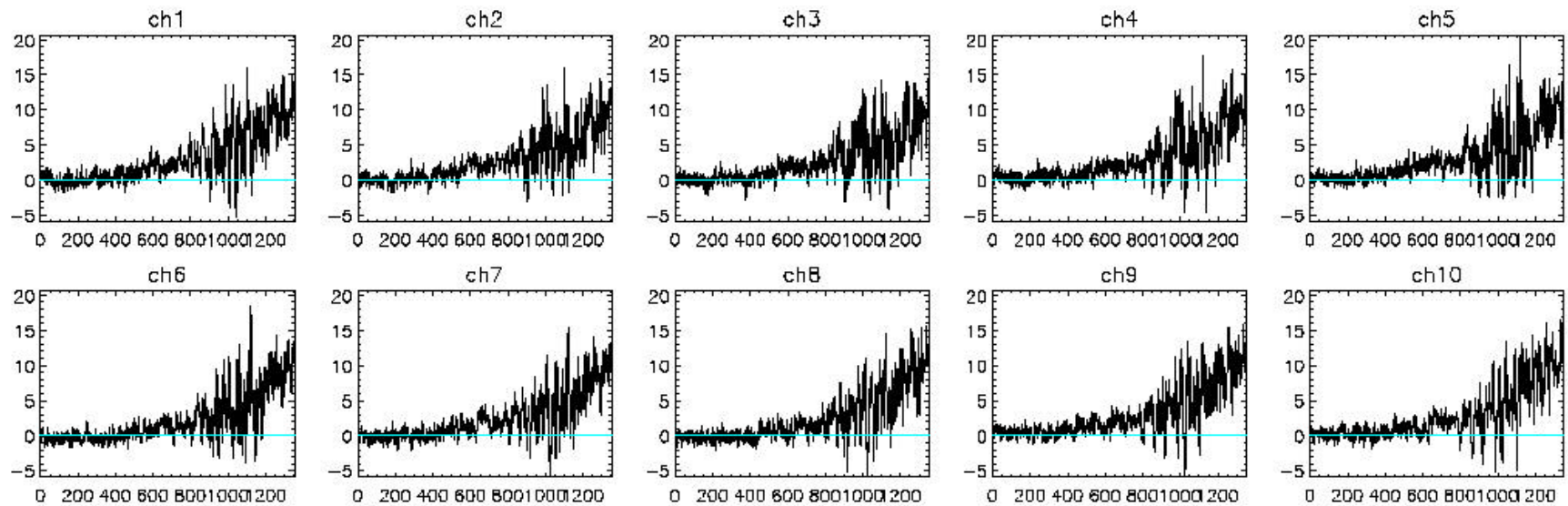
Side 2 ---

dnlev difference

# 2000158.1835 Band 35 dn



## 2000158.1835 Band 35 Mirror Side dn Difference (ms2-ms1)

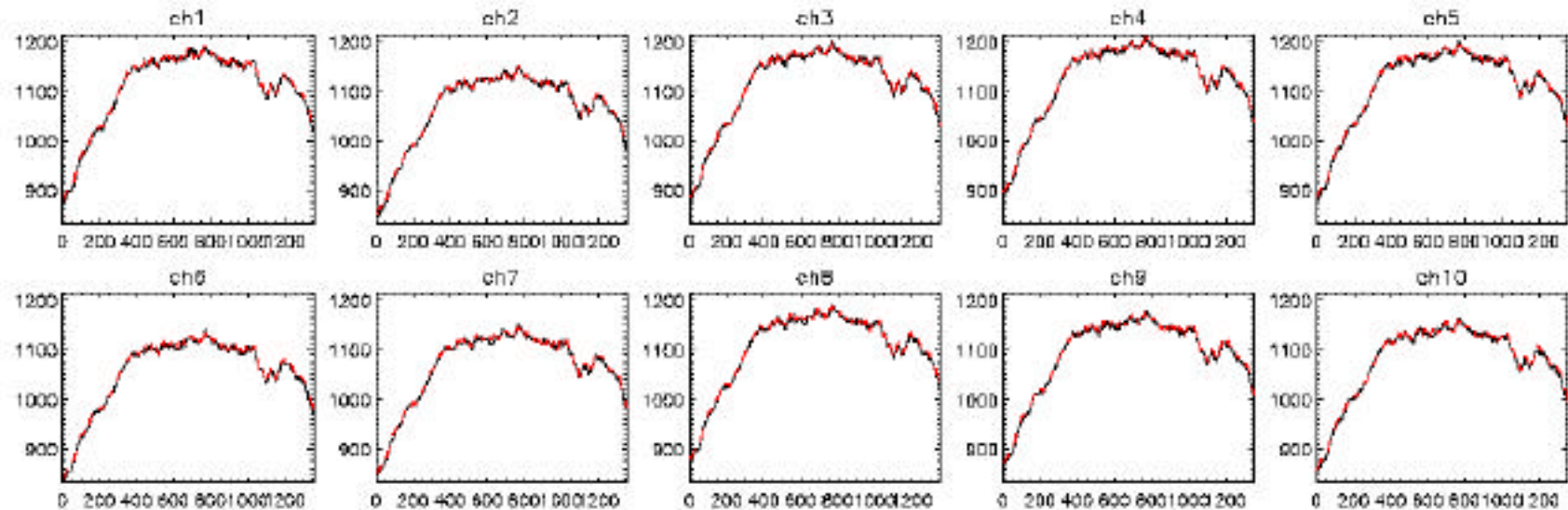


Side 1 —

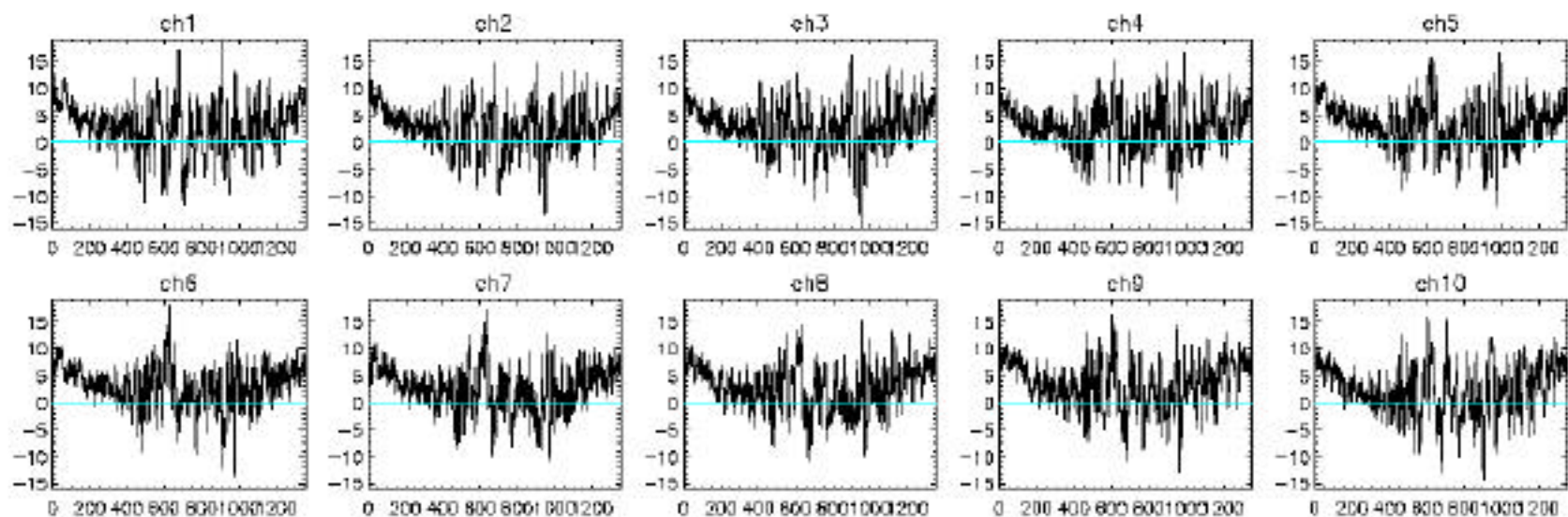
Side 2 ---

dnlev difference

# 2000247.1825 Band 35 dn



## 2000247.1825 Band 35 Mirror Side dn Difference (ms2-ms1)

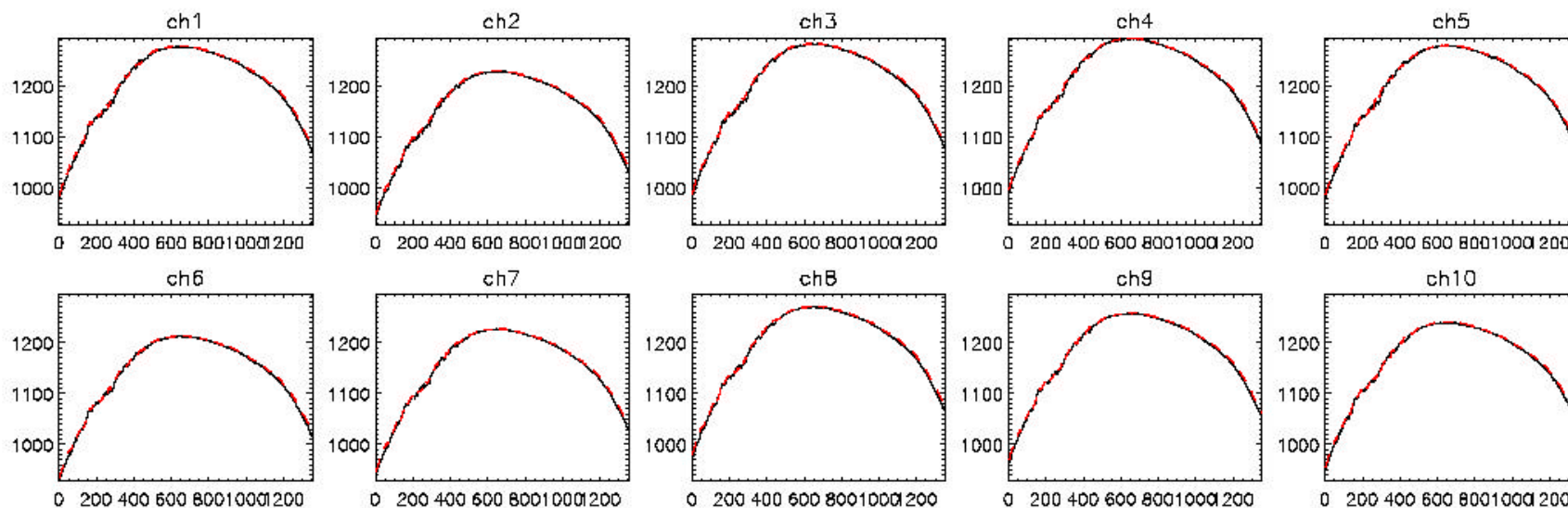


Side 1 —

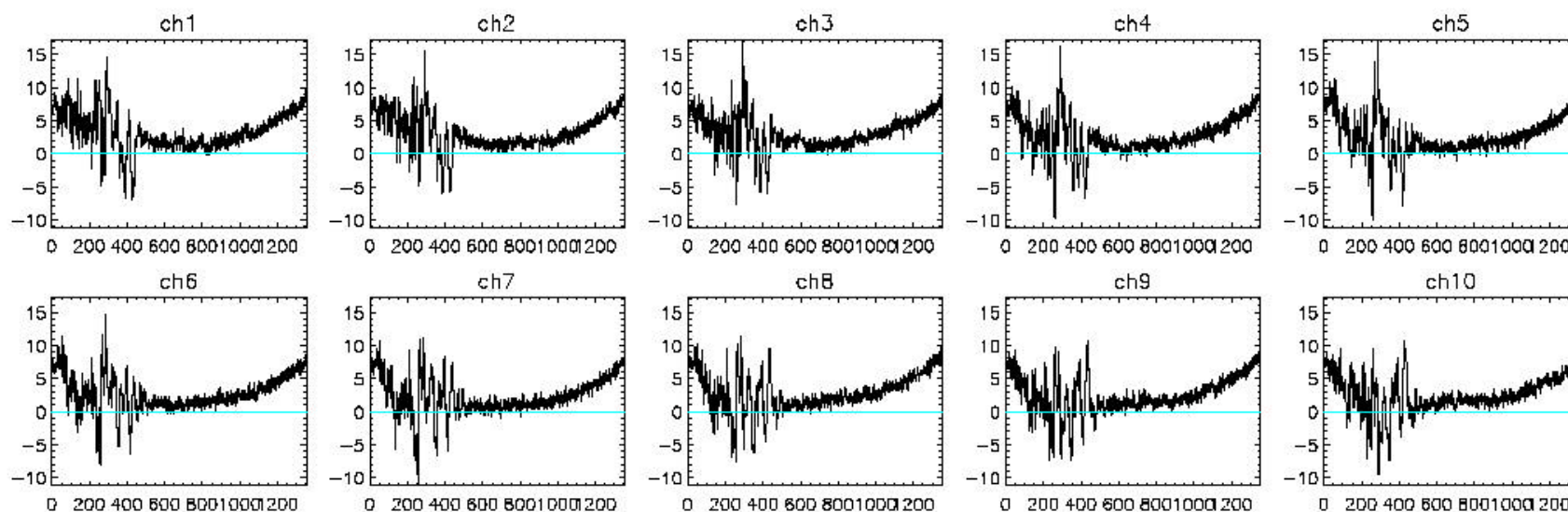
Side 2 ---

dn Lev difference

# 2000278.1110 Band 35 dn



## 2000278.1110 Band 35 Mirror Side dn Difference (ms2-ms1)



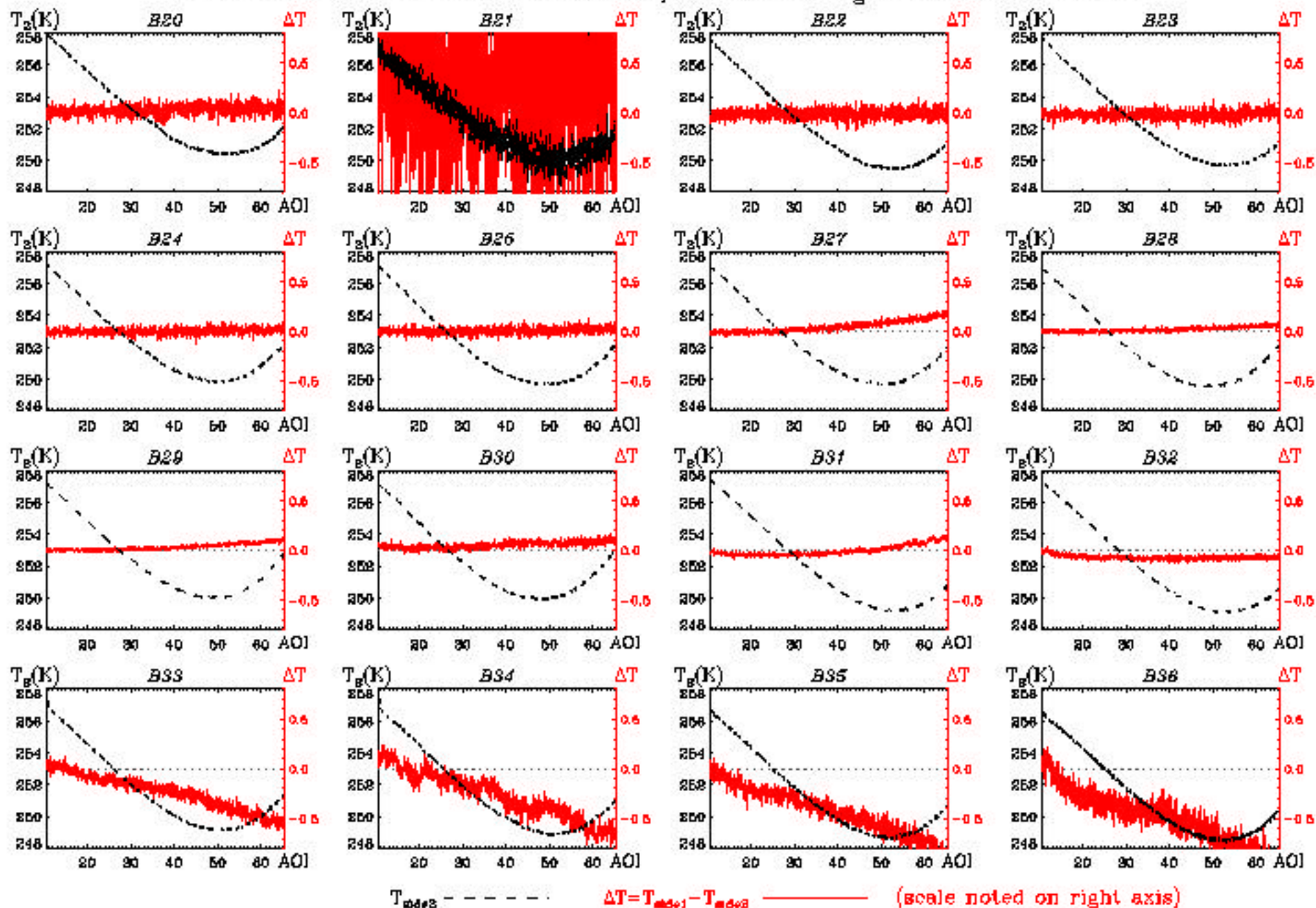
Side 1 —

Side 2 ---

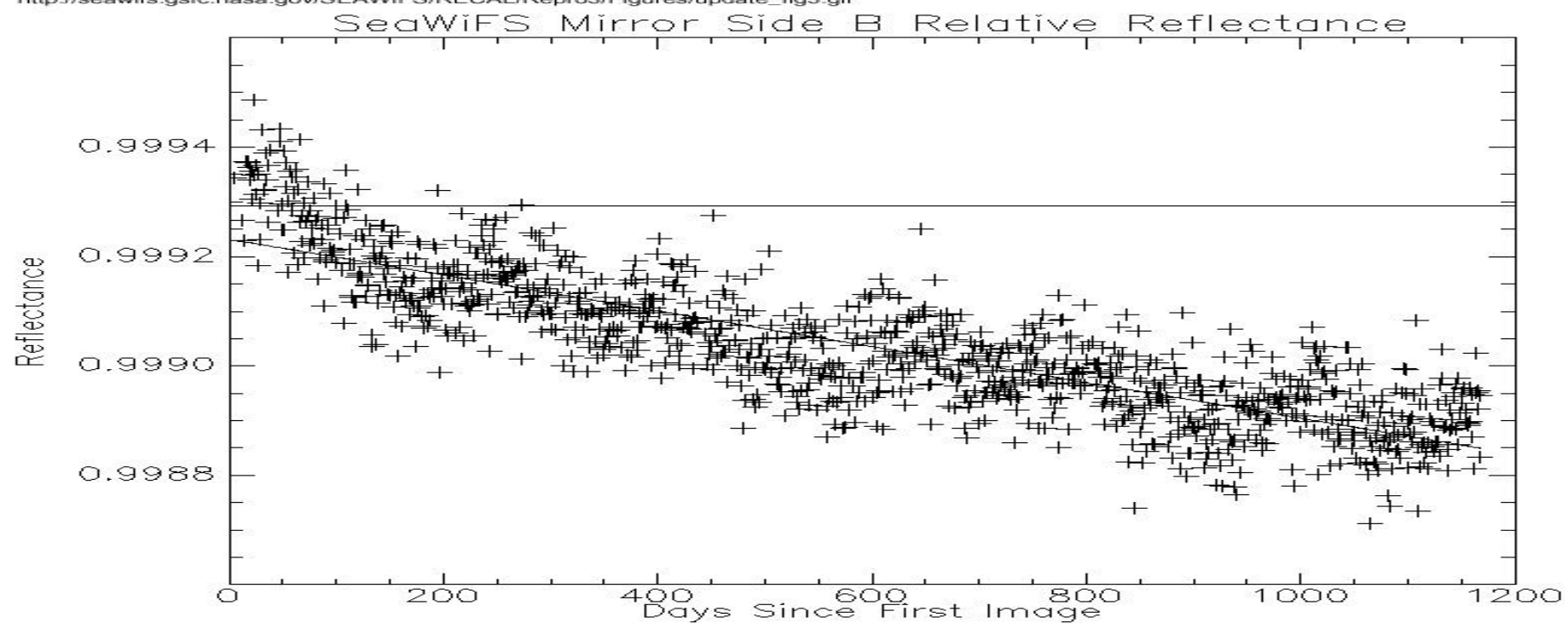
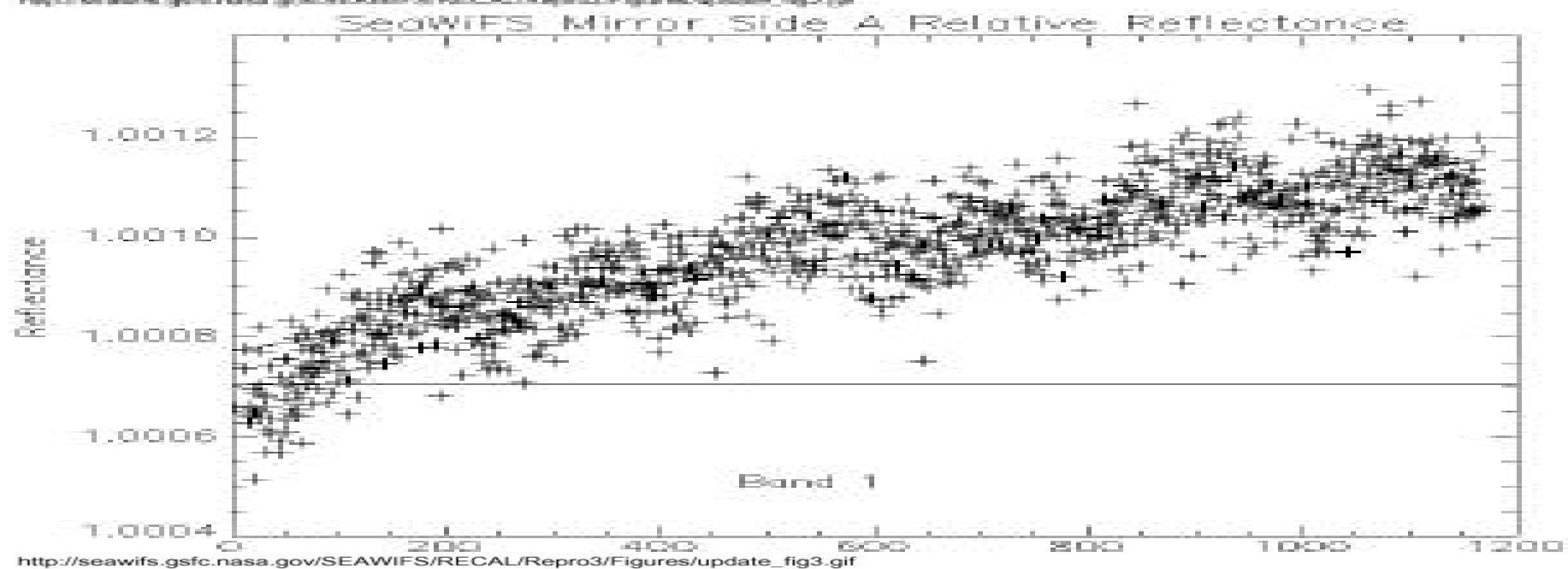
dn\_ev difference

# Brightness Temperature Retrieval vs AOI inside the EV Nadir Door

Data shown for Ch5 on L1A:2000116.1215; 30 scans average for each mirror side

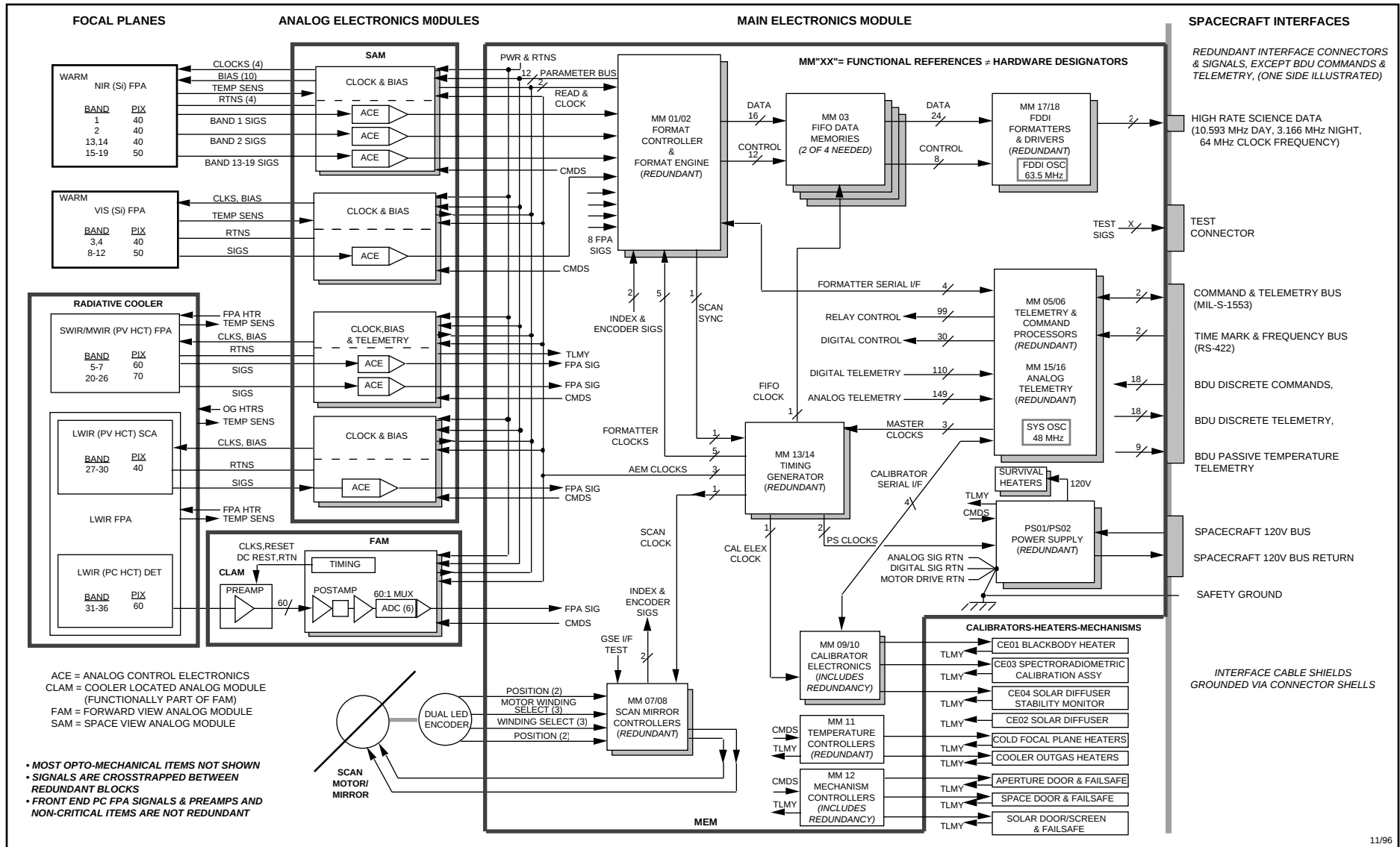






## Radiometric Section 2

# MODIS Electronic Block Diagram





# ADC List



| ADC # | ADC Name   | Bands     |
|-------|------------|-----------|
| 1     | VIS (p)    | 3,4, 8-12 |
| 2     | VIS (r)    |           |
| 3     | NIR1 (p)   | 1         |
| 4     | NIR1 (r)   |           |
| 5     | NIR2 (p)   | 2         |
| 6     | NIR2 (r)   |           |
| 7     | NIR3 (p)   | 13-19     |
| 8     | NIR3 (r)   |           |
| 9     | SMIR5 (p)  | 5,6,7     |
| 10    | SMIR5 (r)  |           |
| 11    | SMIR20 (p) | 20-26     |
| 12    | SMIR20 (r) |           |
| 13    | LWIR (p)   | 27-30     |
| 14    | LWIR (r)   |           |



**PV Bands ADC Conversion Time: 800 ns**

| ADC # | ADC Name  | Bands |
|-------|-----------|-------|
| 15    | PCB31 (p) | 31    |
| 16    | PCB31 (r) |       |
| 17    | PCB32 (p) | 32    |
| 18    | PCB32 (r) |       |
| 19    | PCB33 (p) | 33    |
| 20    | PCB33 (r) |       |
| 21    | PCB34 (p) | 34    |
| 22    | PCB34 (r) |       |
| 23    | PCB35 (p) | 35    |
| 24    | PCB35 (r) |       |
| 25    | PCB36 (p) | 36    |
| 26    | PCB36 (r) |       |



**PC Bands ADC Conversion Time: 15  $\mu$ s**

**p / r: primary / redundant**



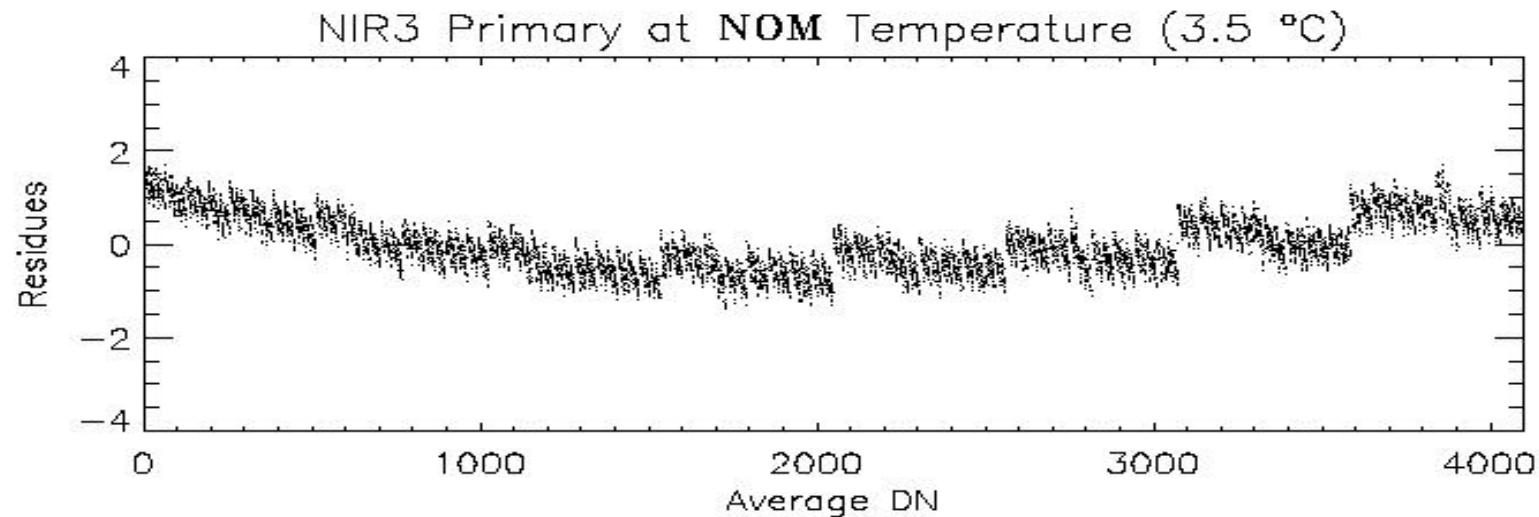
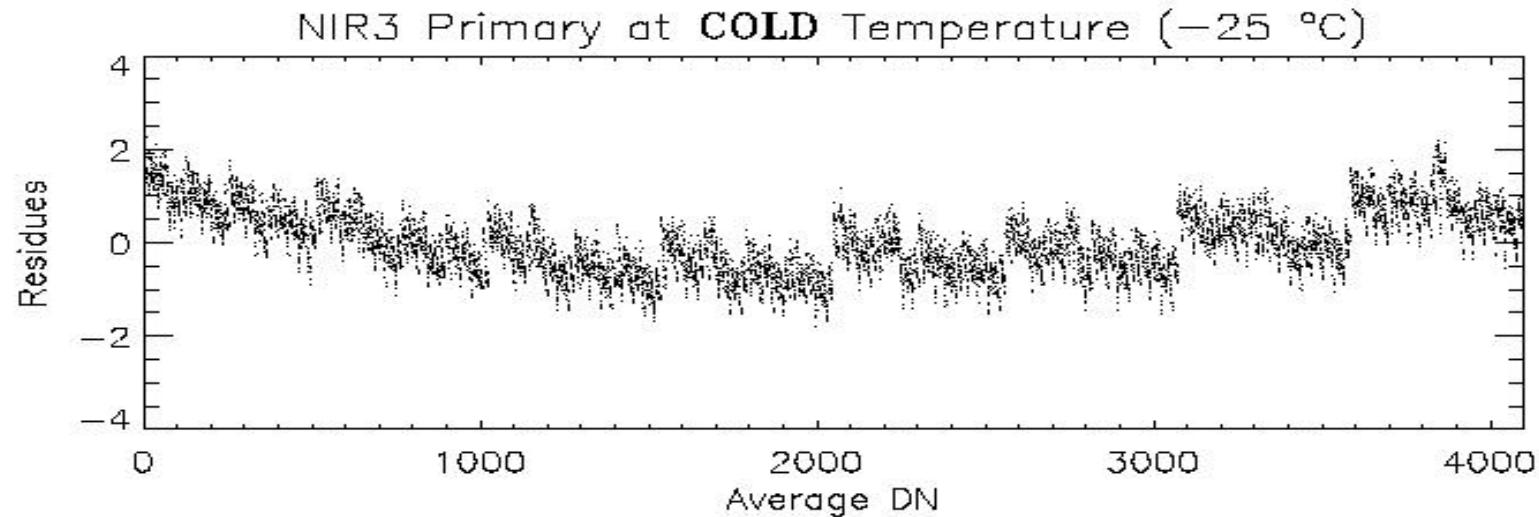
# Component Level Testing



- 16-bit DAC used as calibrator for the 12-bit ADC
- Minimum 2 DAC steps for each of the ADC DN step (level)
- Averaging 200 - 400 samples per DAC step
- BFSL method used to fit the averaged ADC DN data set  
-> Integral Non-linearity (INL)
- No test for LWIR PC bands ADCs
- Periodic pattern observed

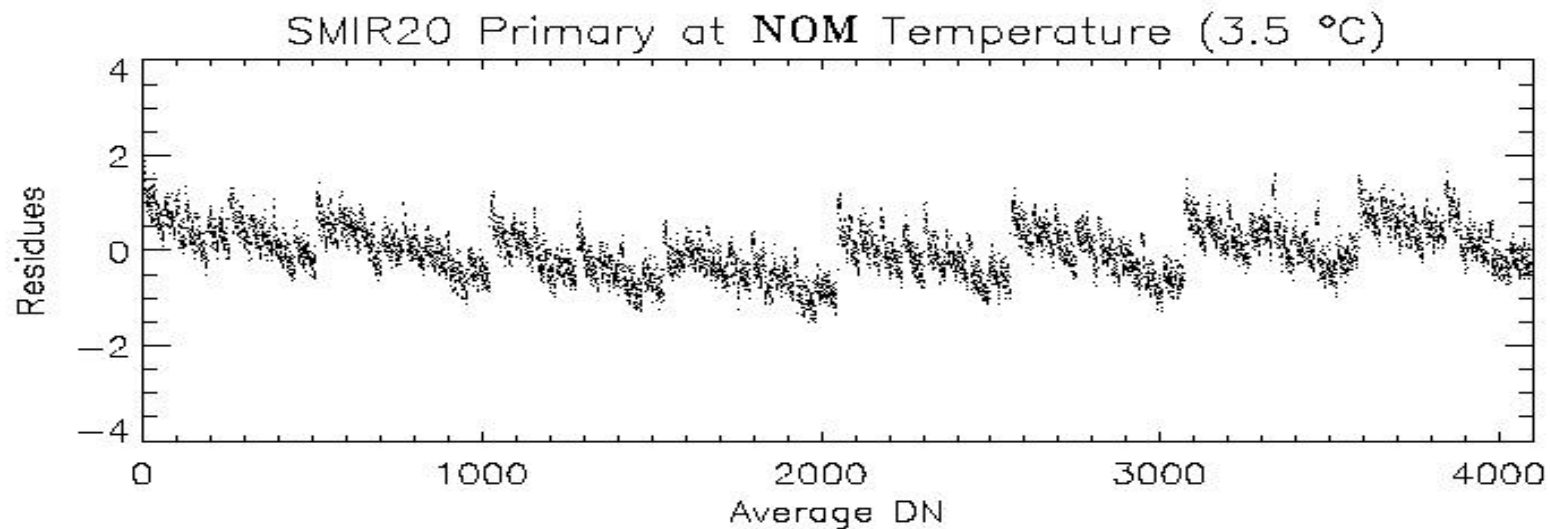
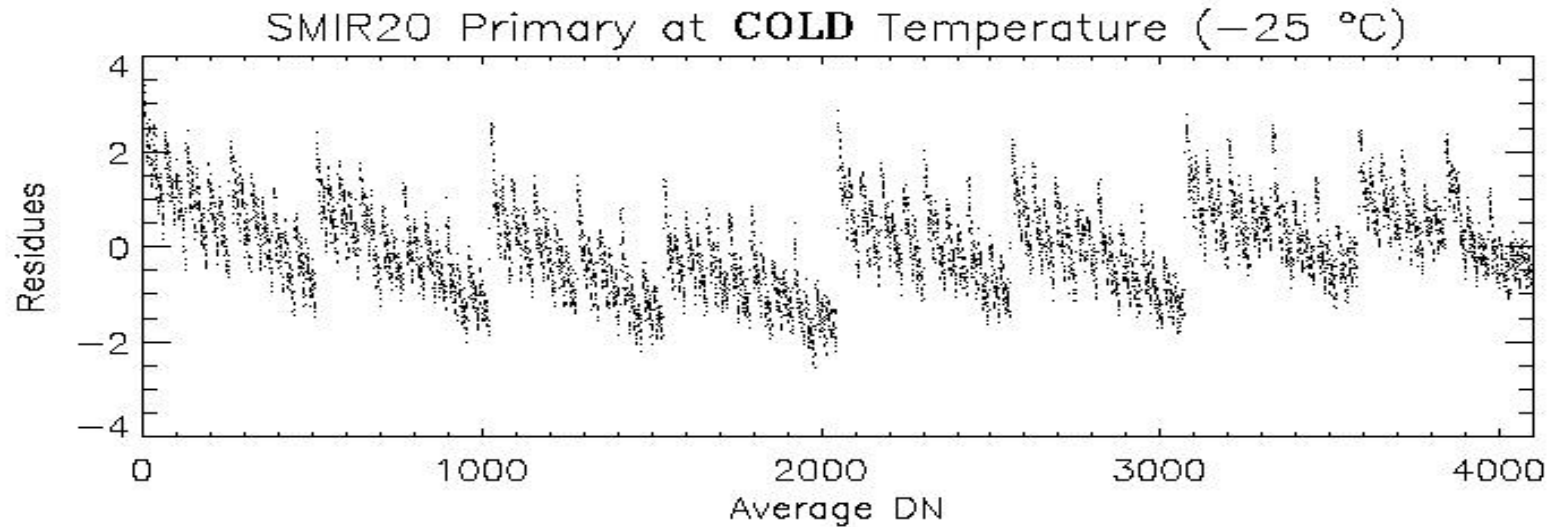


# Component Level Testing



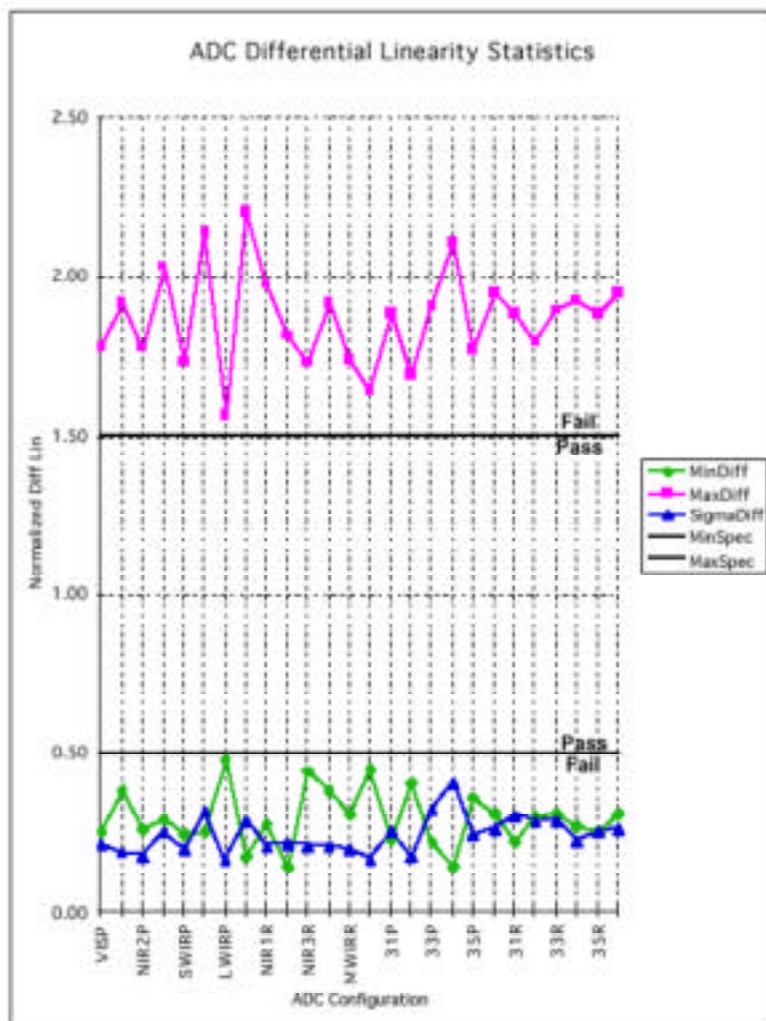


# Component Level Testing





### 3.4.3. Waiver 056: Differential Nonlinearity



This EPS image does not contain a screen preview.  
It will print correctly to a PostScript printer.  
File Name : SMIR20R.EPS  
Creator : GRAF/DRIVE PLUS



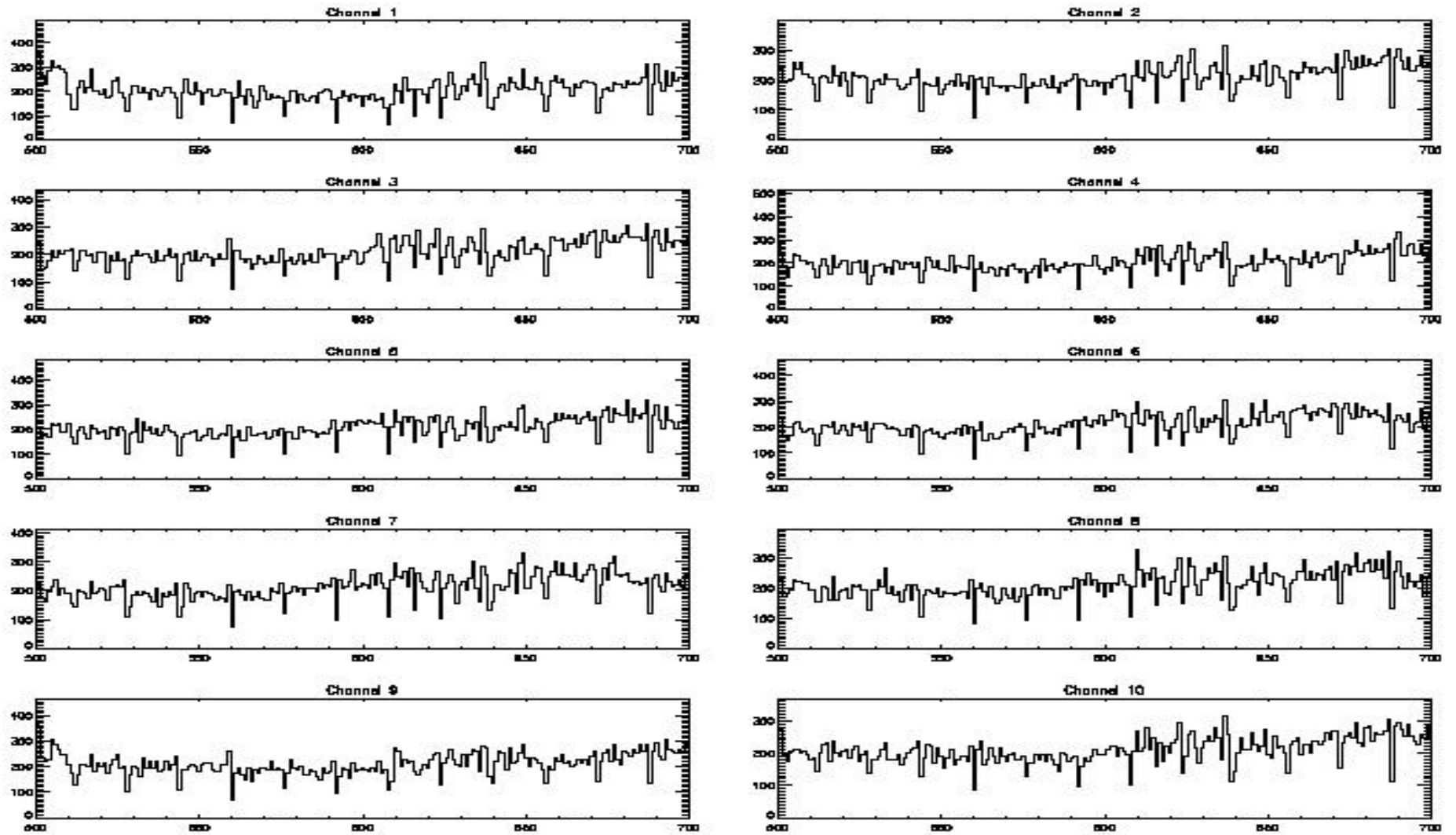
# On-orbit ADC Performance



- MCST EV data histogram for RSB and TEB
  - Some periodic patterns observed
  - PC bands (B31-36) show same channel dependent structure
- B-side (day 313, 25 granules) and A-side (day 77, 12 granules) show similar histogram

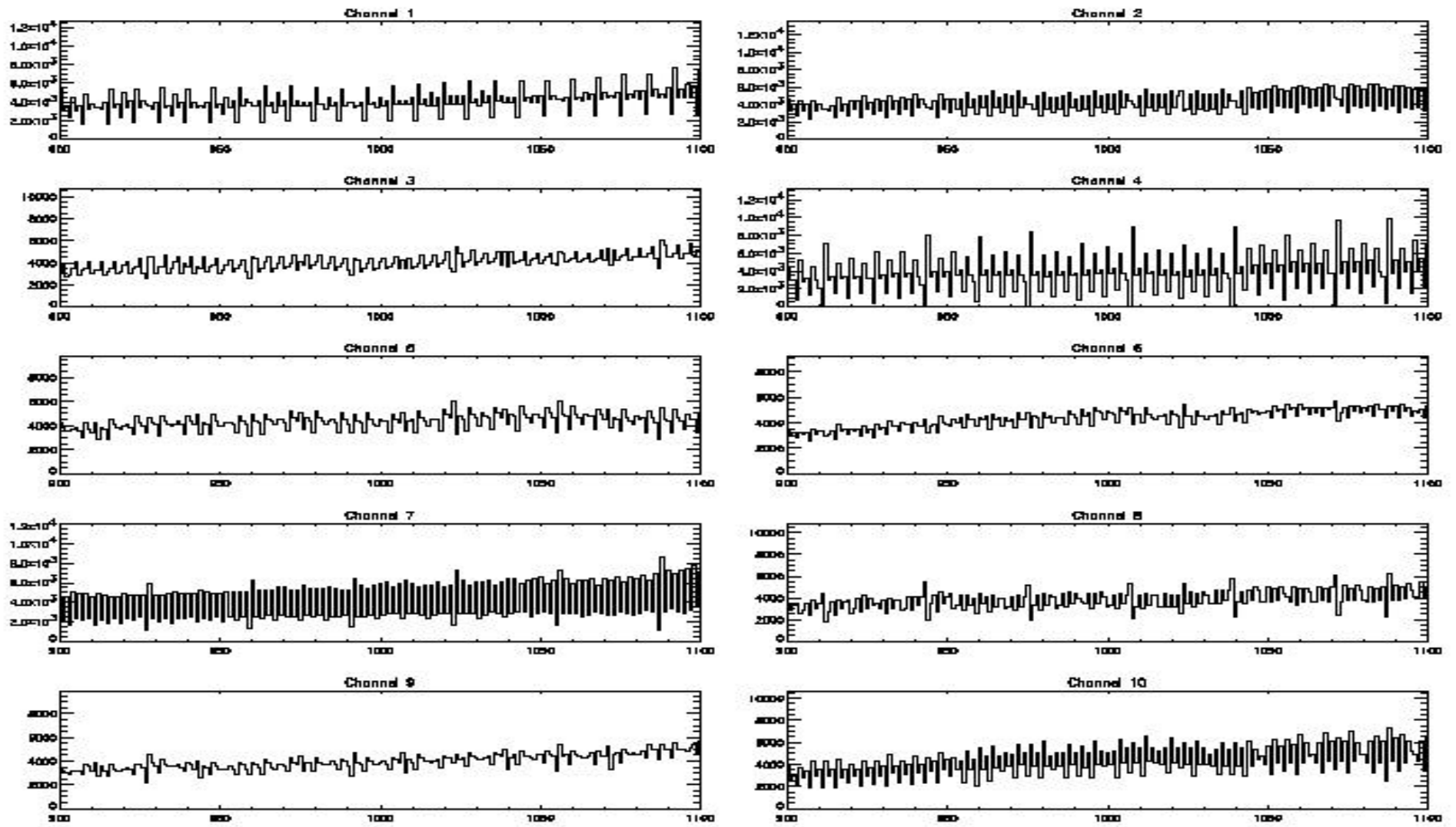
# A-Side

ADC Histogram of Events vs DN For Band 16



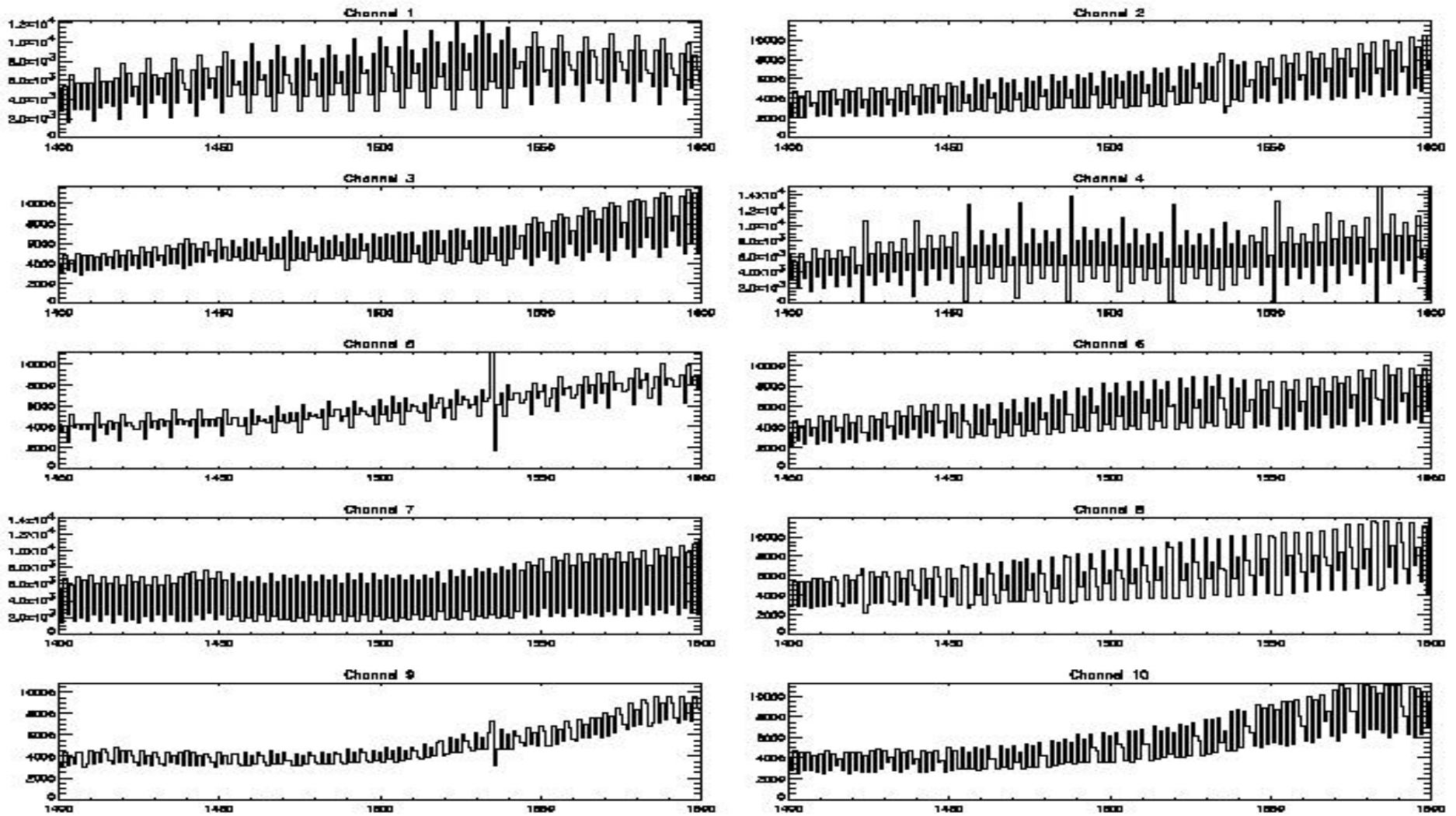
# A-Side

ADC Histogram of Events vs DN For Band 32

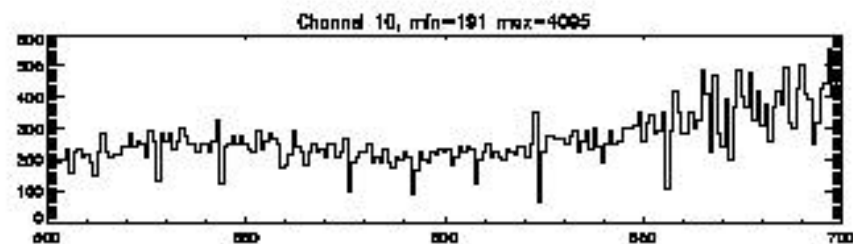
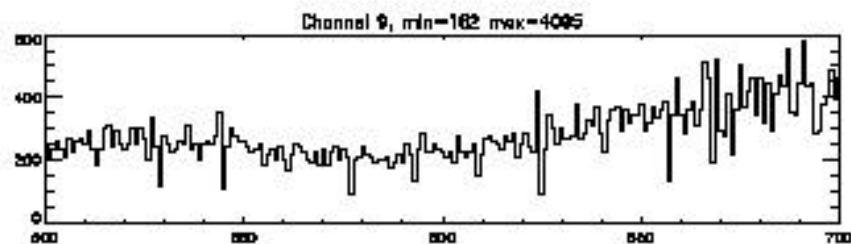
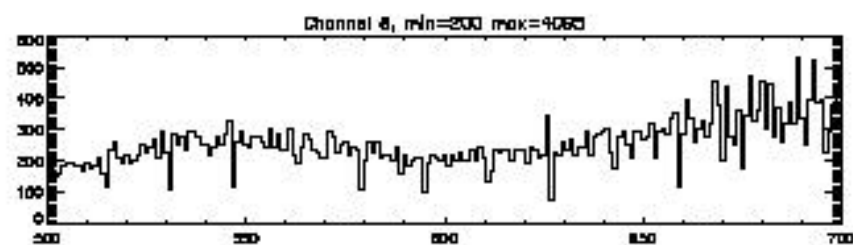
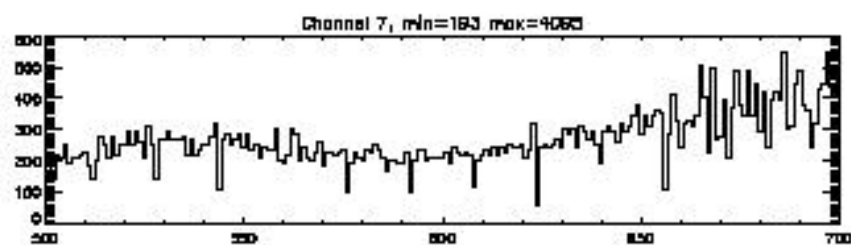
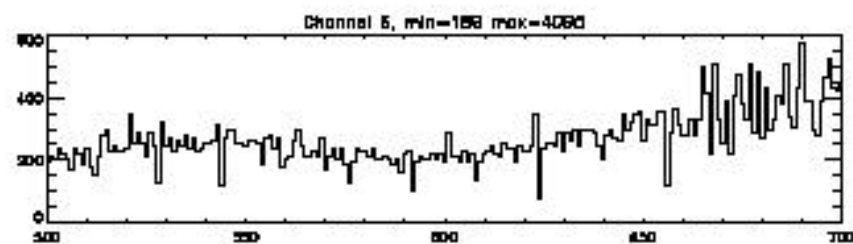
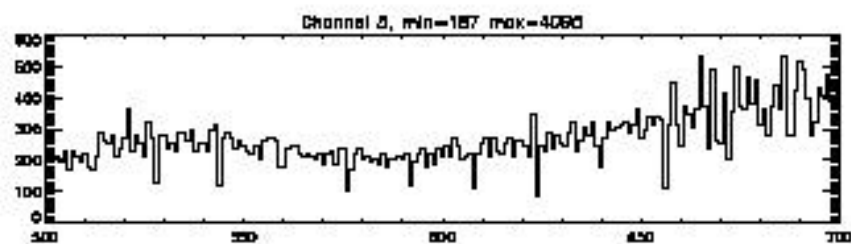
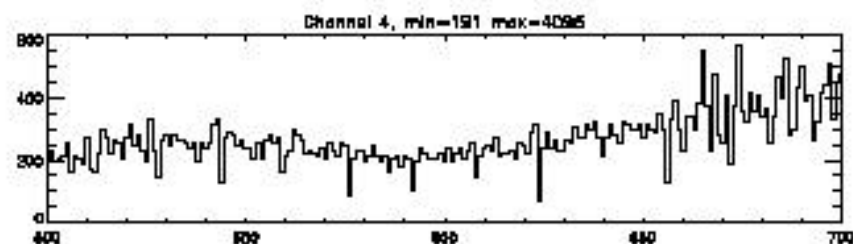
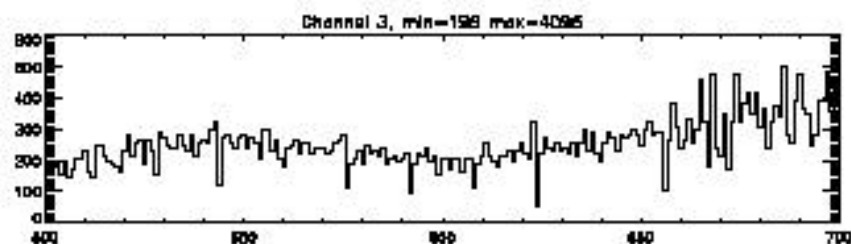
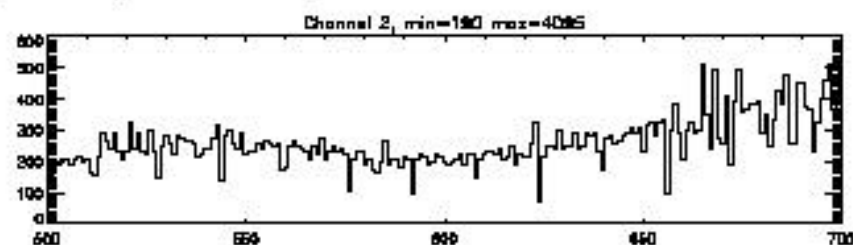
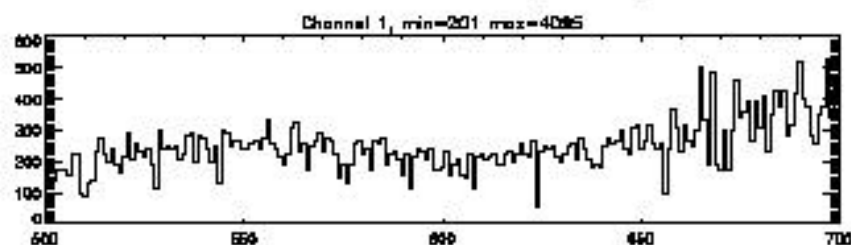


# A-Side

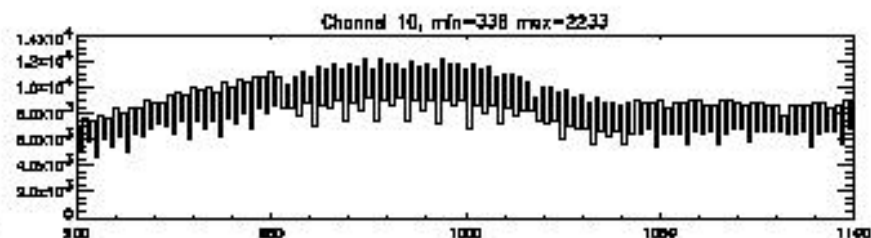
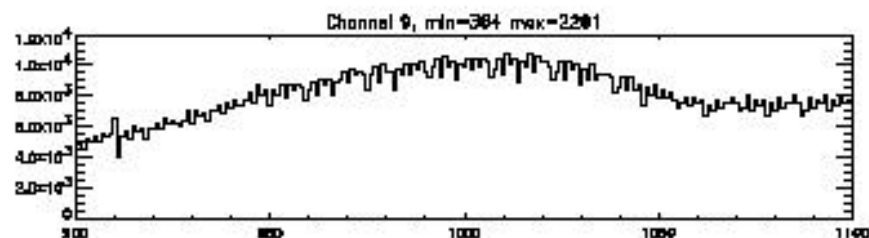
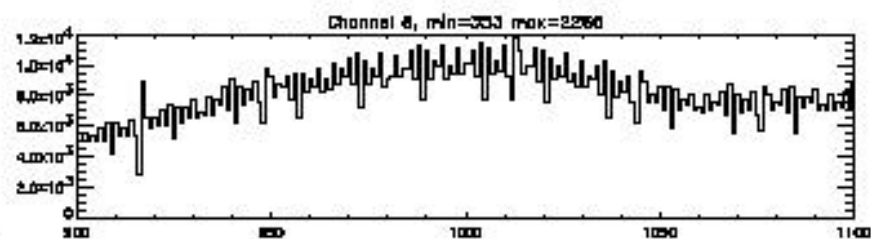
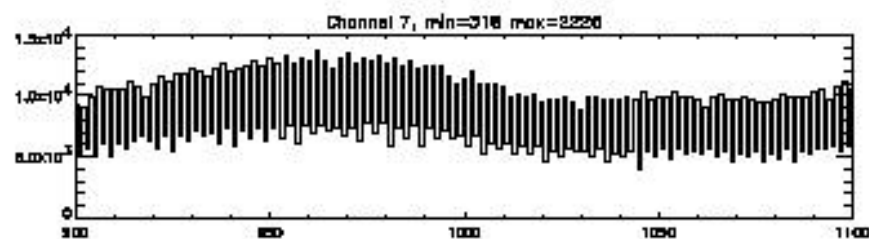
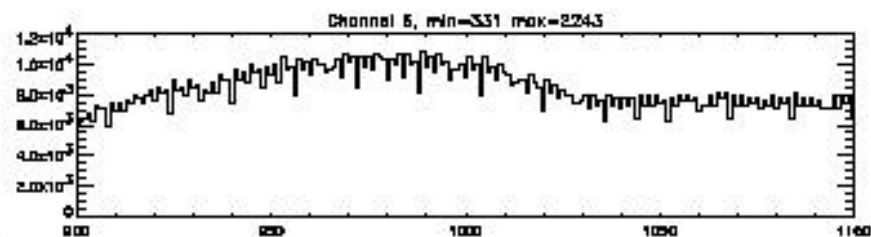
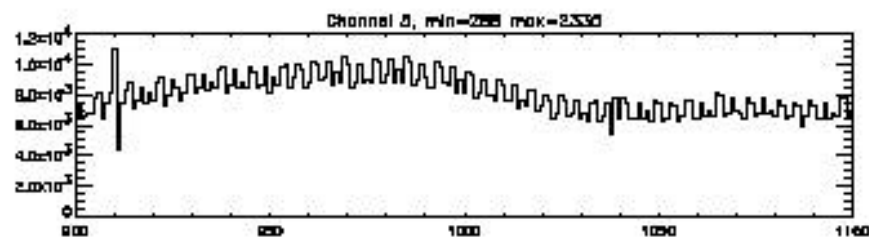
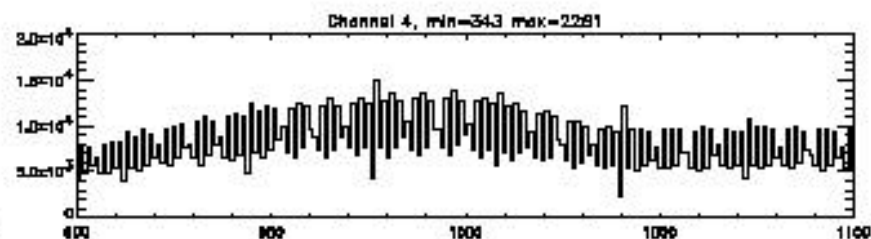
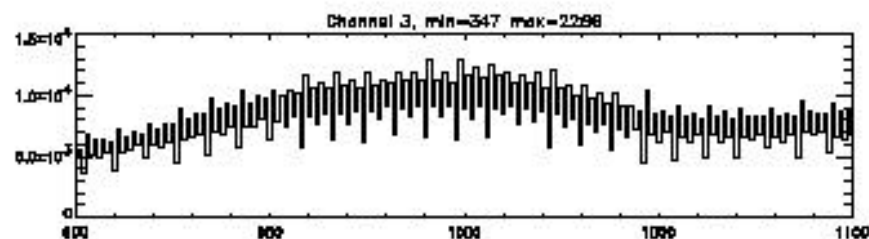
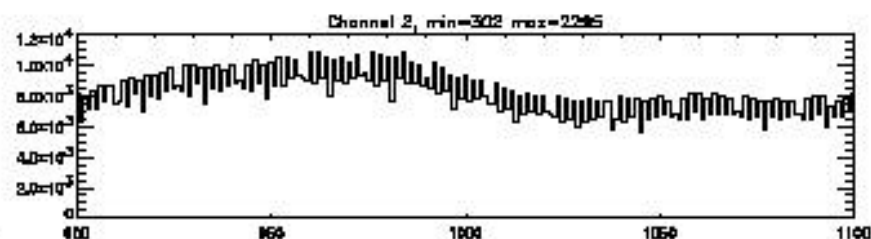
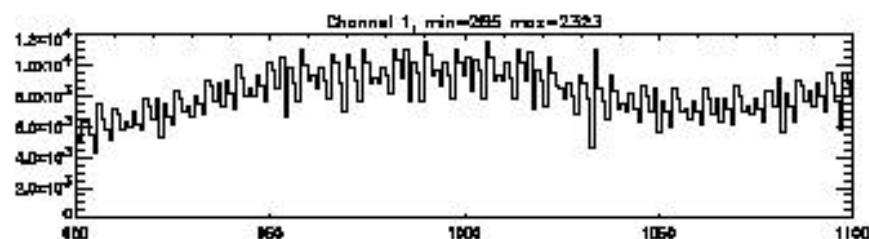
ADC Histogram of Events vs DN For Band 34



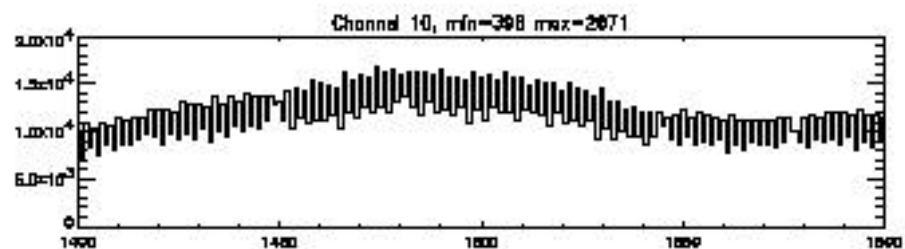
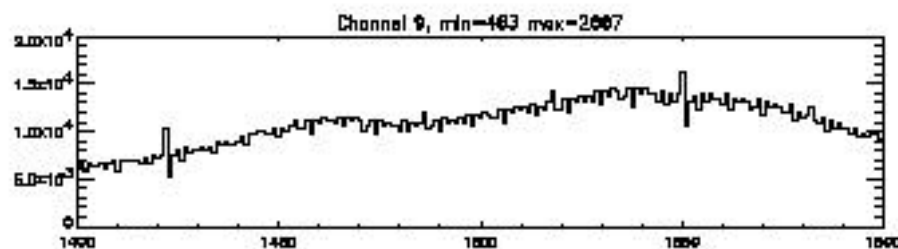
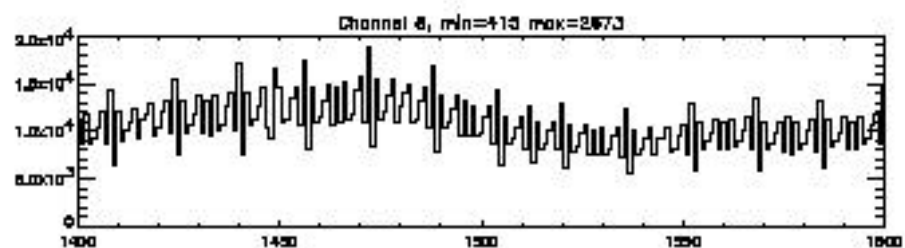
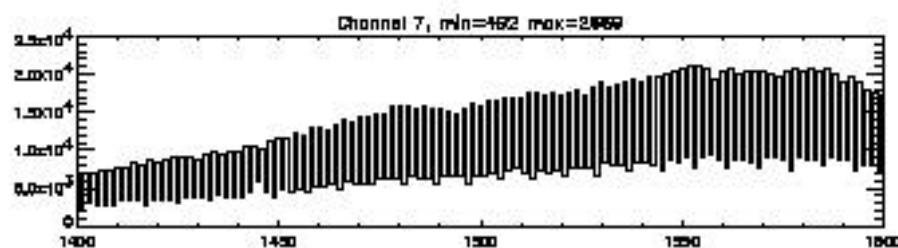
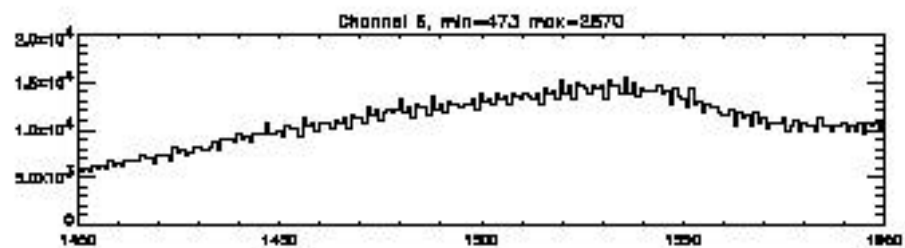
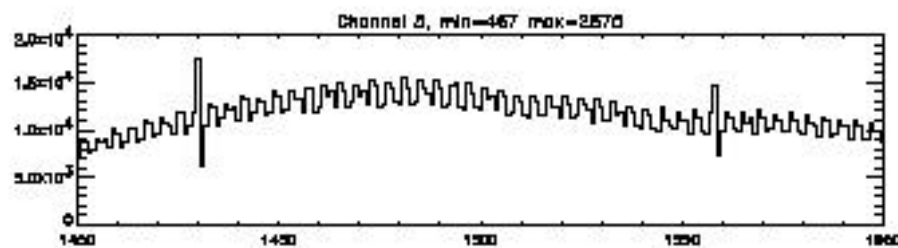
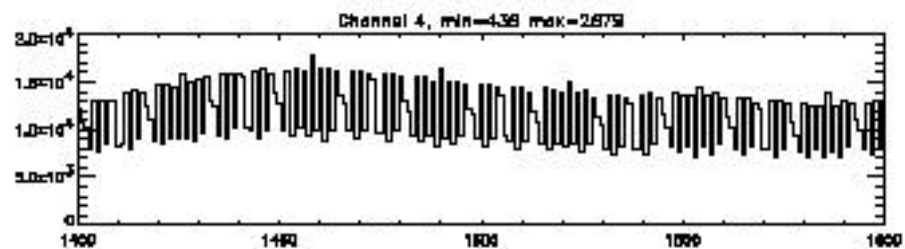
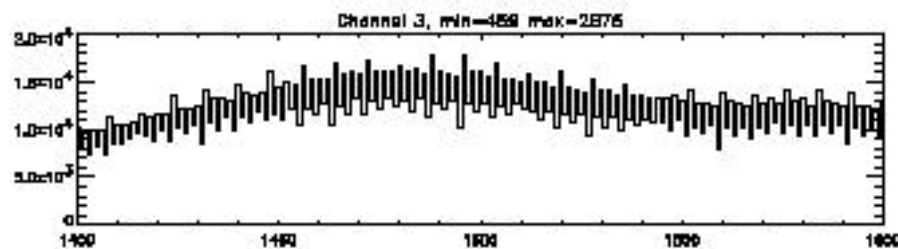
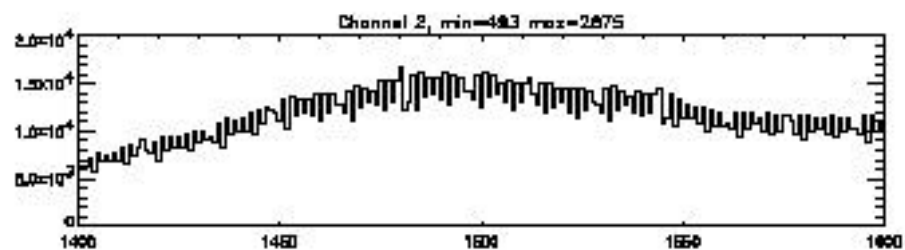
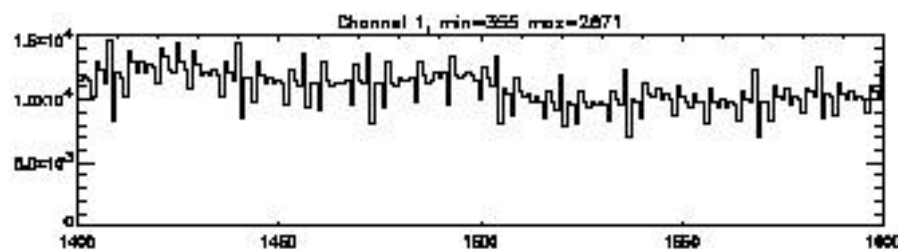
*MODIS ADC vs Events for Band 16*  
Bside, Day313(08:20-10:20), Itwk/Ved=79/110



# *MODIS ADC vs Events for Band 32* Bside, Day313(08:20-10:20), Itwk/Vedt=79/110



# *MODIS ADC vs Events for Band 34* Bside, Day313(08:20-10:20), Itwk/Vedt=79/110



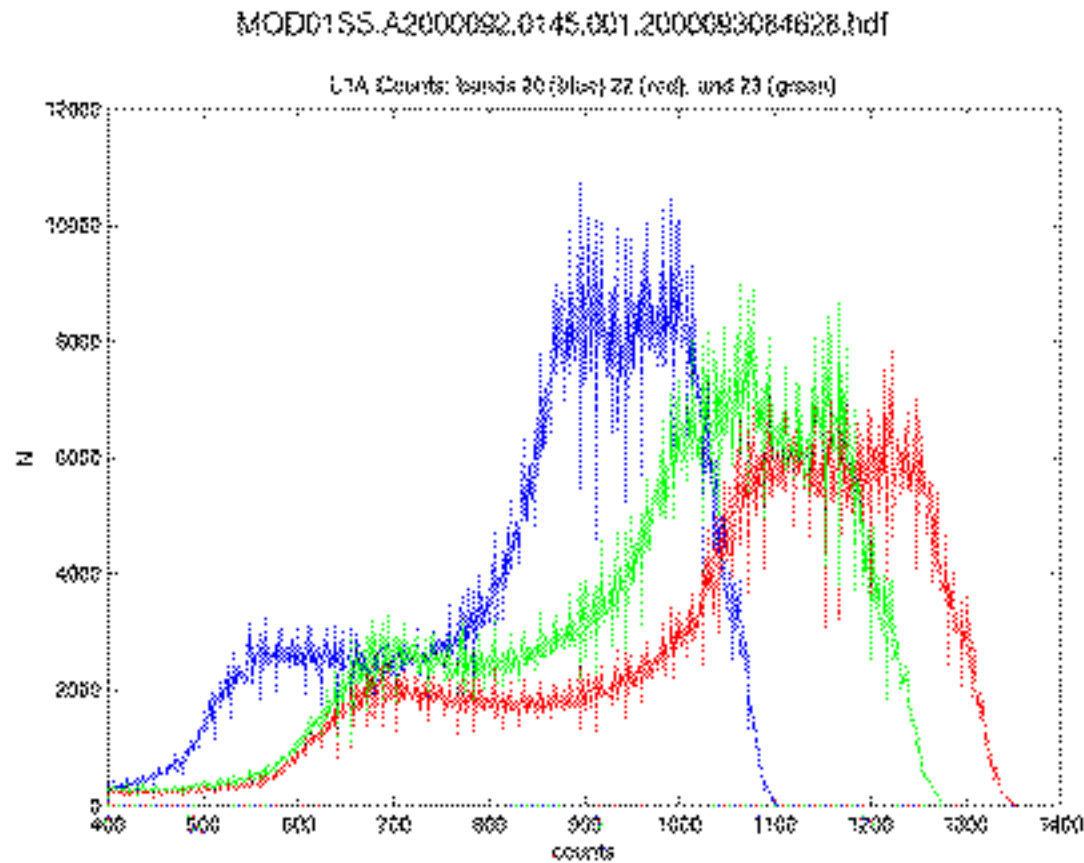


Figure 1. Histogram of the occurrence of digital counts in Bands 20, 22 and 23 at Level 1a processing for one granule. Output from all channels (detectors) are shown. (U. Miami/Minnett, et. al.

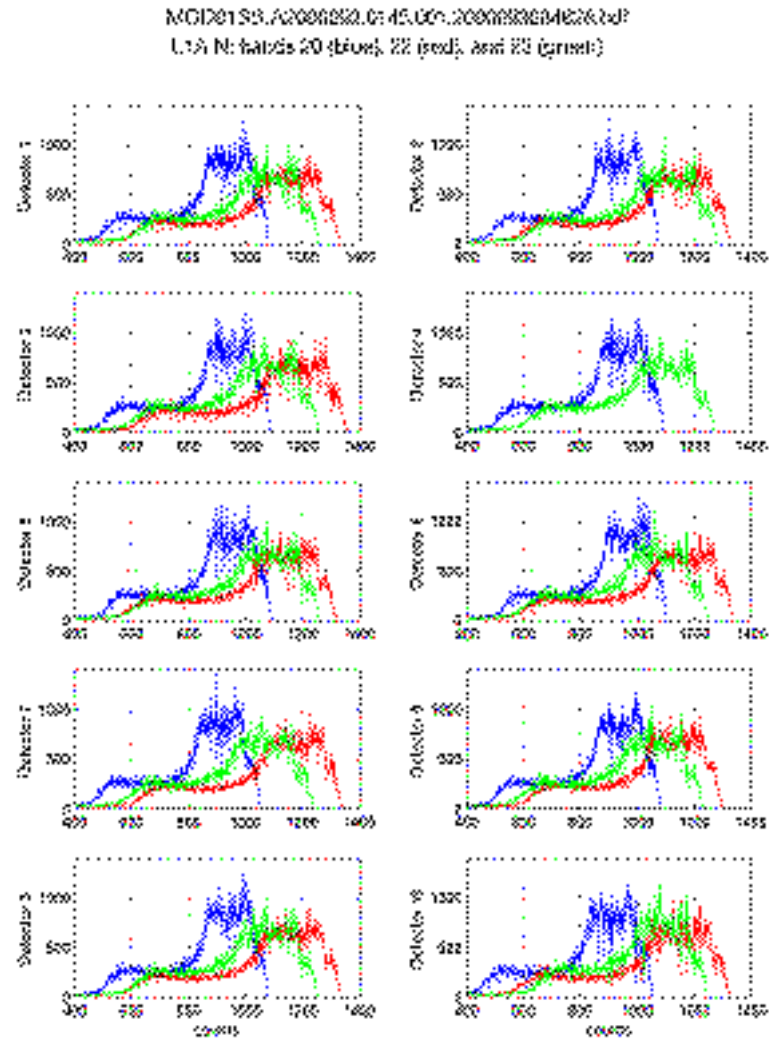


Figure 4. As figure 1, but output from each channel (detector) shown separately. Note the output from detector 4 of band 22 is biased off-scale. (U. Miami/ Minnett, et. al.

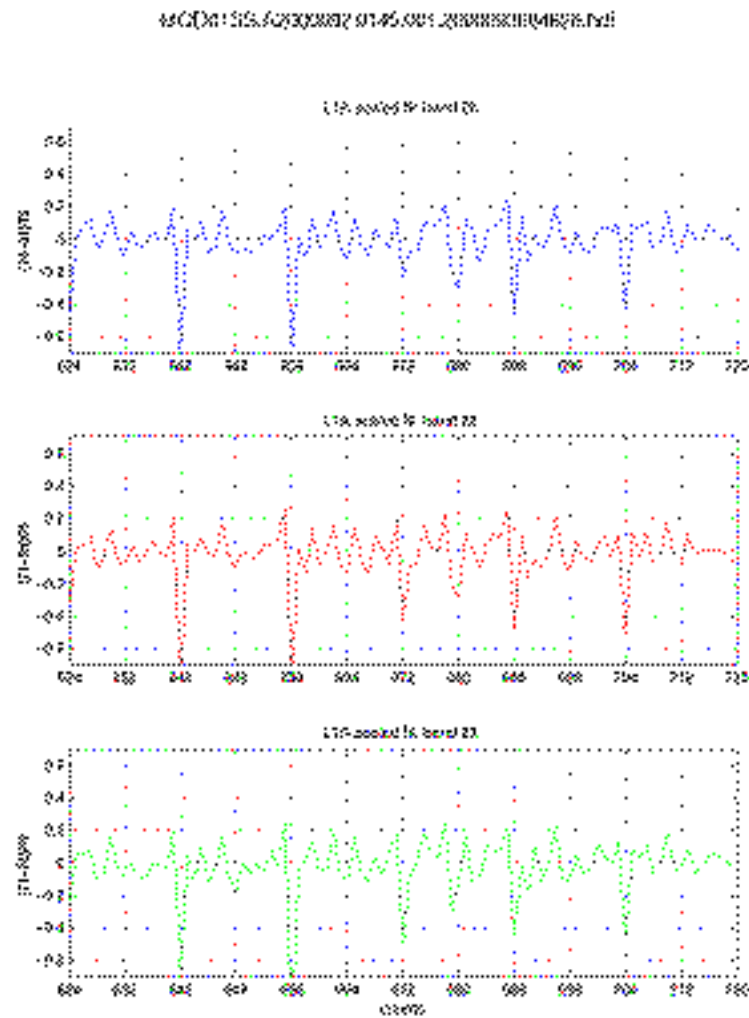


Figure 3. As Figure 2, but on an expanded horizontal scale. (U. Miami/Minnett, et. al.)

M0D01S3.A20080920.07\_45.006.20080920EIM42K.bdf  
L1A N : nans 57 (bad) atz 52 (free)

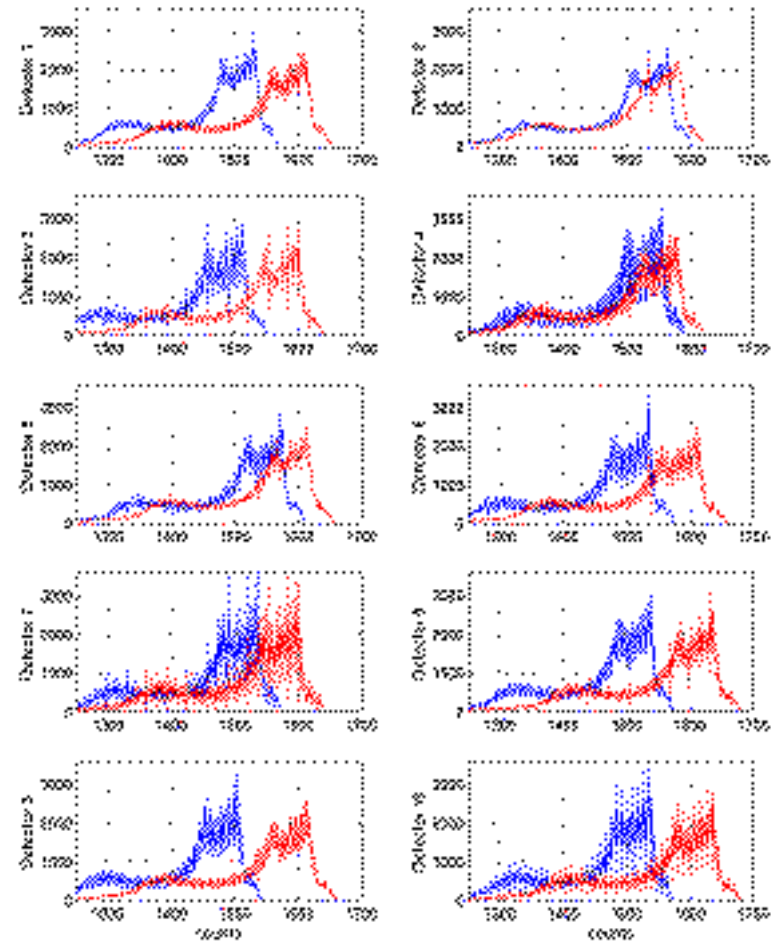


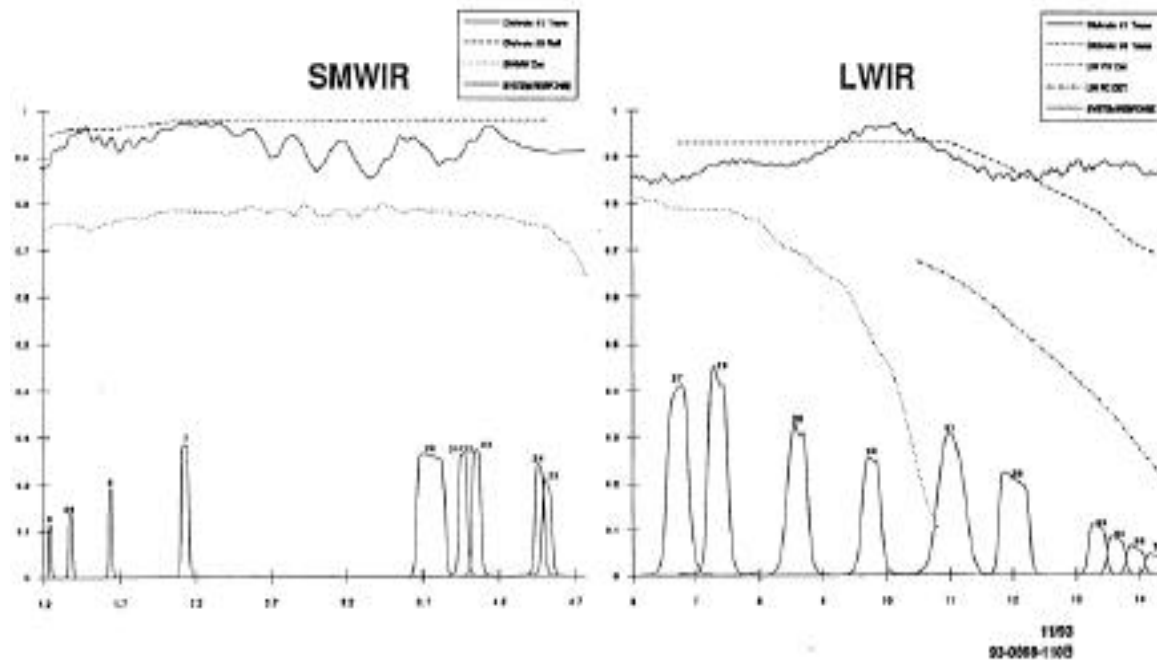
Figure 15. As figure 12, but output from each channel (detector) shown separately. (U. Miami/Minnett, et. al.)



## MODEL IDENTIFIES KEY CONTRIBUTORS TO SYSTEM WAVELENGTH, TRANSMISSION (SMWIR, LWIR)

**HUGHES**

SANTA BARBARA RESEARCH CENTER  
A DIVISION OF



Note: Model anticipates SW/MWIR detector roll-off at 4.7  $\mu\text{m}$



# On-orbit to Pre-launch 500m $m_1$ Coefficients



**Fractional On-orbit to Pre-launch Differences  
for the MODIS 500m Bands  
Expanded Scale**

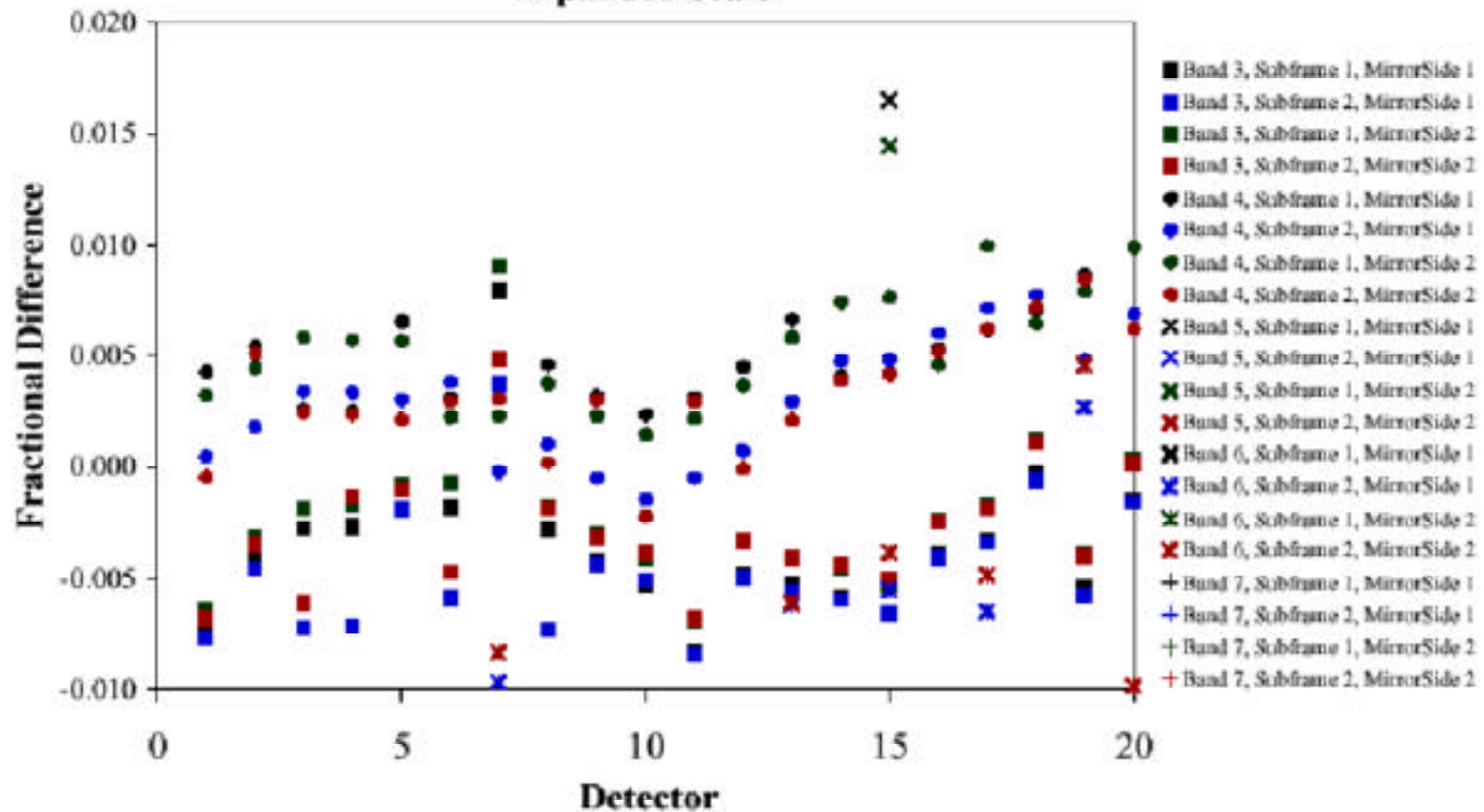
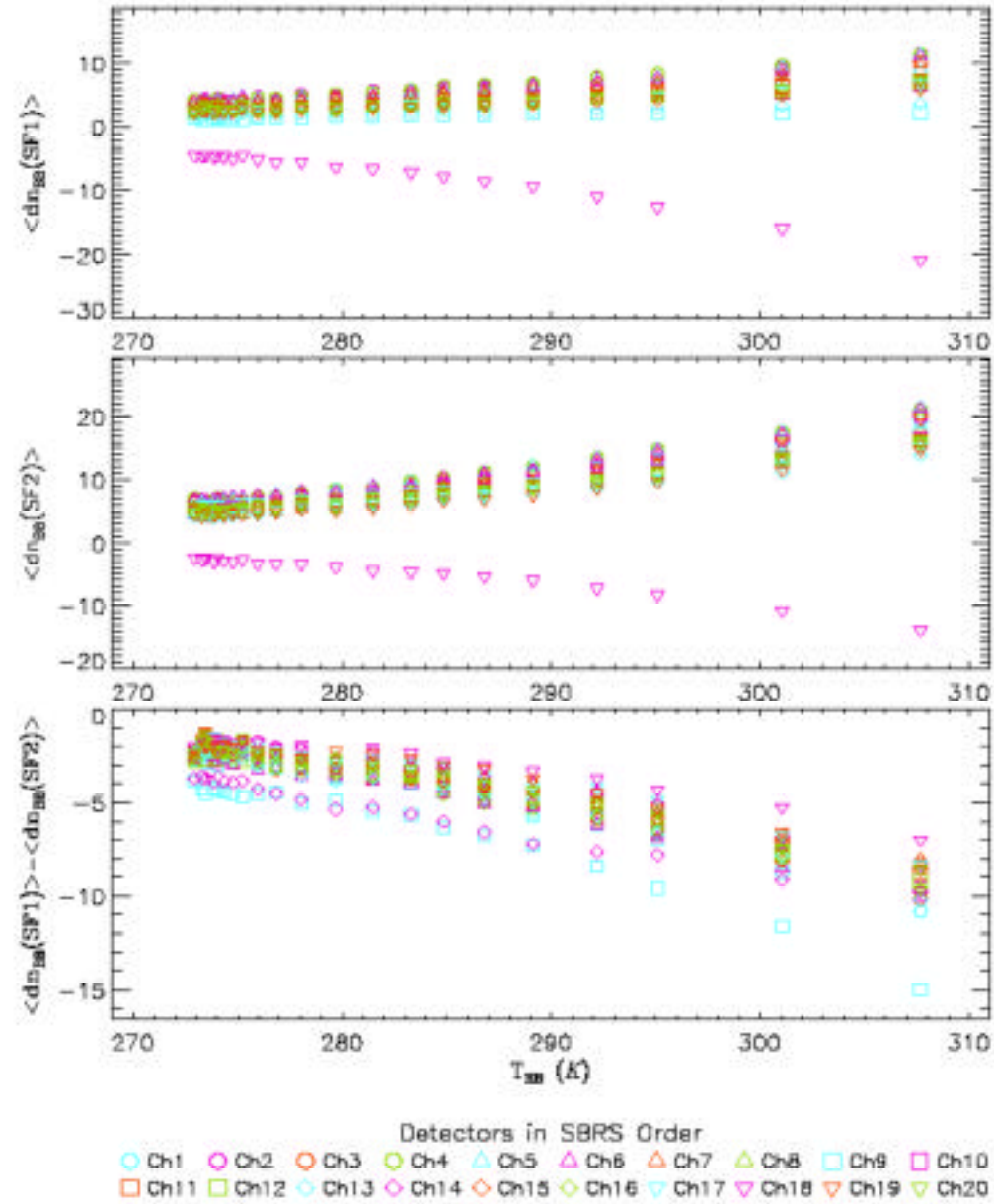
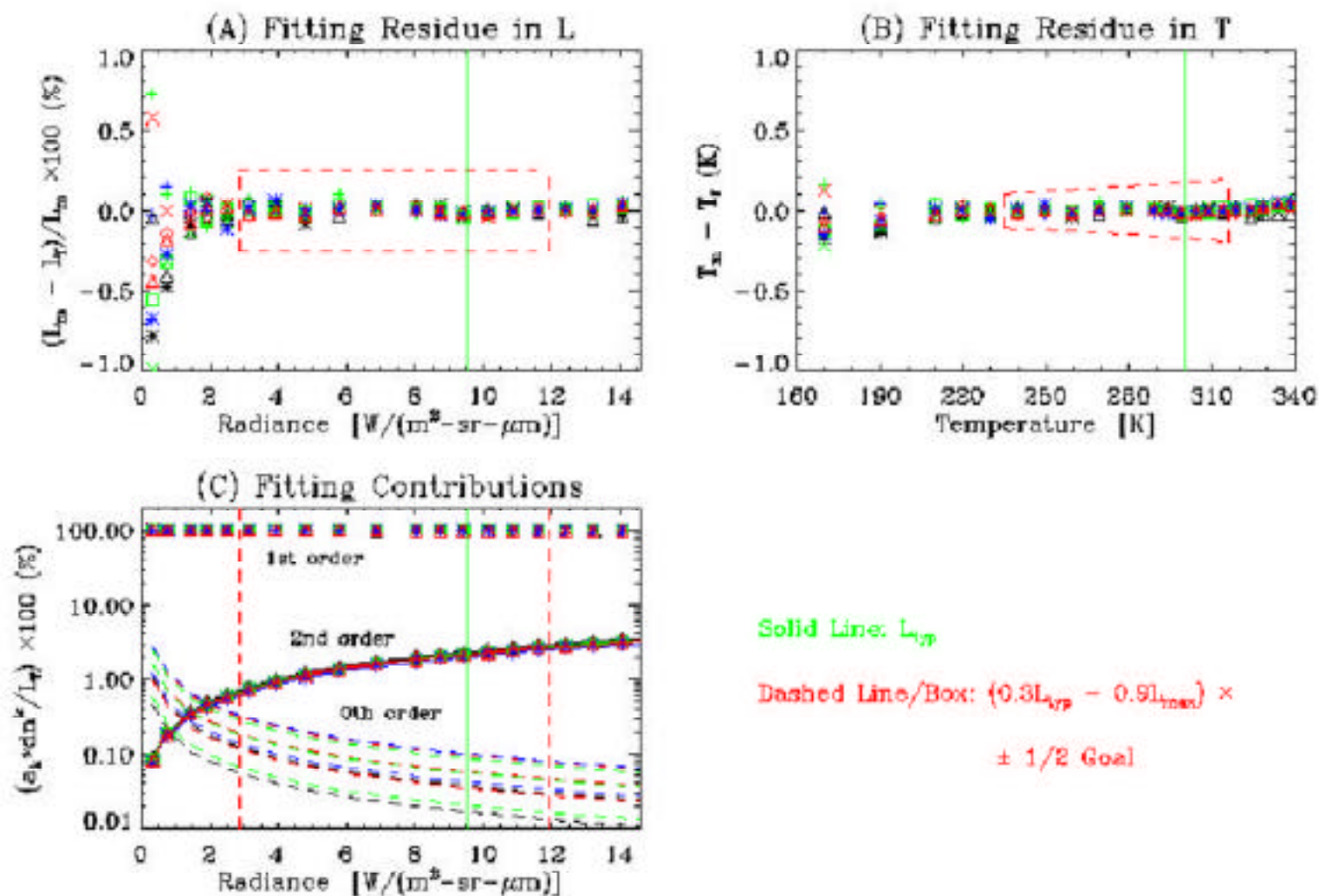


Figure 2. PFM B6  $\langle dn_{\text{SB}} \rangle$  of Sub-frame 1, 2, and their Difference  
OBC Data Sets Selected during OA-26 (Day 83-84)  
(Blackbody Cool-down Activity; Itwk:100/Vdet:210)

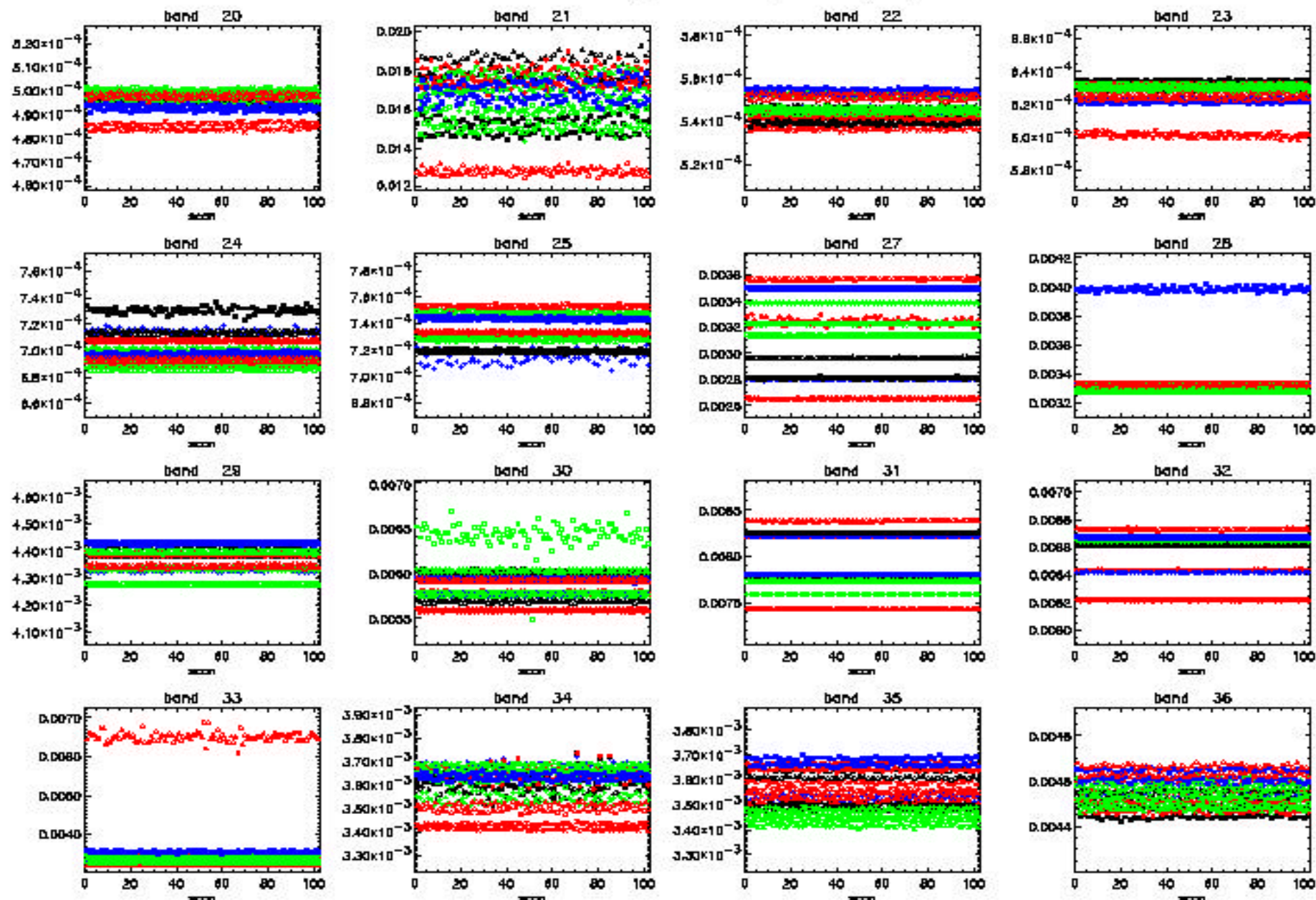




# B31 (11.0 $\mu\text{m}$ ) L vs DN Quadratic Fitting Results at Nominal Plateau (273 K)

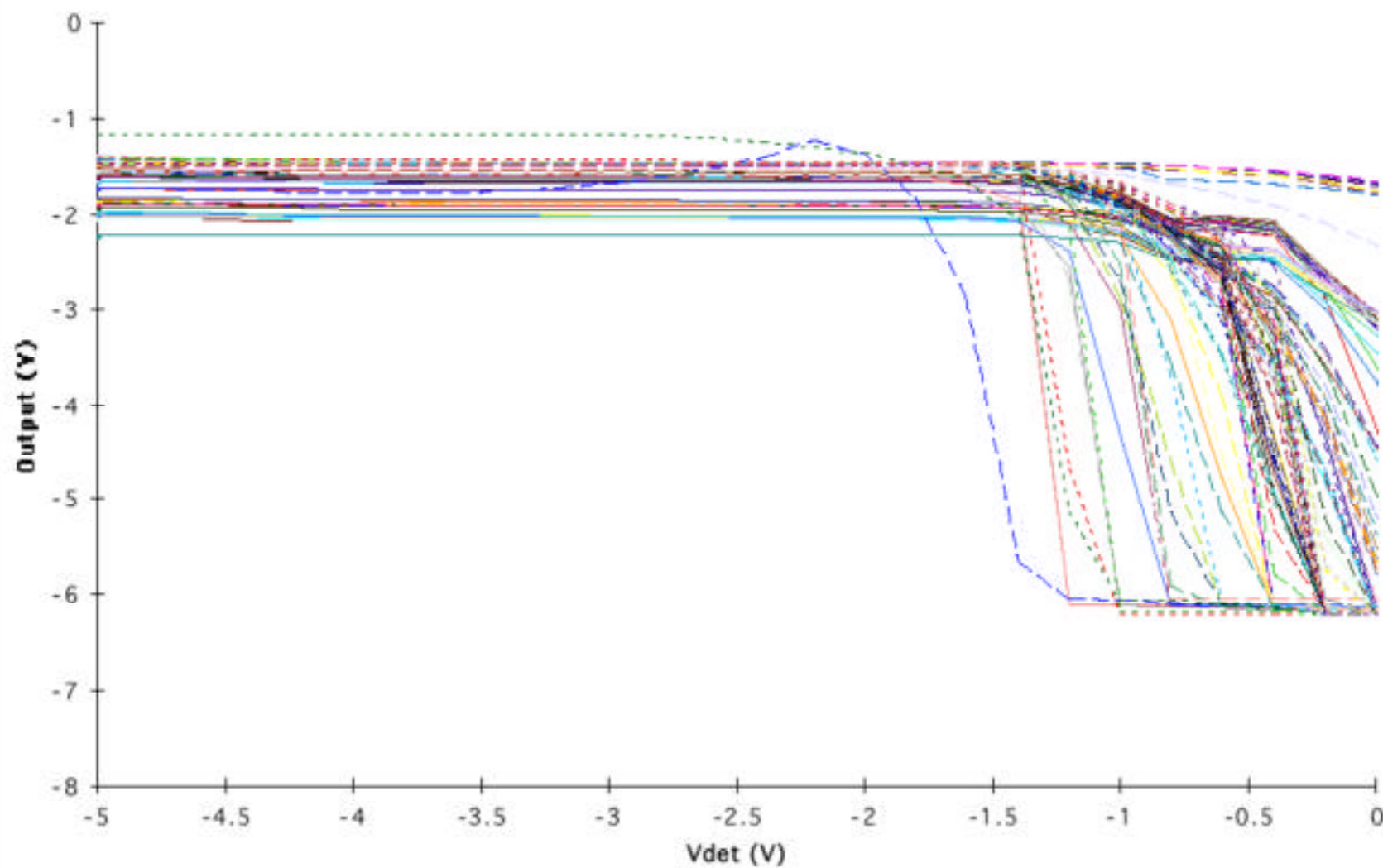


# MODIS PFM On-orbit TEB System Response (b1) vs Scan

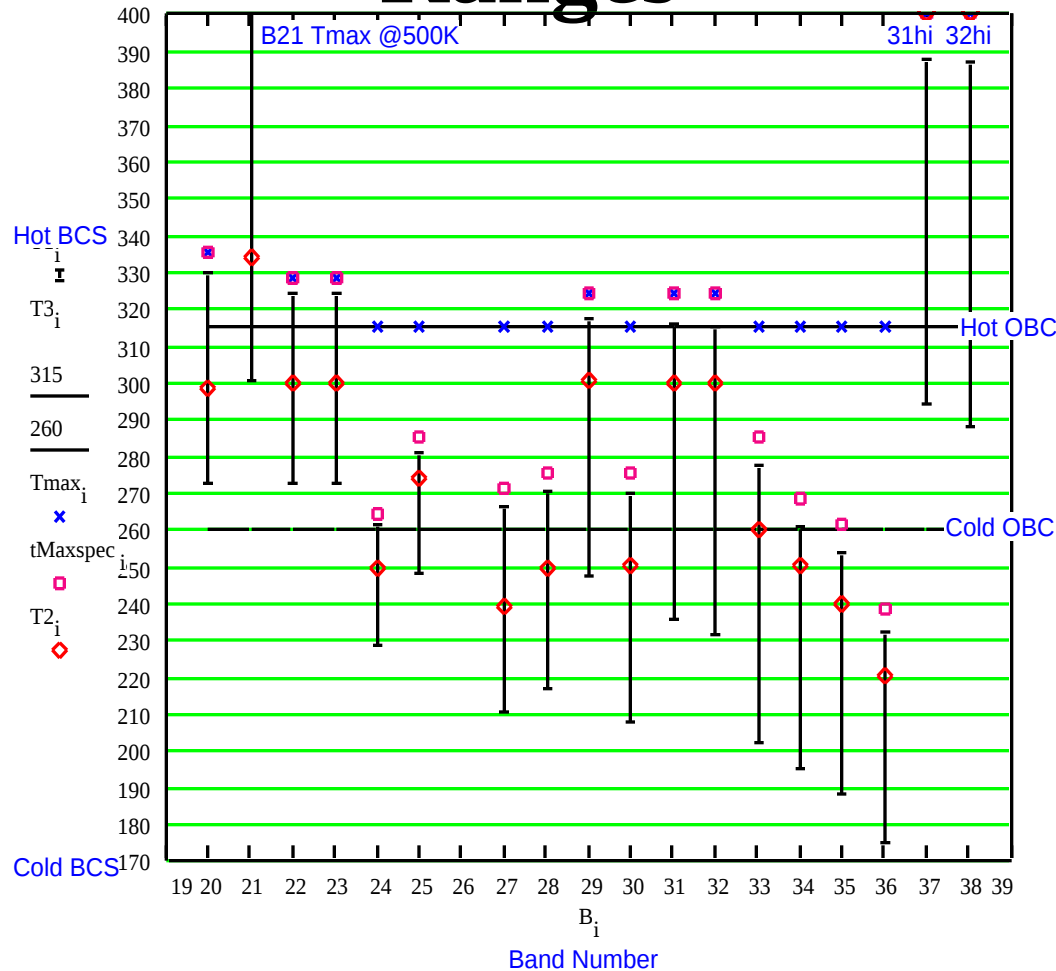


Ch1:Red x Ch2:Blu + Ch3:Blk \* Ch4:Blk Δ Ch5:Red ◇ Ch6:Grn □ Ch7:Grn x Ch8:Grn + Ch9:Blu \* Ch10:Red Δ  
 Data collect time:306.1623 T<sub>avg</sub>:[274.162 - 274.196] detector in SBRS order Itwk/Vdet = 79/110

PFM SCA Data, Vdet Sweep, PFS109, Primary A, 85K, Tseq# 72148494,  
1/22/95



# MODIS Operational and Calibration Ranges



Vertical bars represent temperatures of 0.3Ltyp and 0.9 Lmax. Box symbols represent  $Tmax\_spec$ , X symbols represent the value of  $Tmax\_set$  (the higher of  $tMaxspec$  or 315K), and the diamond symbols represent the temperature of  $Ltyp$ .



## RSB SIS-100 Round Robin (RR)



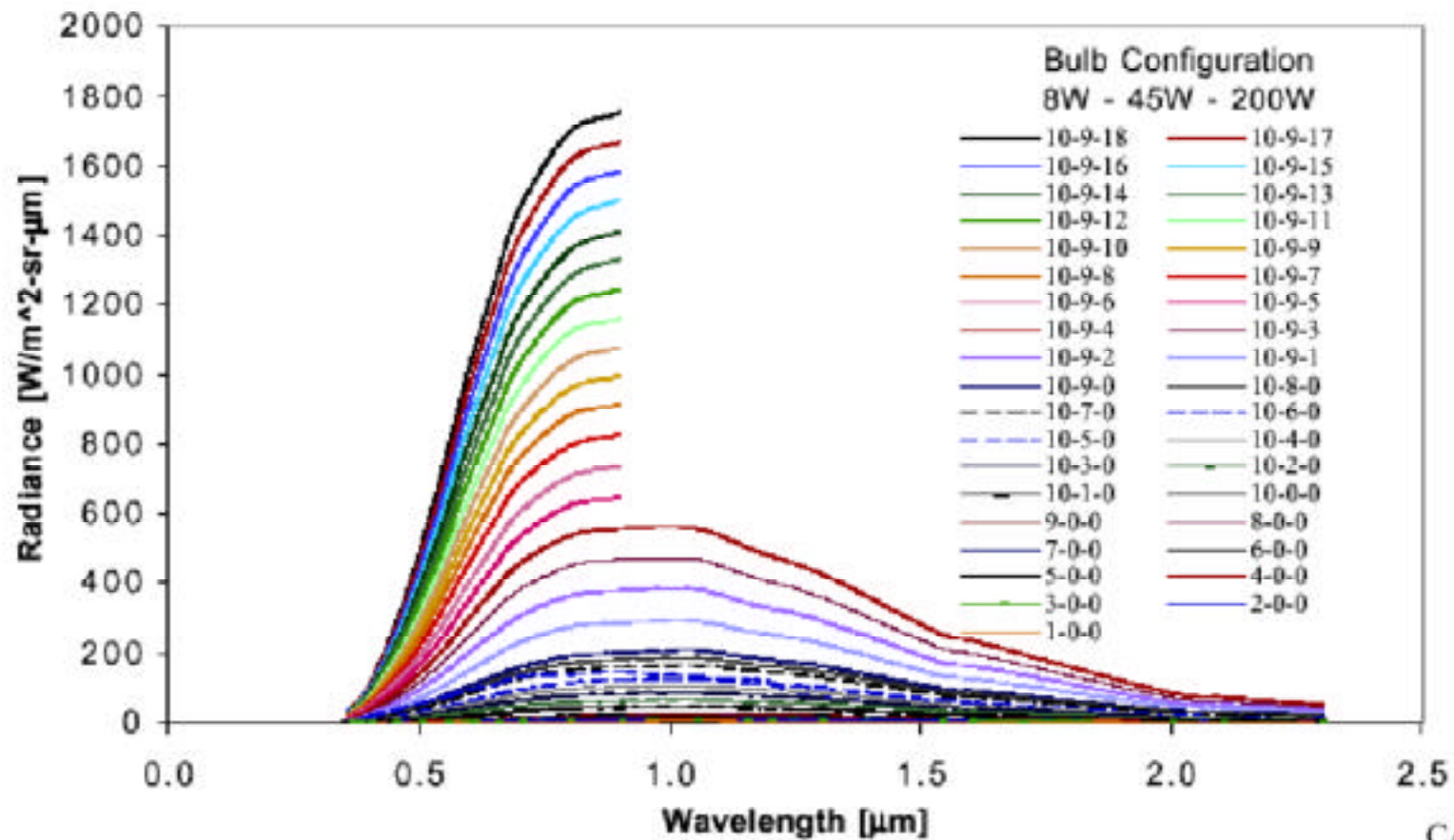
- SIS(100) was calibrated at SBRS for the RR in 1996 (PFM) and 1998 (FM1)
- Due to burnout of the 45W bulbs, the PFM SIS(100) calibration has not been RR validated
- The 1998 data has been turned in for the RR and the FM1 calibration will be RR validated when results are published.
- Preliminary results indicate consistency between NIST standards and the SBRS SIS(100) scale at about 2-3% for FM1 Testing



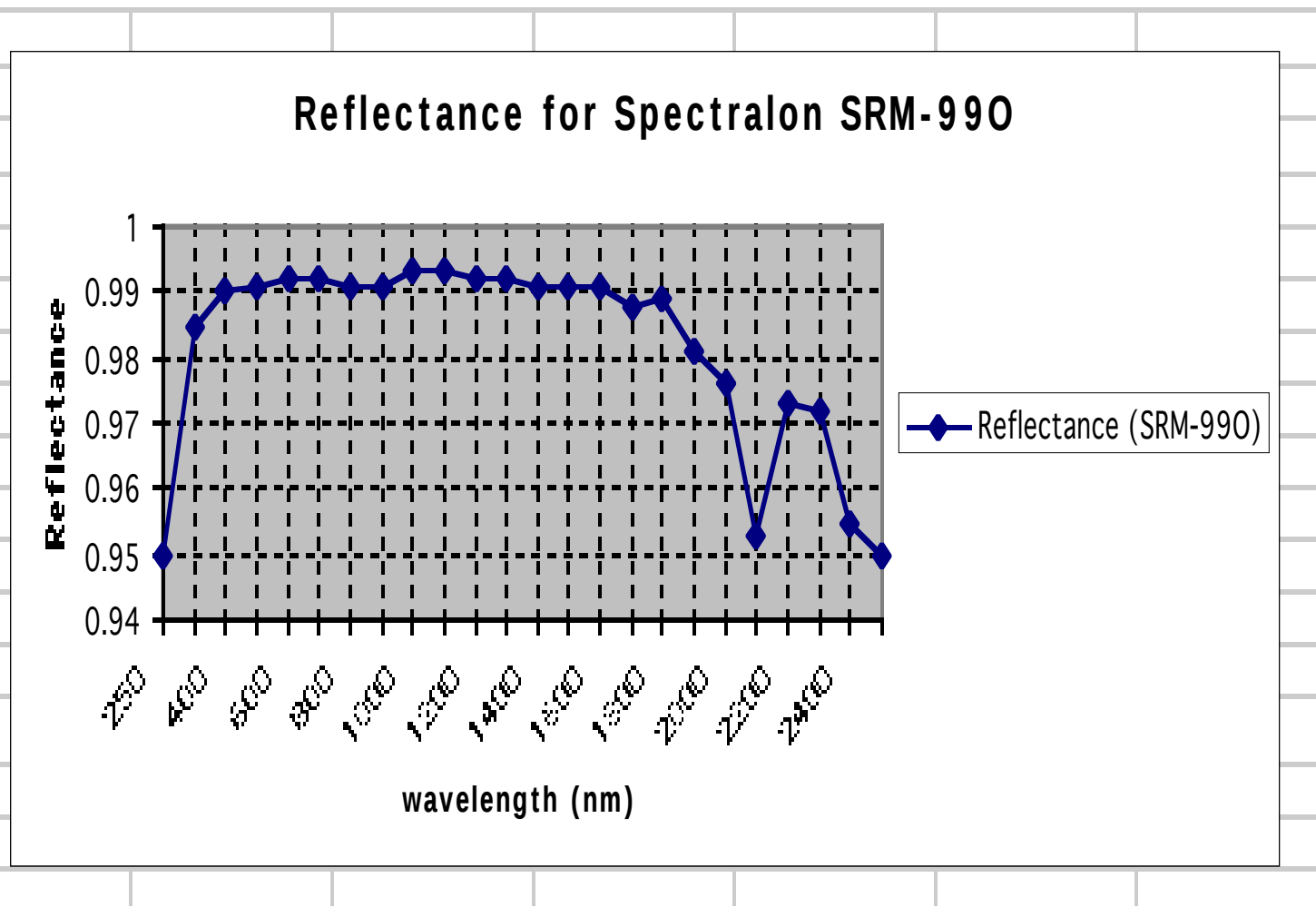
# SBRS SIS(100) Radiance as used for FM1



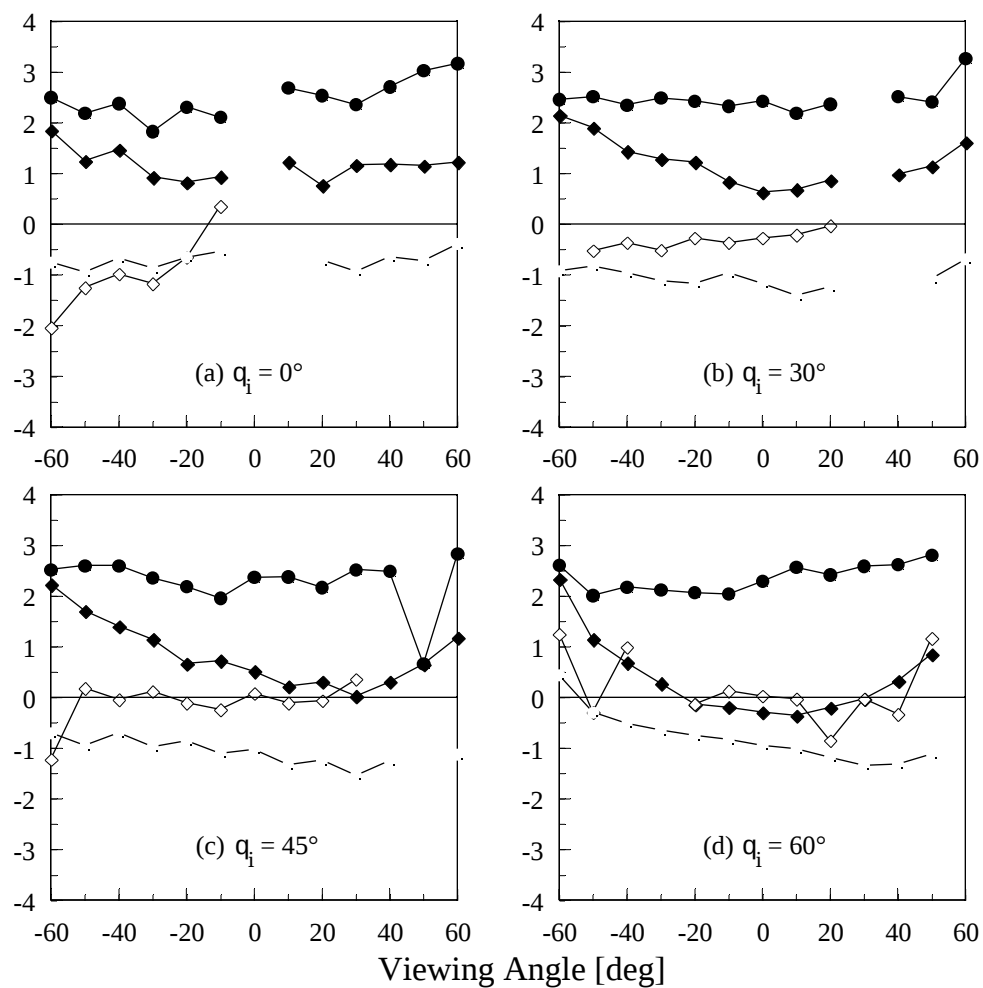
SIS(100) Radiance



# Reflectance of MODIS Solar Diffuser Material



Spectralon,  $\lambda = 440 \text{ nm}$



Laboratory

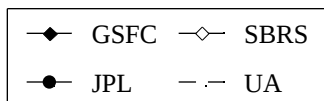


Figure 10. Difference of BRDF relative to the NIST value as a function of viewing angle for Spectralon at 440 nm. The incident angles are indicated in the panels, and the participants are given in the legend. From Early, et.al., J. A. O. Tech, **17**, 1077-1091 (2000)

Figure 10

Spectralon,  $\lambda = 633 \text{ nm}$

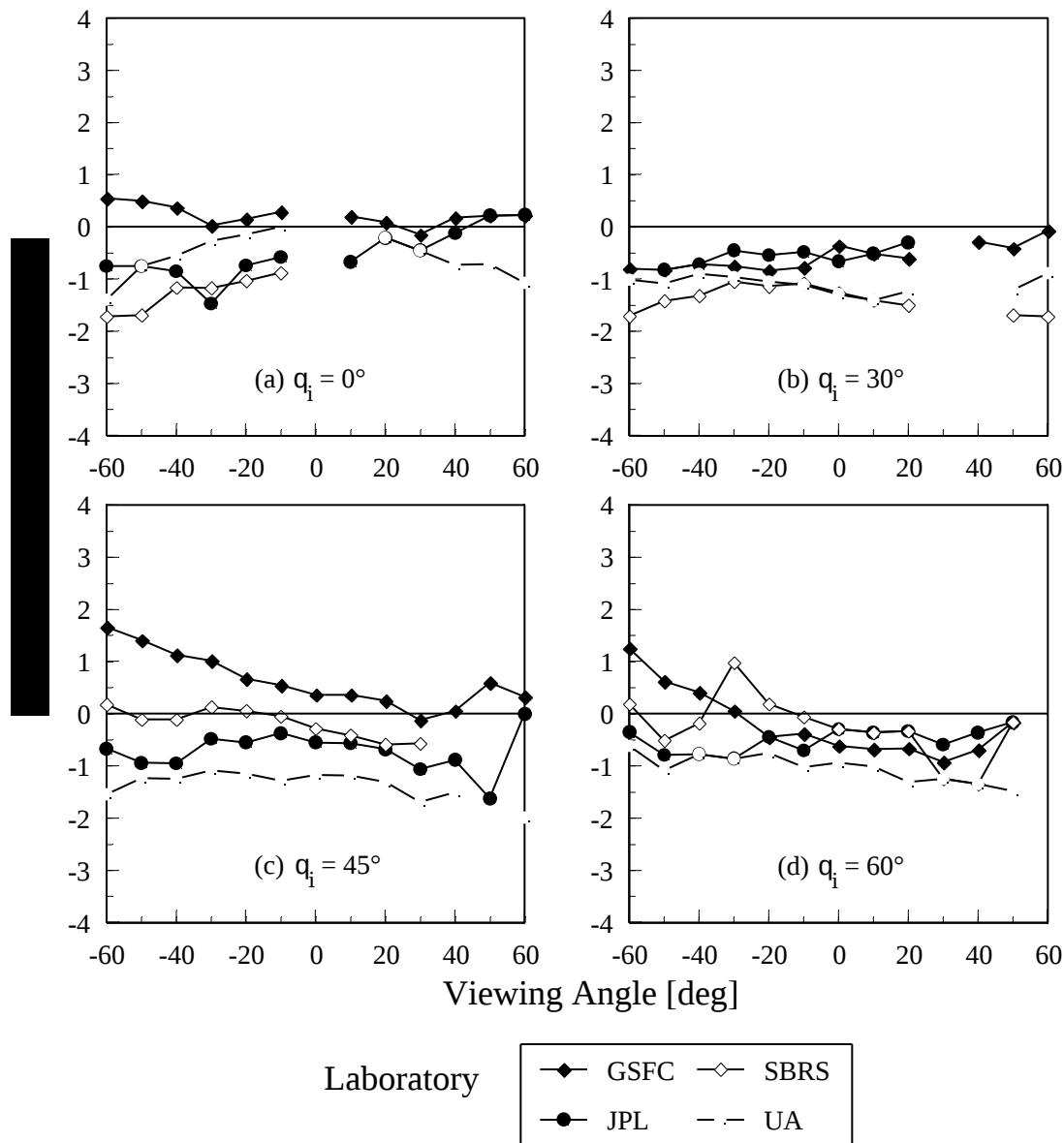


Figure 11. Difference of BRDF relative to the NIST value as a function of viewing angle for Spectralon at 633 nm. The incident angles are indicated in the panels, and the participants are given in the legend. From Early, et.al., J. A. O. Tech, **17**, 1077-1091 (2000)

Figure 11

Spectralon,  $\lambda = 860 \text{ nm}$

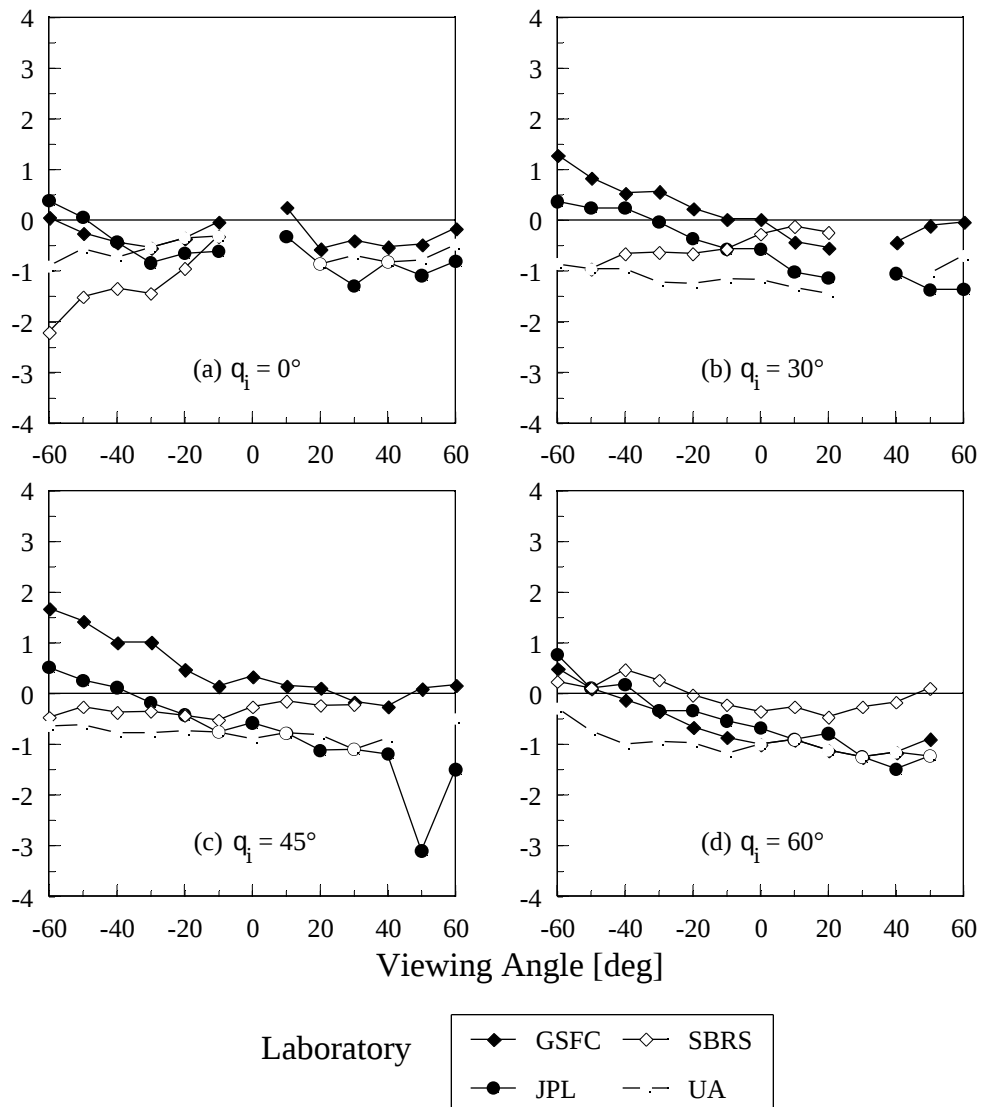
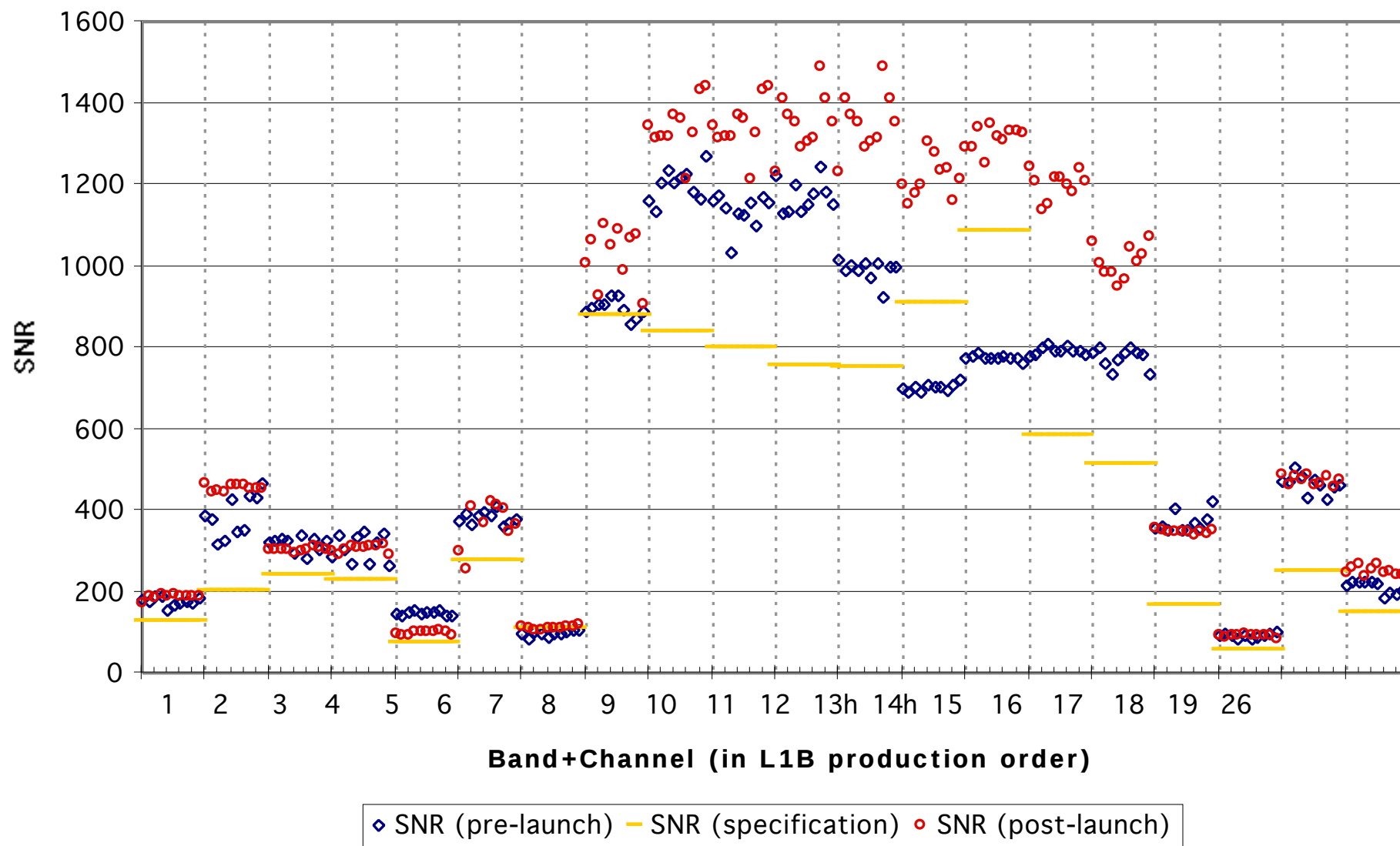


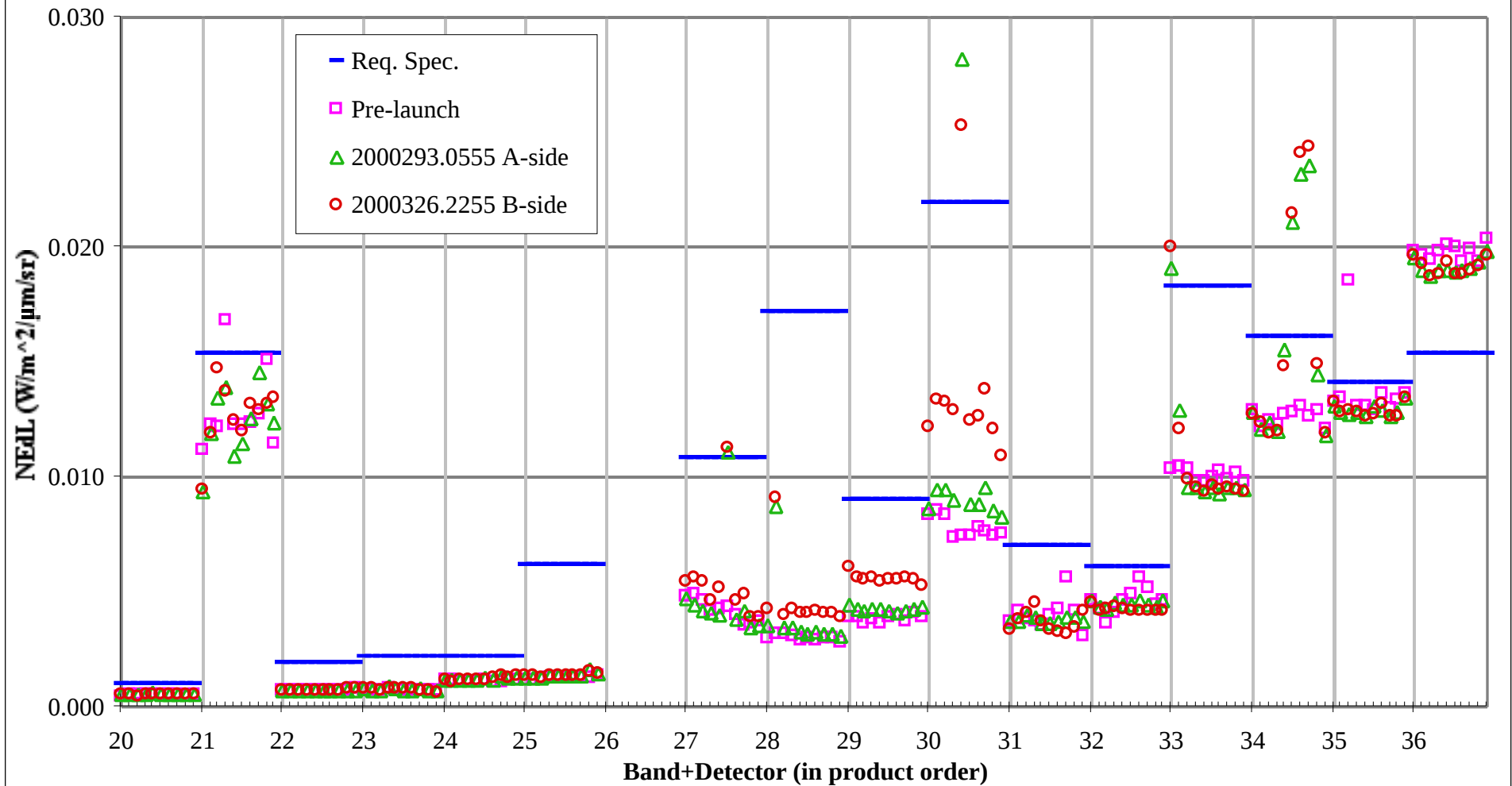
Figure 11. Difference of BRDF relative to the NIST value as a function of viewing angle for Spectralon at 633 nm. The incident angles are indicated in the panels, and the participants are given in the legend. From Early, et.al., J. A. O. Tech, **17**, 1077-1091 (2000)

Figure 12

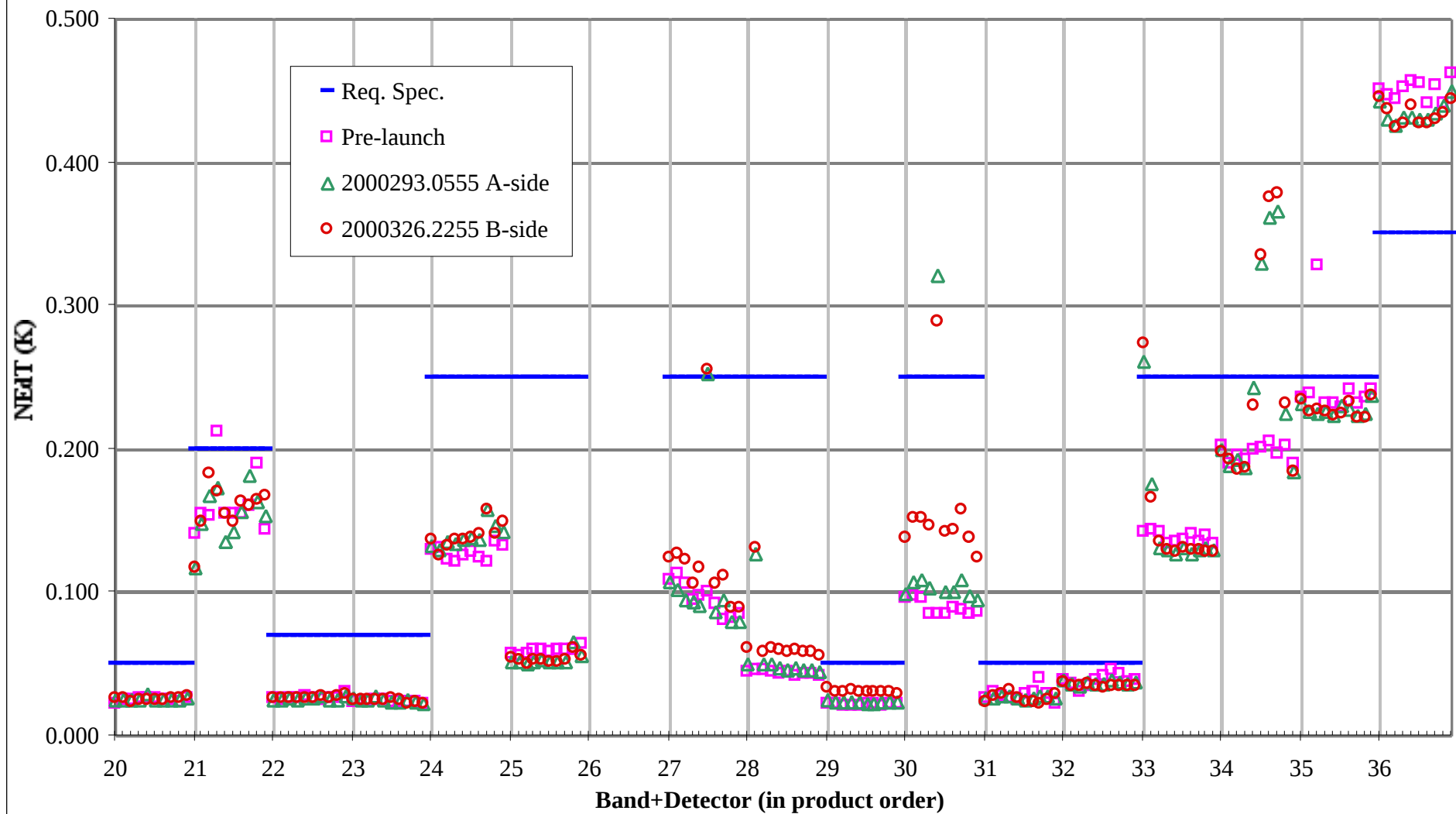
Figure 1. MODIS RSB SNR from Pre-launch, Post-launch and Specification at Ltyp



MODIS PFM TEB On-orbit **NEdL** (@ Ltyp) vs. Pre-launch and Specification



MODIS PFM TEB On-orbit **NEdT** (@ Ltyp) vs. Pre-launch and Specification



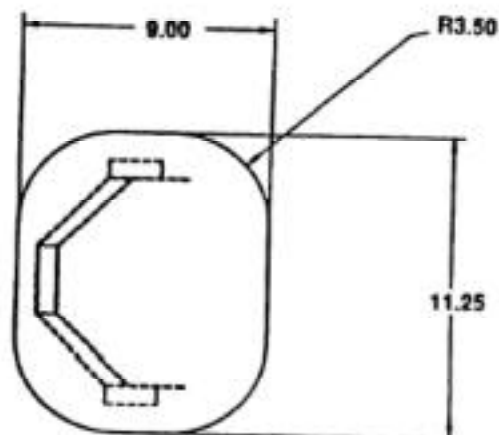
# RADIOMETRIC CHARTS - Section 3



# MODIS SOLAR DIFFUSER AND SOLAR VIEW DOOR ASSEMBLIES

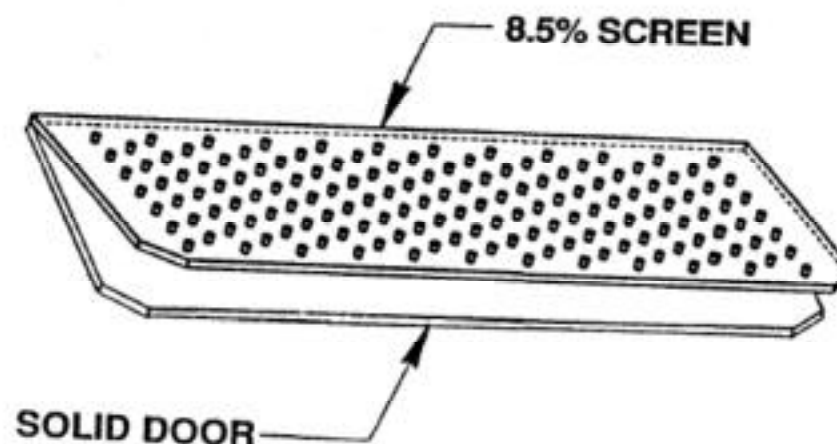
**HUGHES**

SANTA BARBARA RESEARCH CENTER  
a subsidiary



- BRDF: 0.315 sr
- EFFECTIVE ALBEDO: 53%

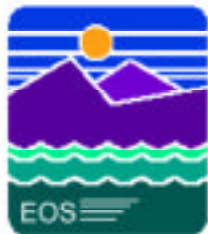
**SOLAR DIFFUSER  
ASSEMBLY**



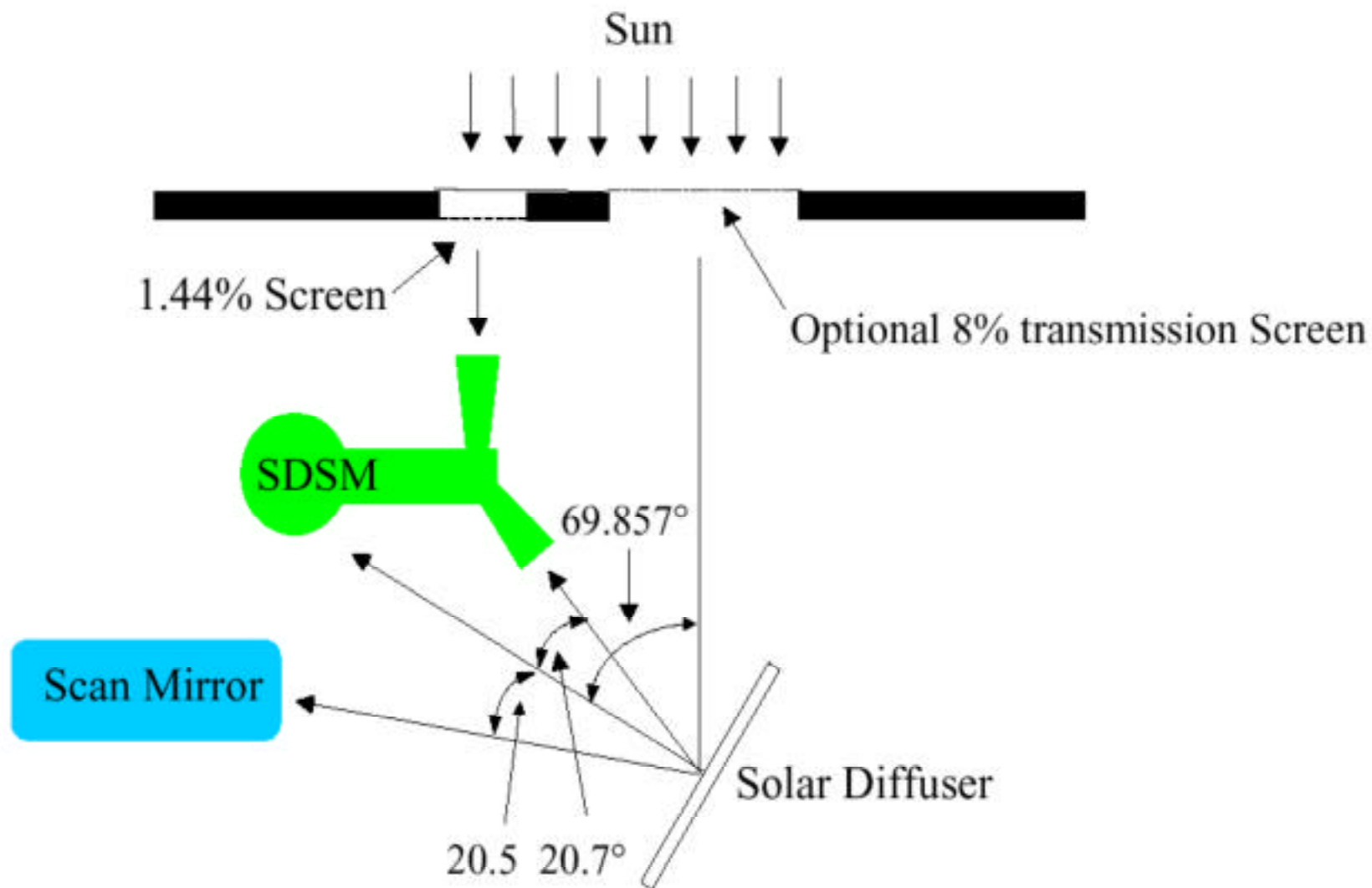
**SOLAR VIEW DOOR  
ASSEMBLY**

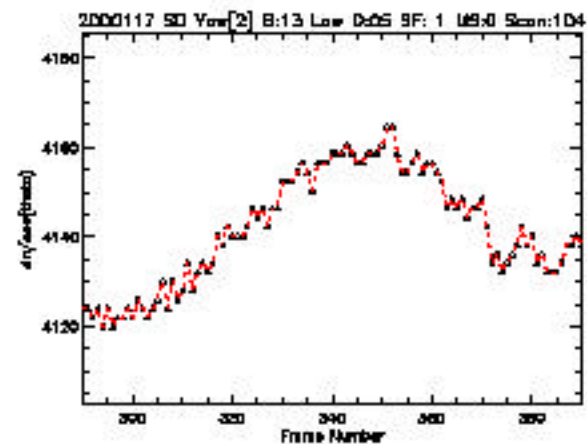
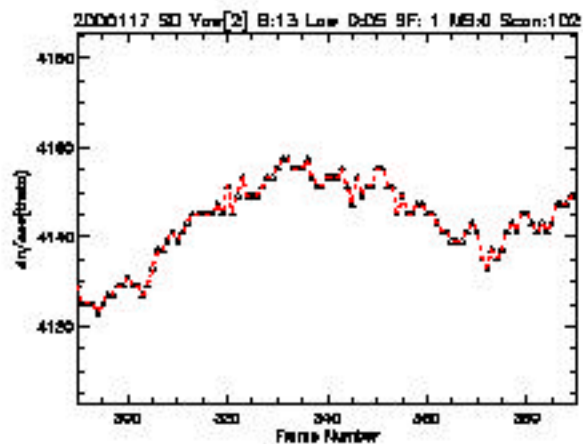
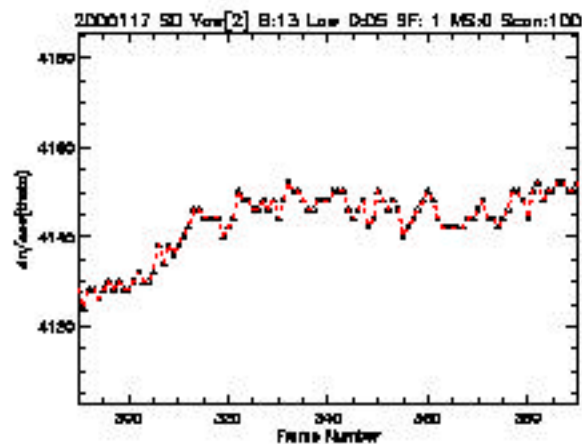
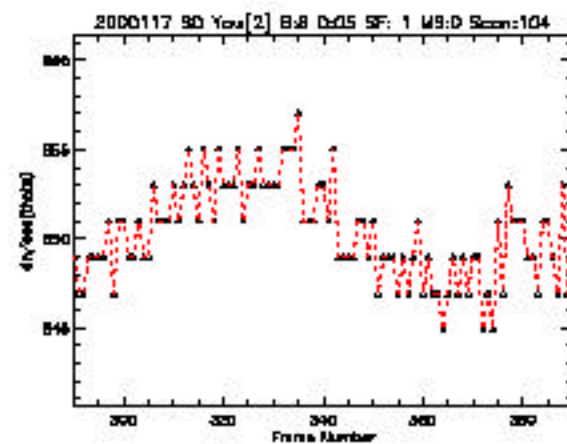
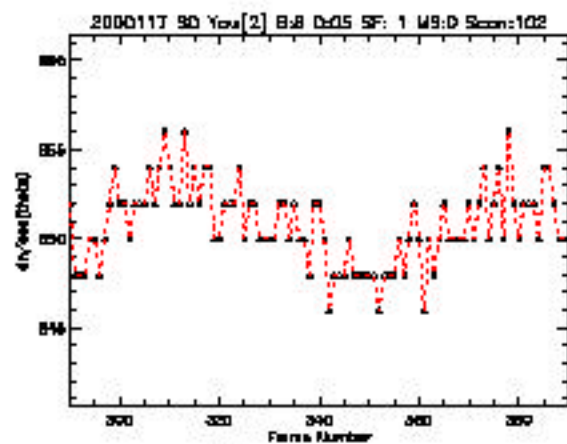
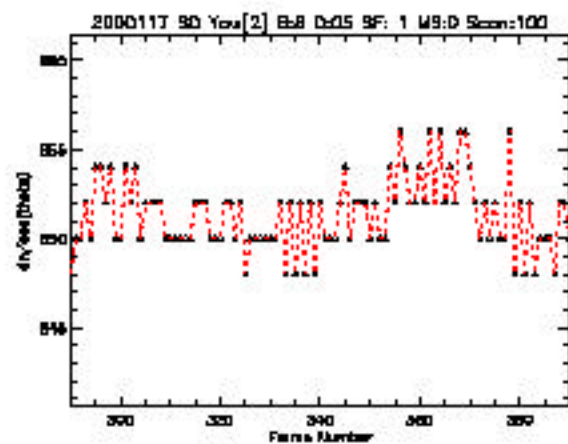
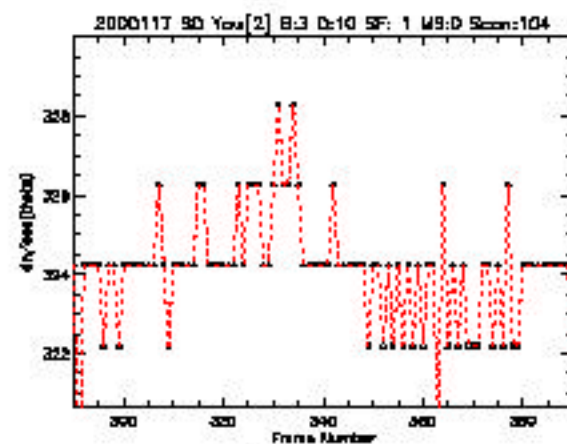
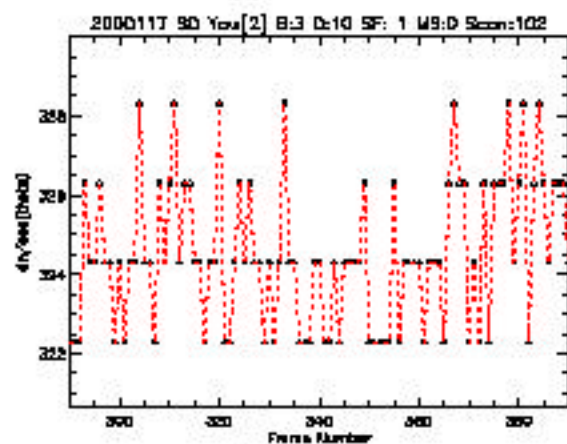
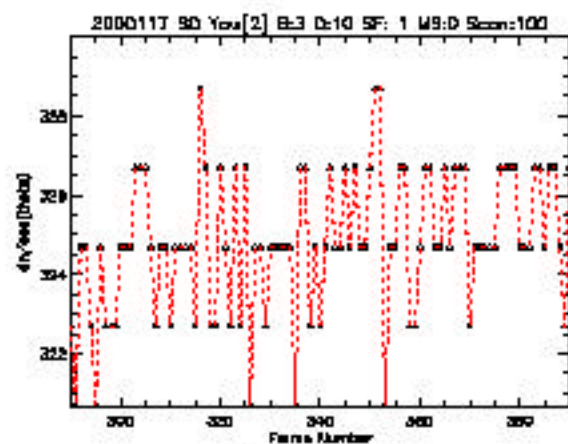
- DIFFUSER MATERIAL IS SPECTRALON (PTFE)

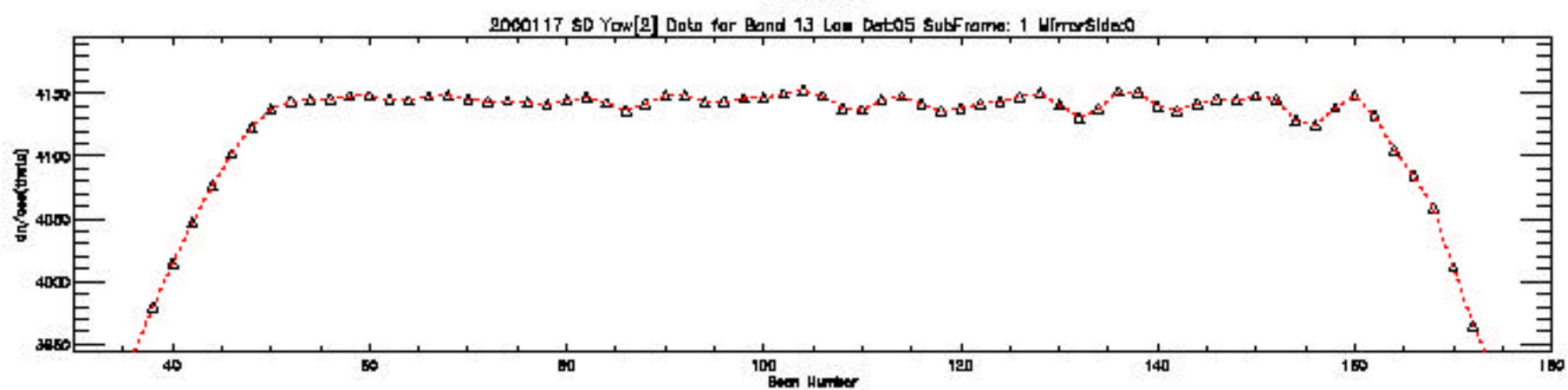
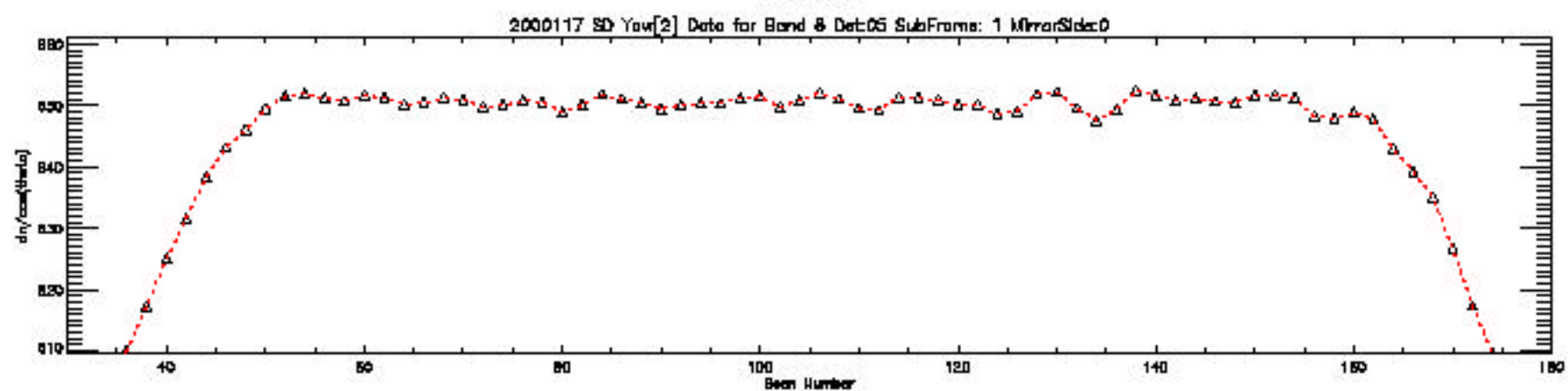
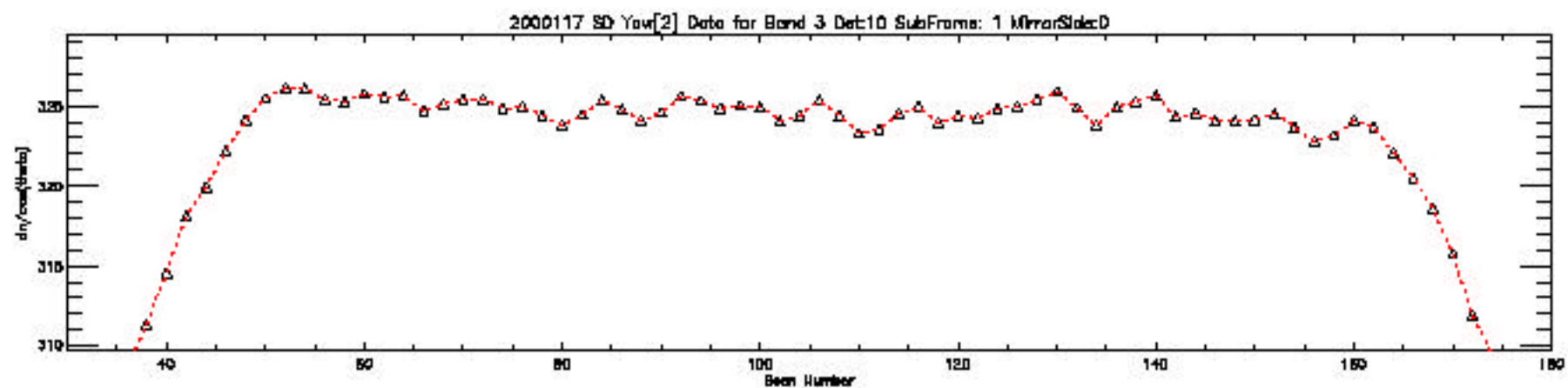
4/93  
93-0191-146



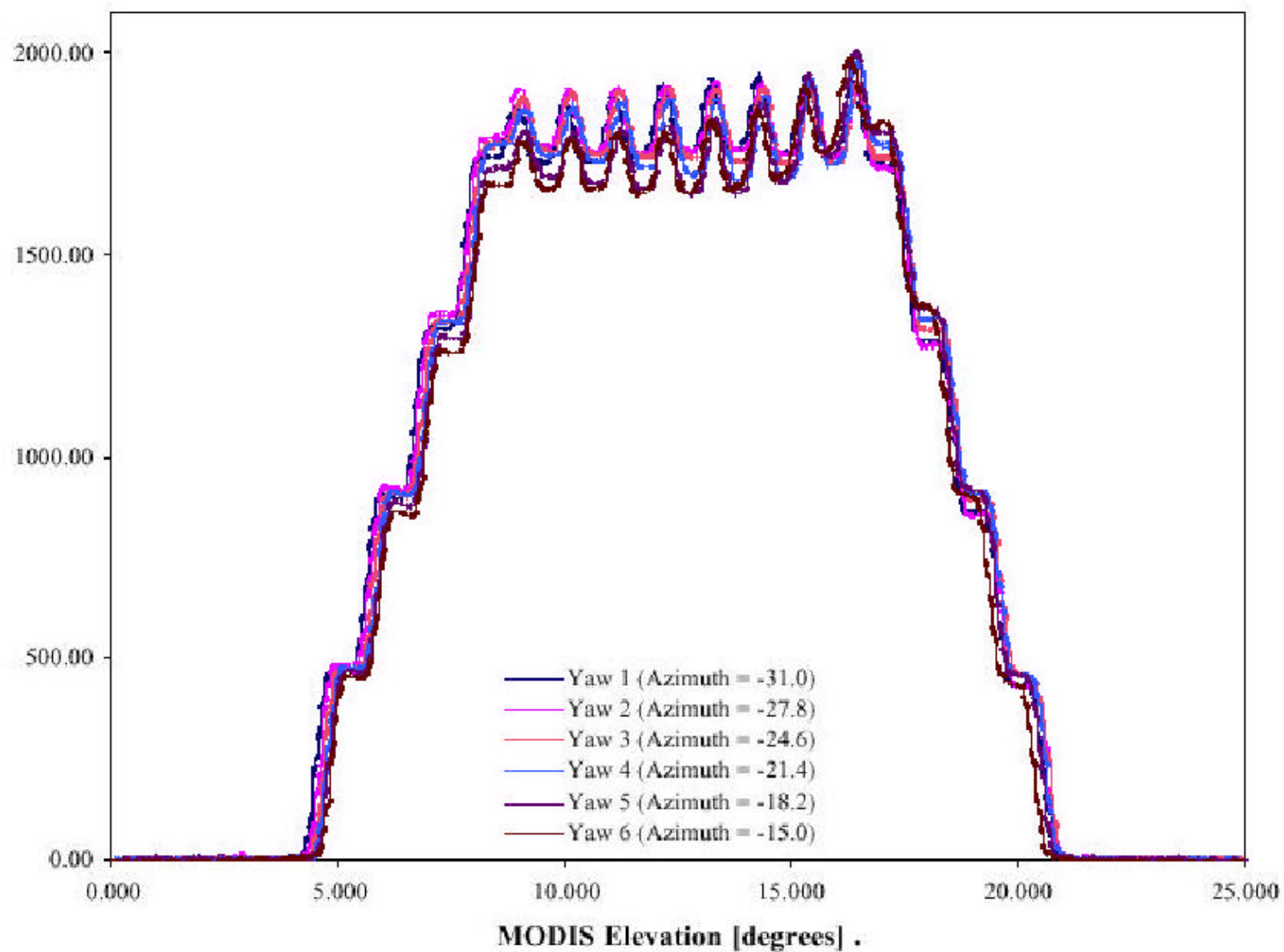
# Solar Diffuser Stability Monitor (SDSM)



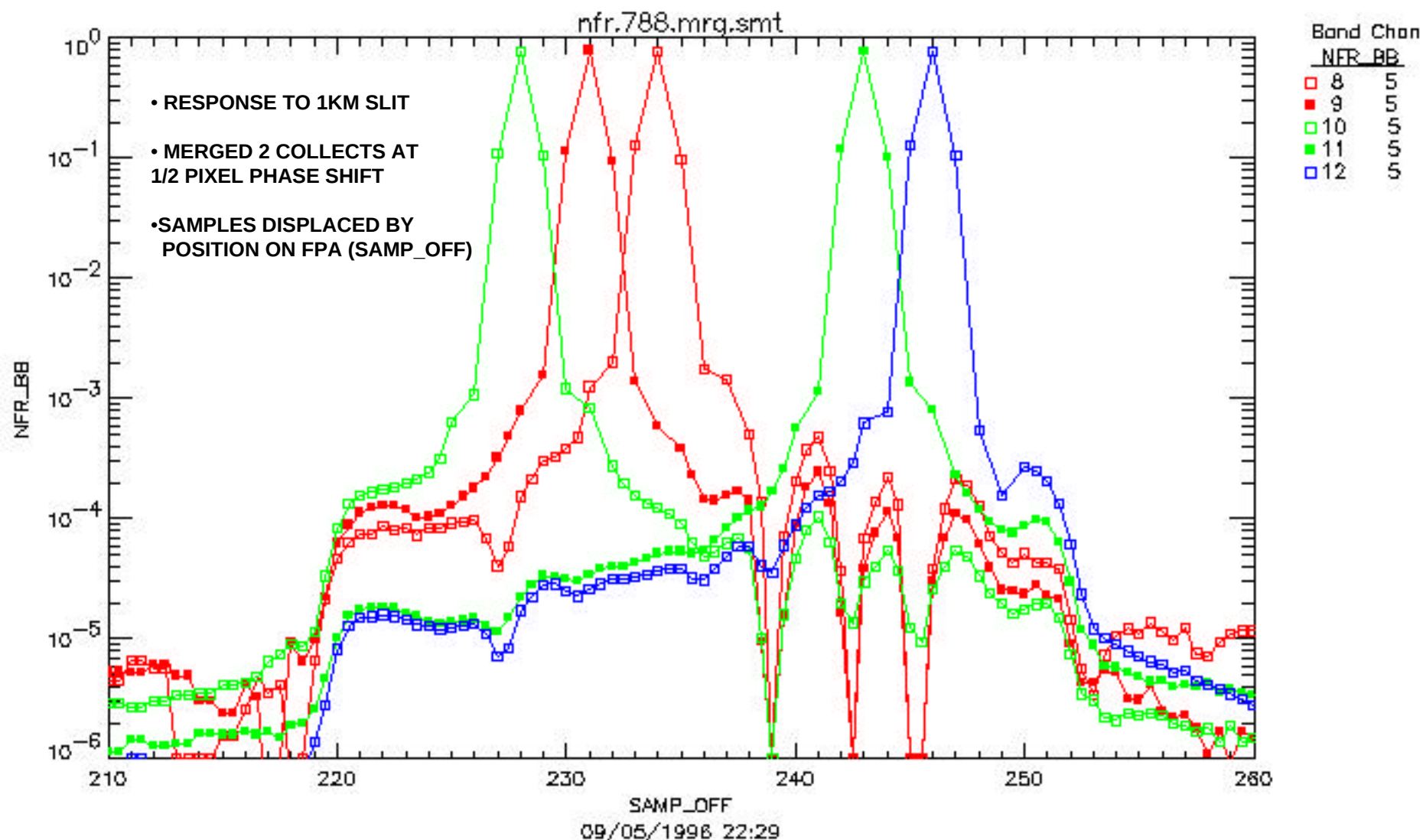




**SDSM Sun View .**  
**Detector 1**

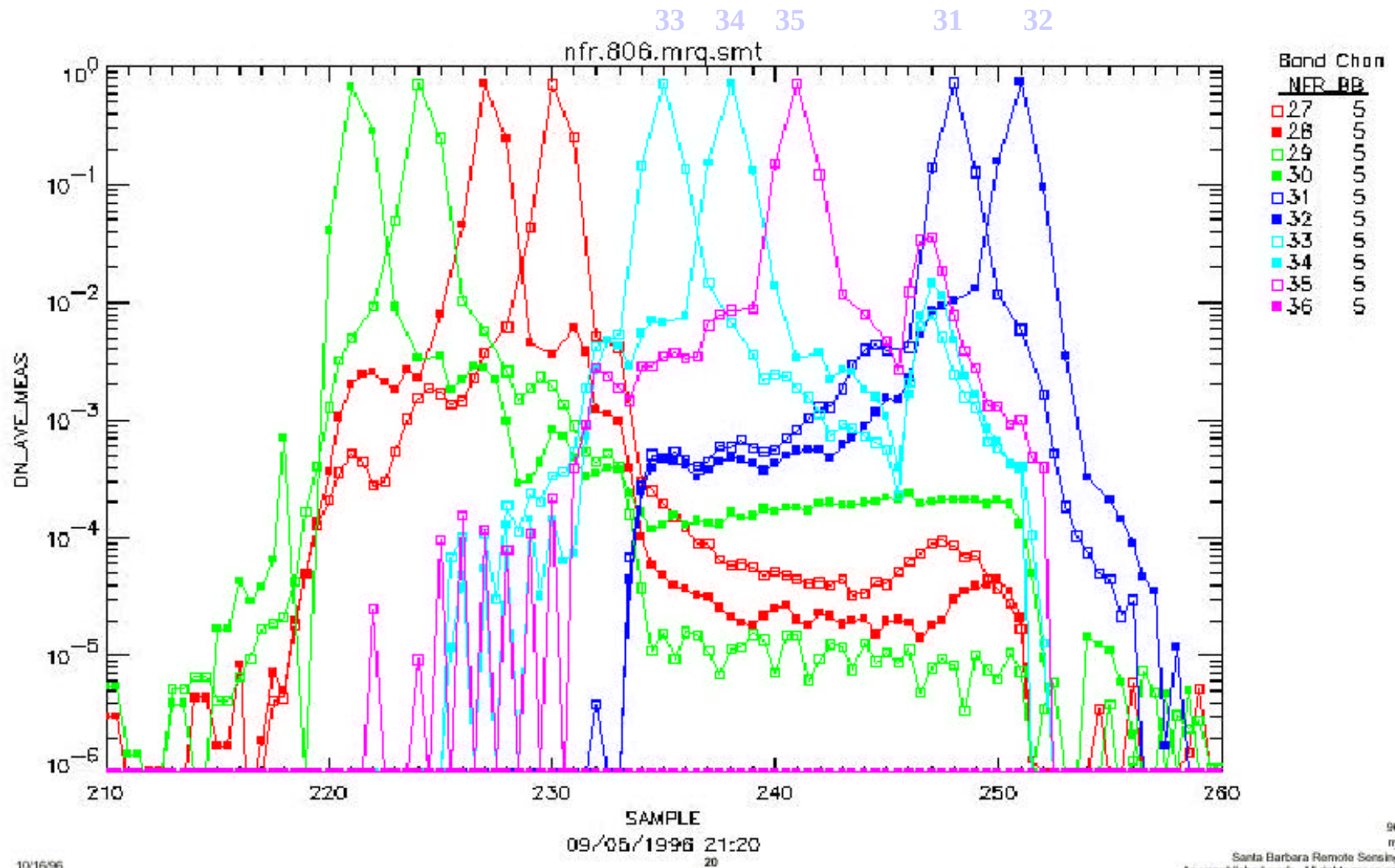
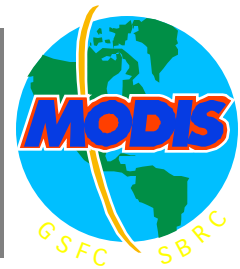


# VIS 1 KM LOG SCALE

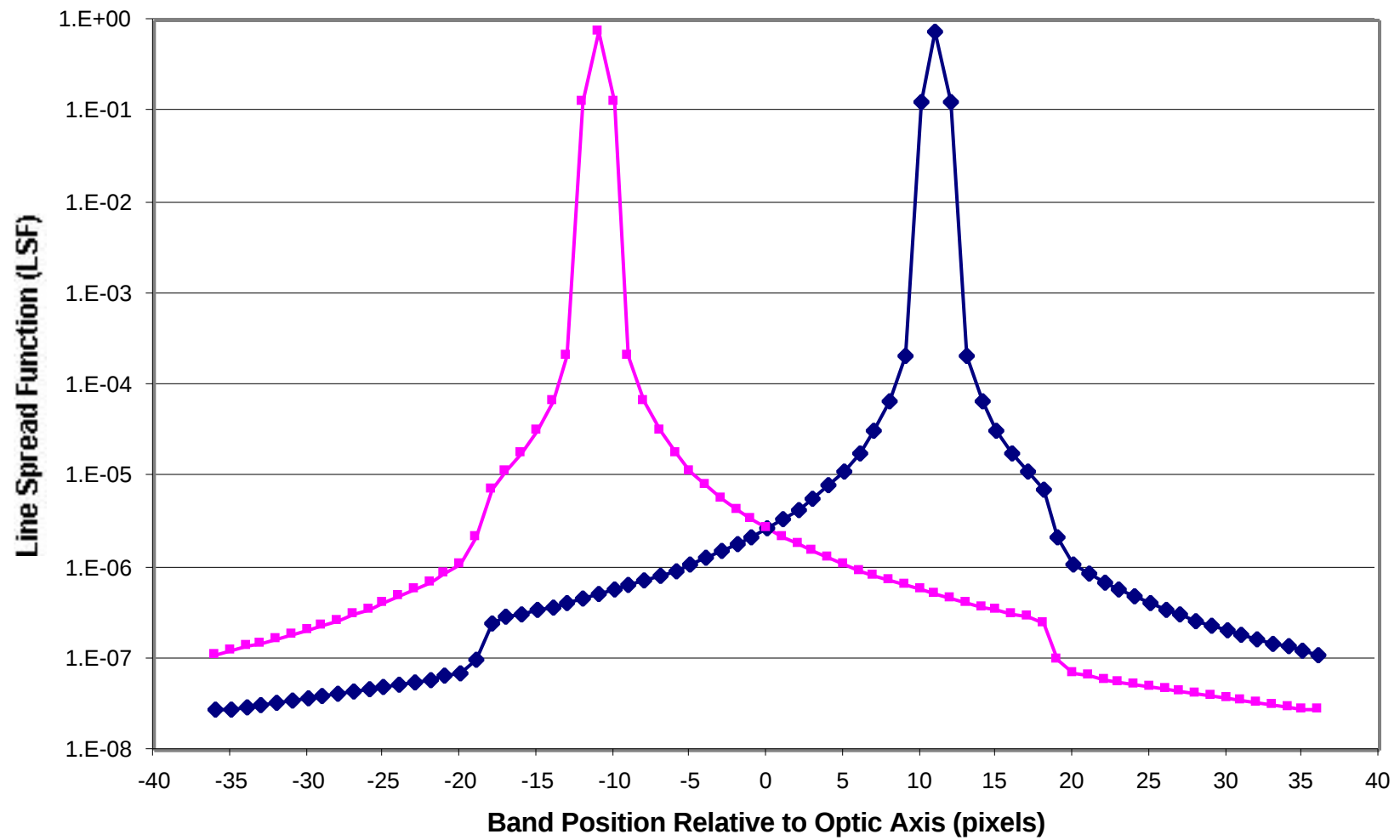




# Pre-launch Crosstalk Observation (Near Field Response Testing)

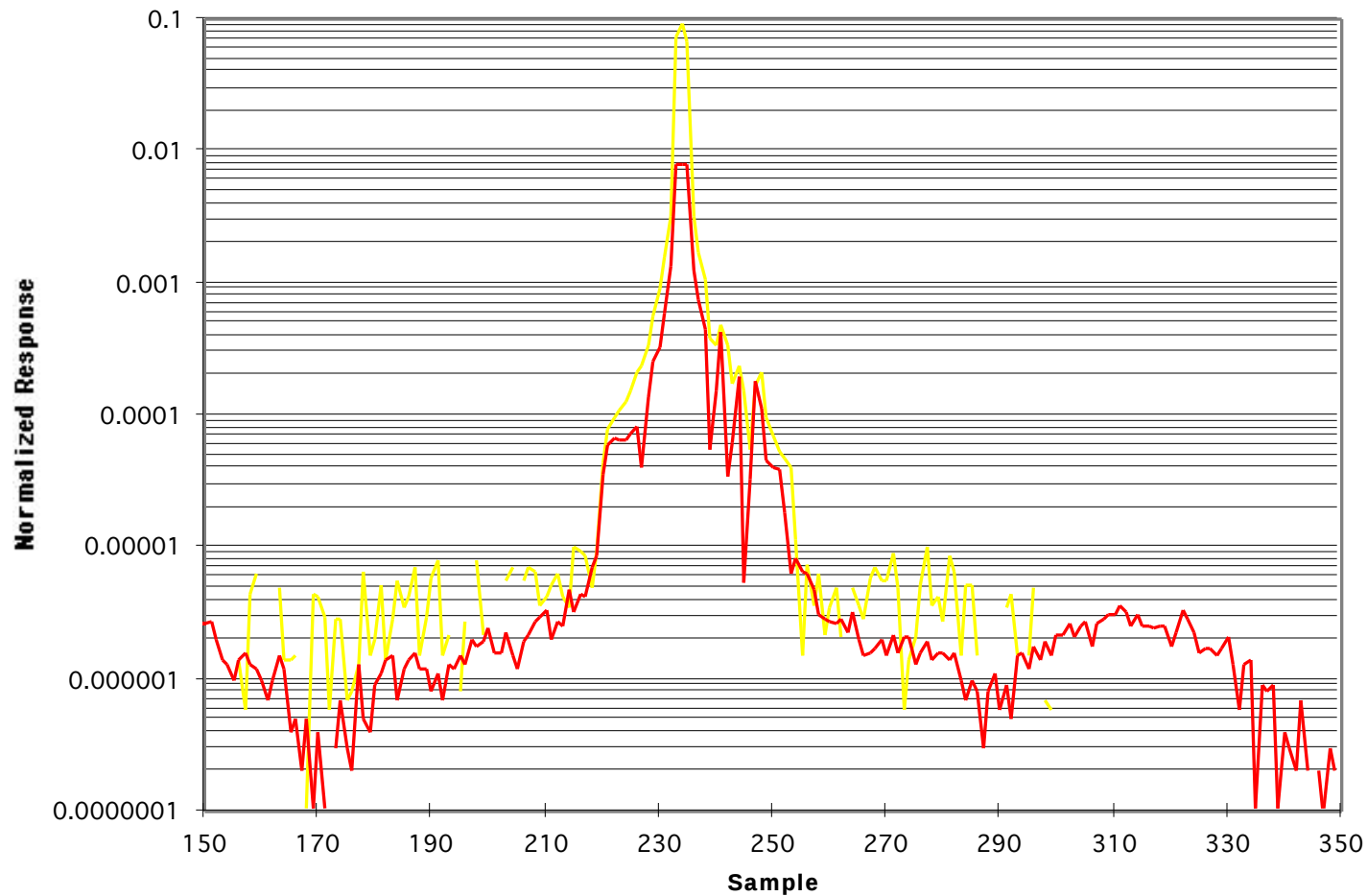


**Figure 5.5-1: Line Spread Functions for two identical bands located on far-left and far-right of FPA**



# PFM RESPONSE CONSIDERABLY BETTER THAN EM

**Comparison of Band 8 EM and PFM NFR**

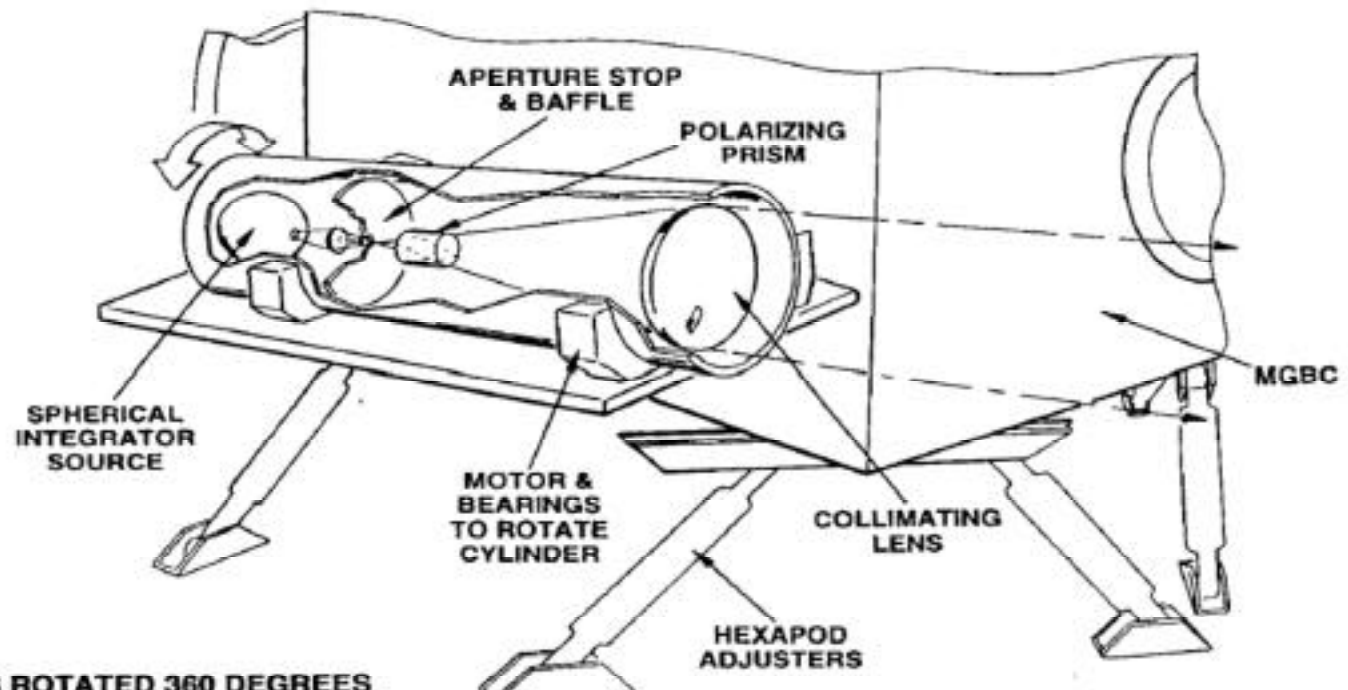




## MODIS POLARIZANCE IS CHARACTERIZED WITH A STAND-ALONE POLARIZER SUBSYSTEM

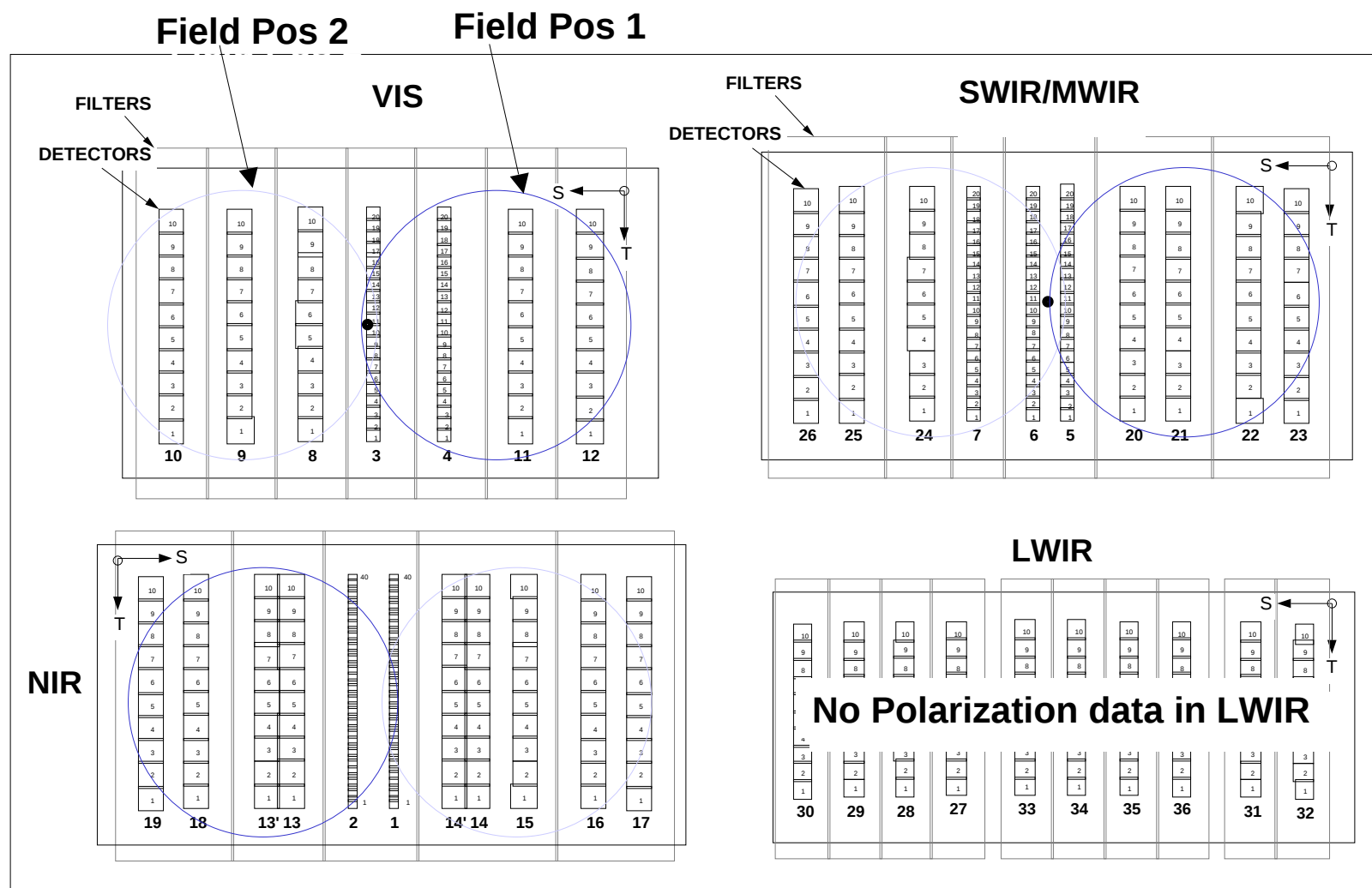
**HUGHES**

SANTA BARBARA RESEARCH CENTER  
a subsidiary

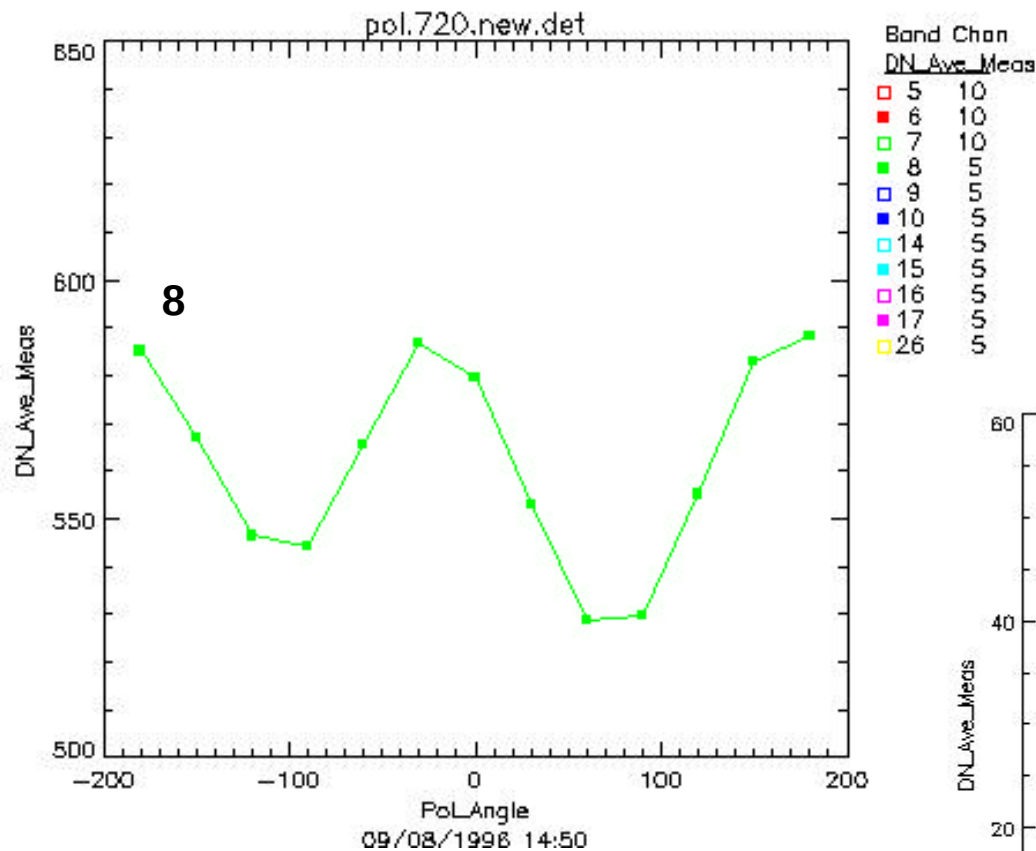


- CYLINDER IS ROTATED 360 DEGREES

# DATA ACQUIRED IN 2 FIELD POSITIONS

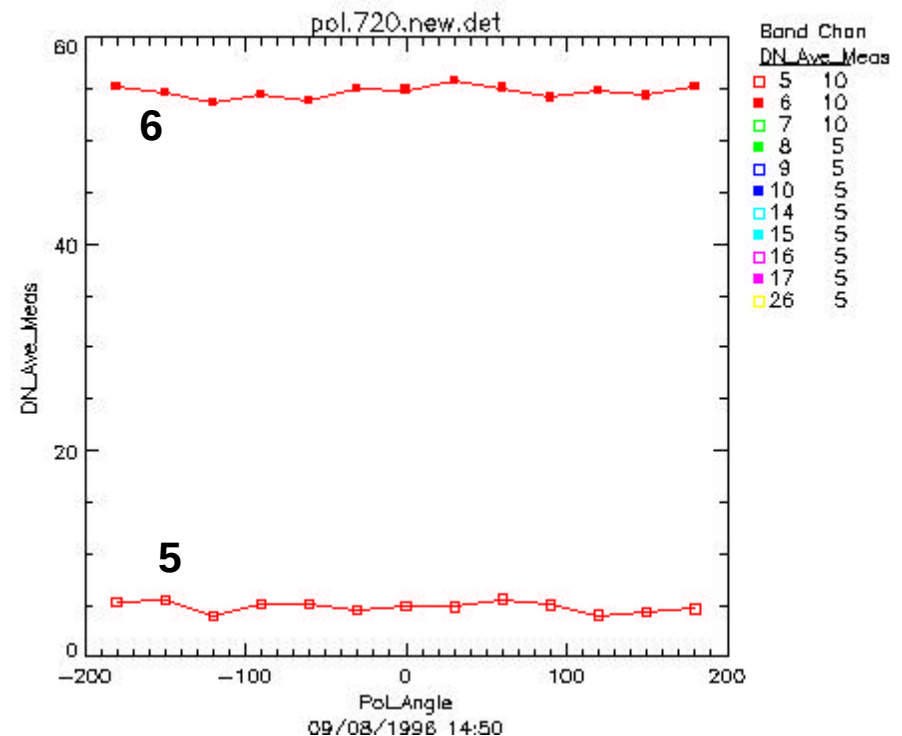


# MODIS RESPONSE TO FULLY POLARIZED BEAM BANDS 3, 6 - 10, 14-17, 26 UAID 720, 721, ZERO DEGREE AOI (Cont'd)

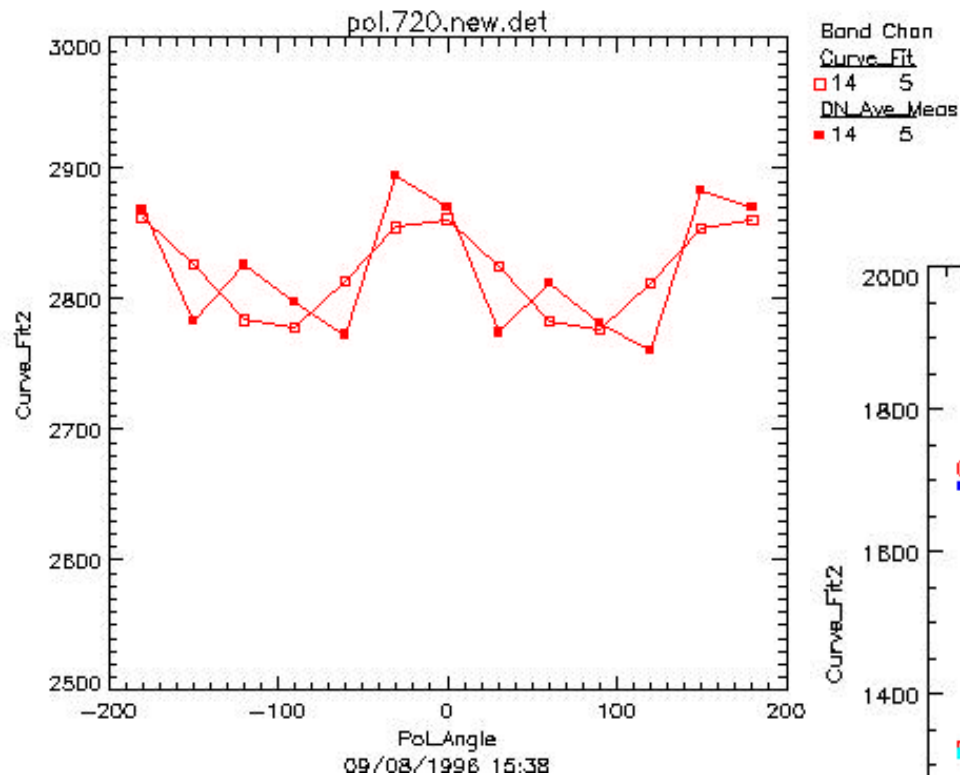


- LARGEST POLARIZATION IN B8

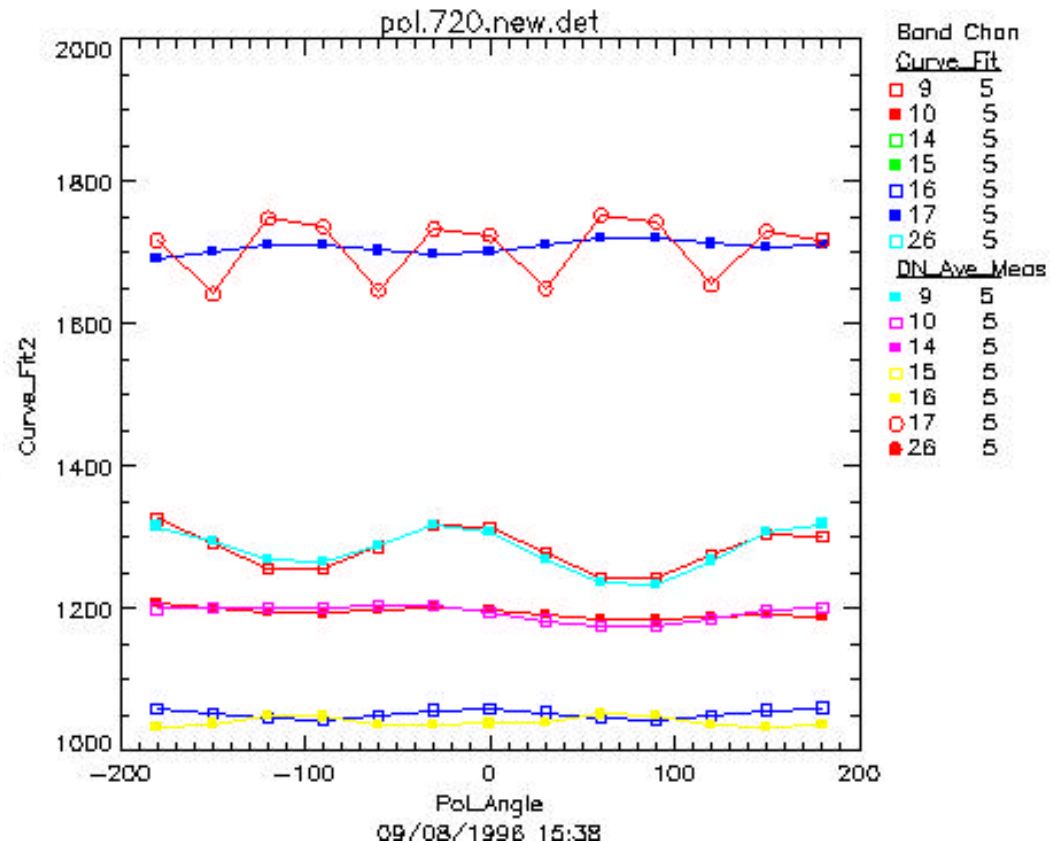
- LOW BUT RETRIEVABLE POLARIZATION IN BAND 6
- BAND 5 SIGNAL TOO LOW



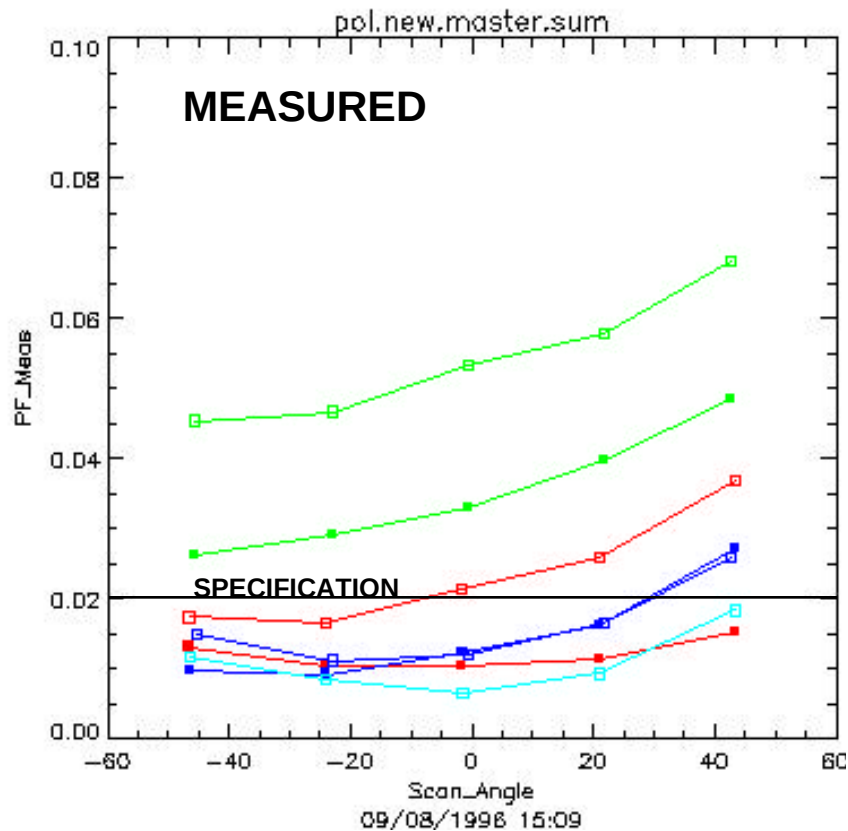
# CURVE FIT EXTRACTS COS2q COMPONENT FROM SIGNAL OBSERVED



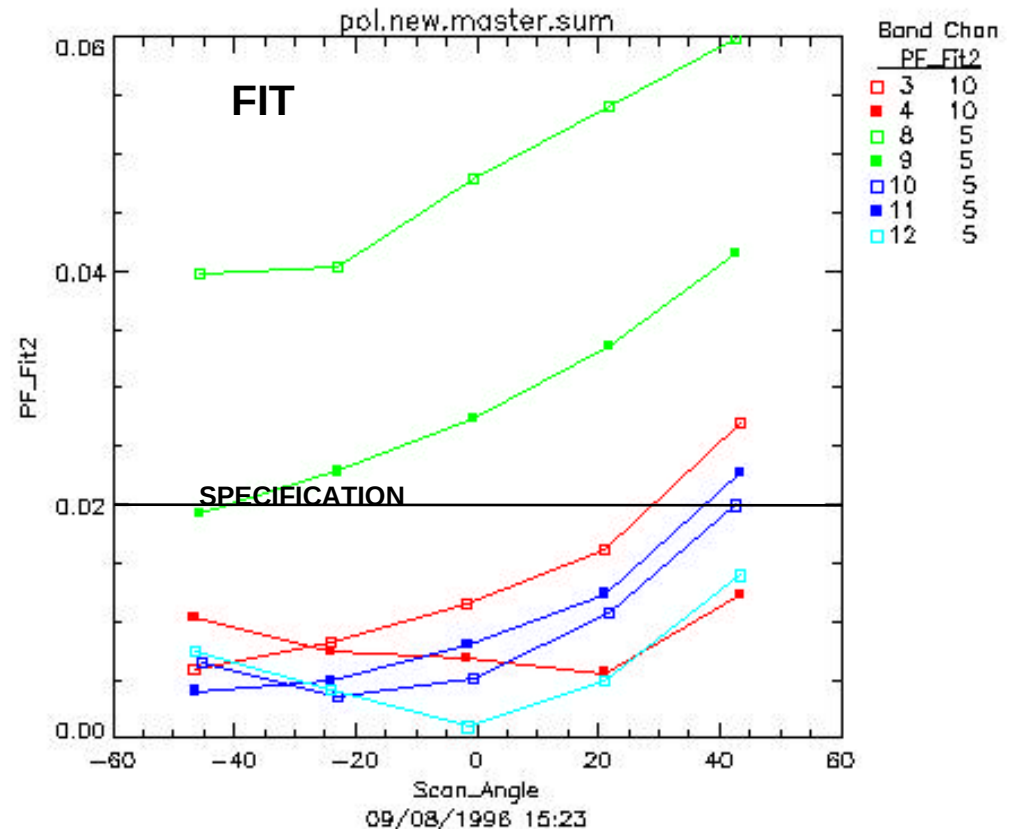
- RATIONALE FOR CURVE FIT TO BE DISCUSSED LATER IN ADVANCED TOPICS SECTION



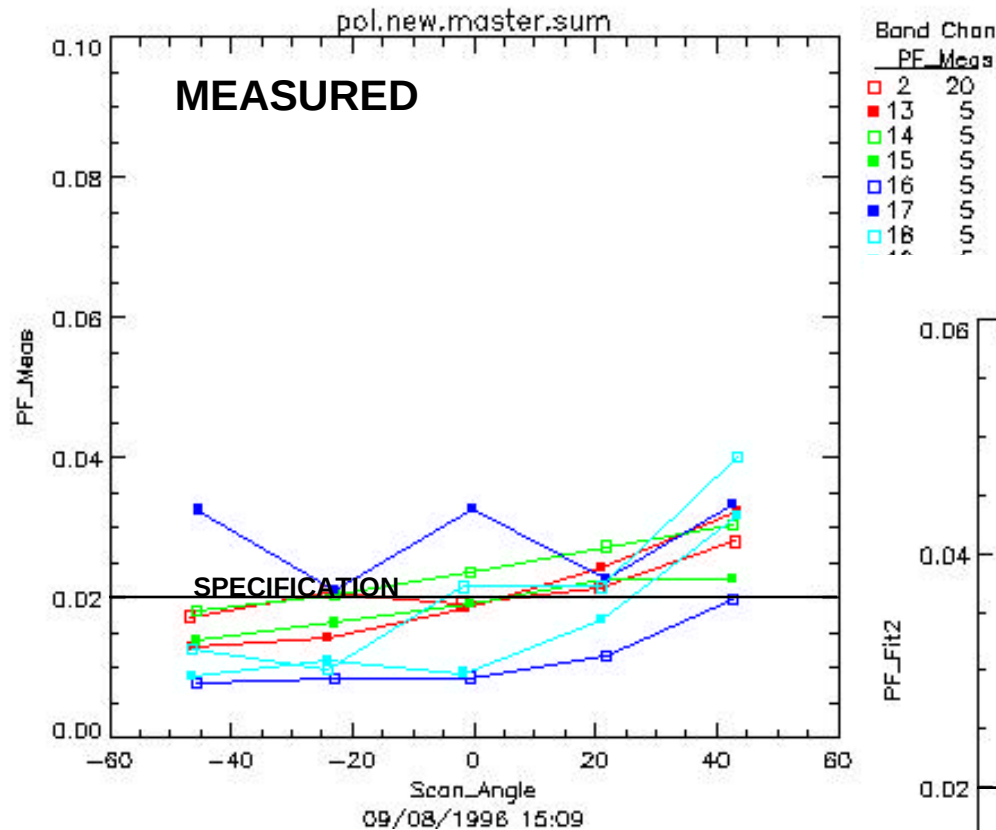
# RESULTS: VIS MEASURED AND FIT POLARIZATION FACTOR



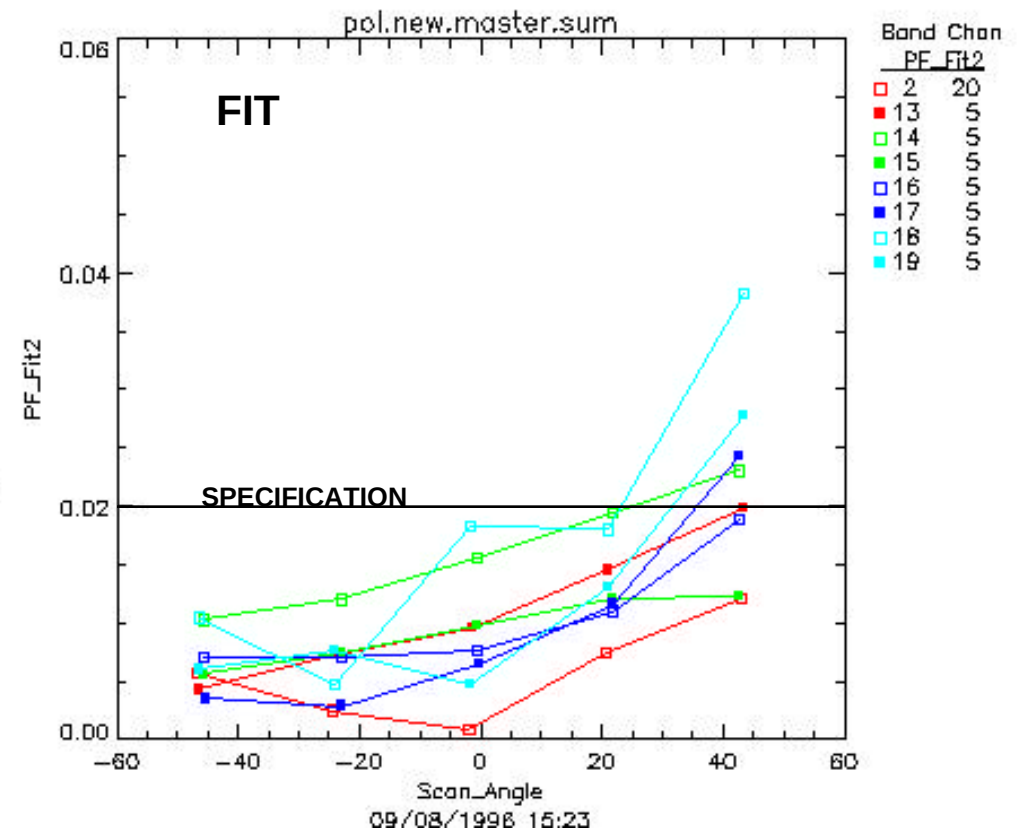
- NO SPEC ON BAND 8
- BAND 9 FAILS AT ALL SCAN ANGLES
- BANDS 3 AND 11 FAIL AT 45°



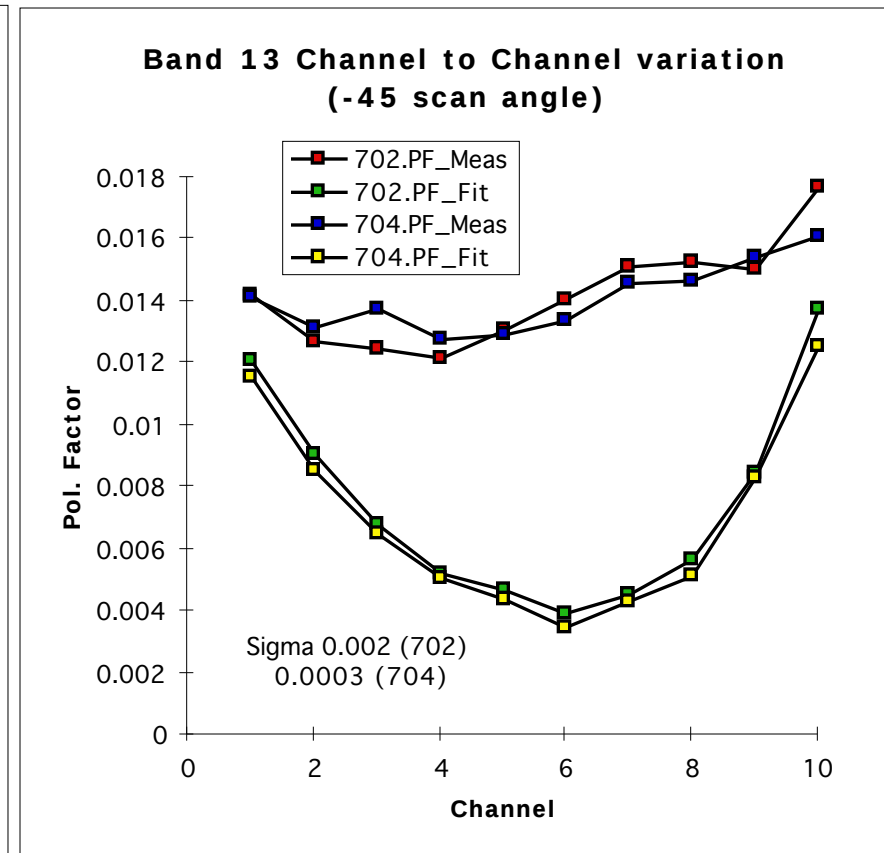
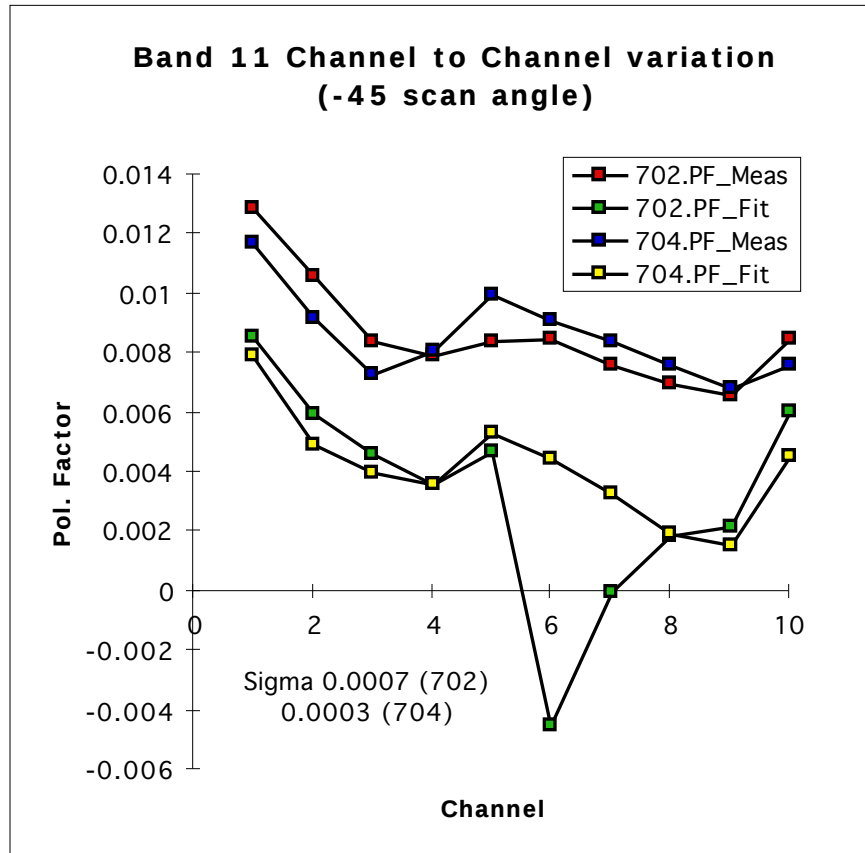
# RESULTS: NIR MEASURED AND FIT POLARIZATION FACTOR



- BANDS 14, 17, 18 AND 19 FAIL AT 45°



# REPEATABILITY ANALYSIS PERFORMED ON EARLY DATA SETS



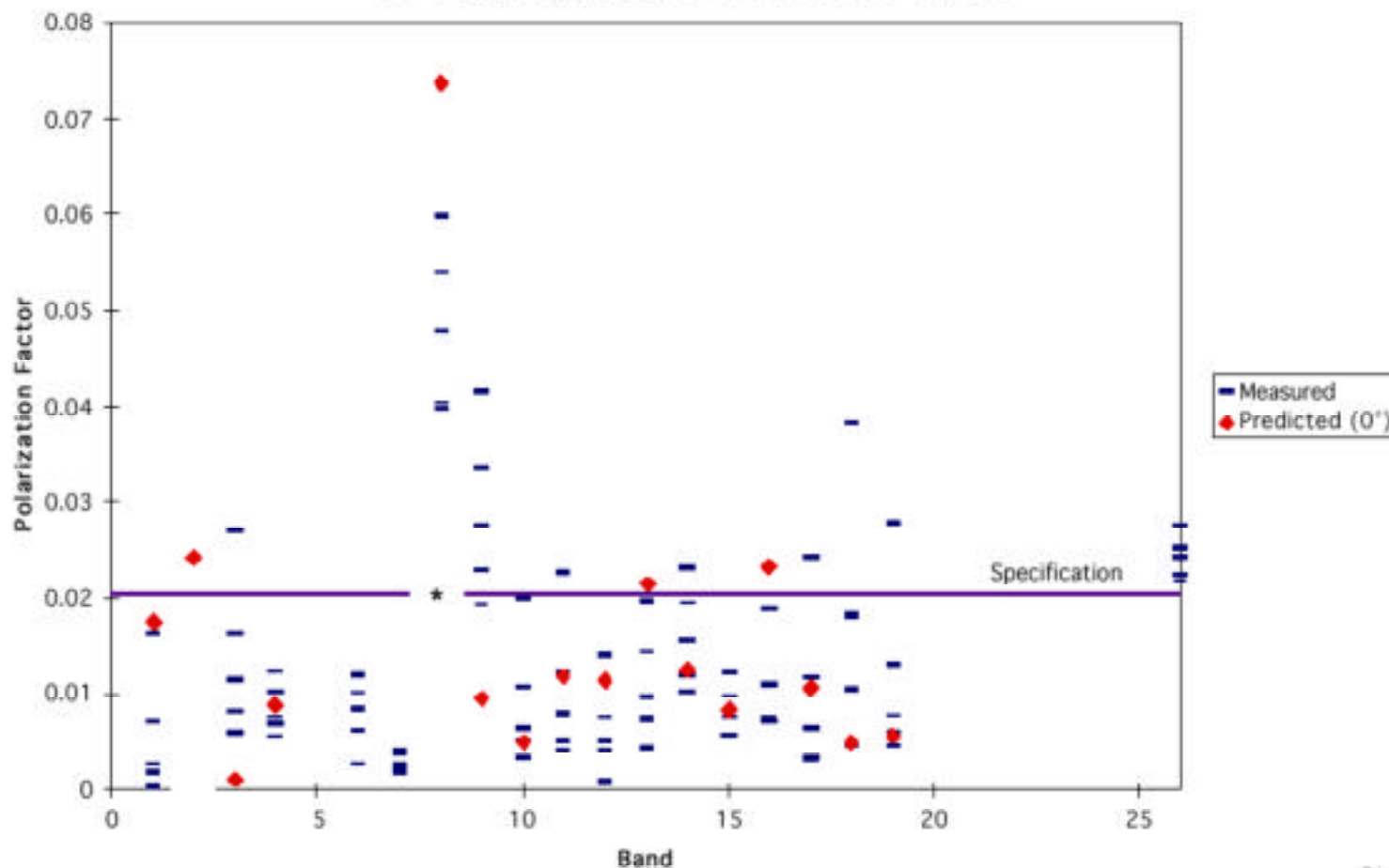
- TWO DATA SETS, 702 AND 704 SEPARATED BY TWO DAYS
- ANALYSIS PERFORMED BY E. KNIGHT



### 3.3.5. Waiver 055: Polarization



Polarization Factor for the MODIS Bands  
for Scan Angles of  $\pm 45^\circ$ ,  $\pm 22.5^\circ$  and  $0^\circ$

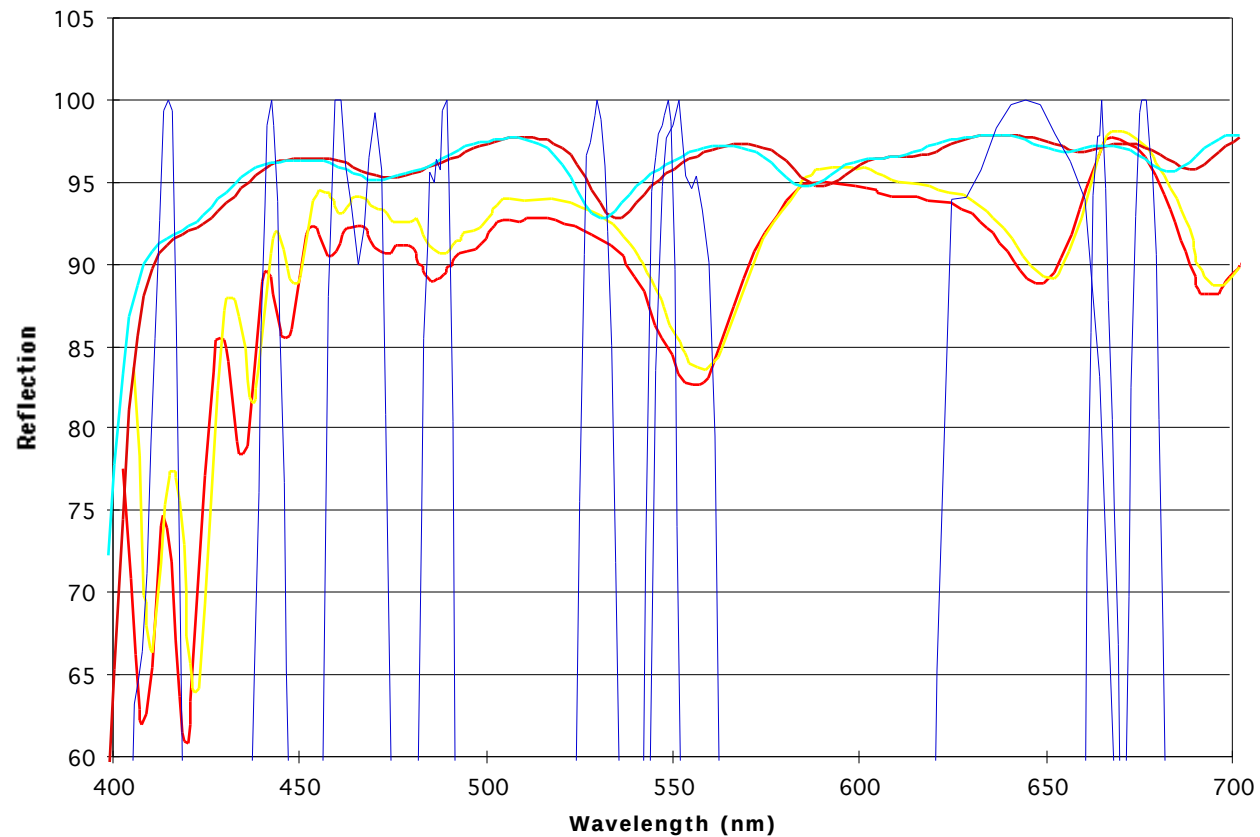


\*Note: Low residual polarization desired for band 8, but not covered as a specification requirement

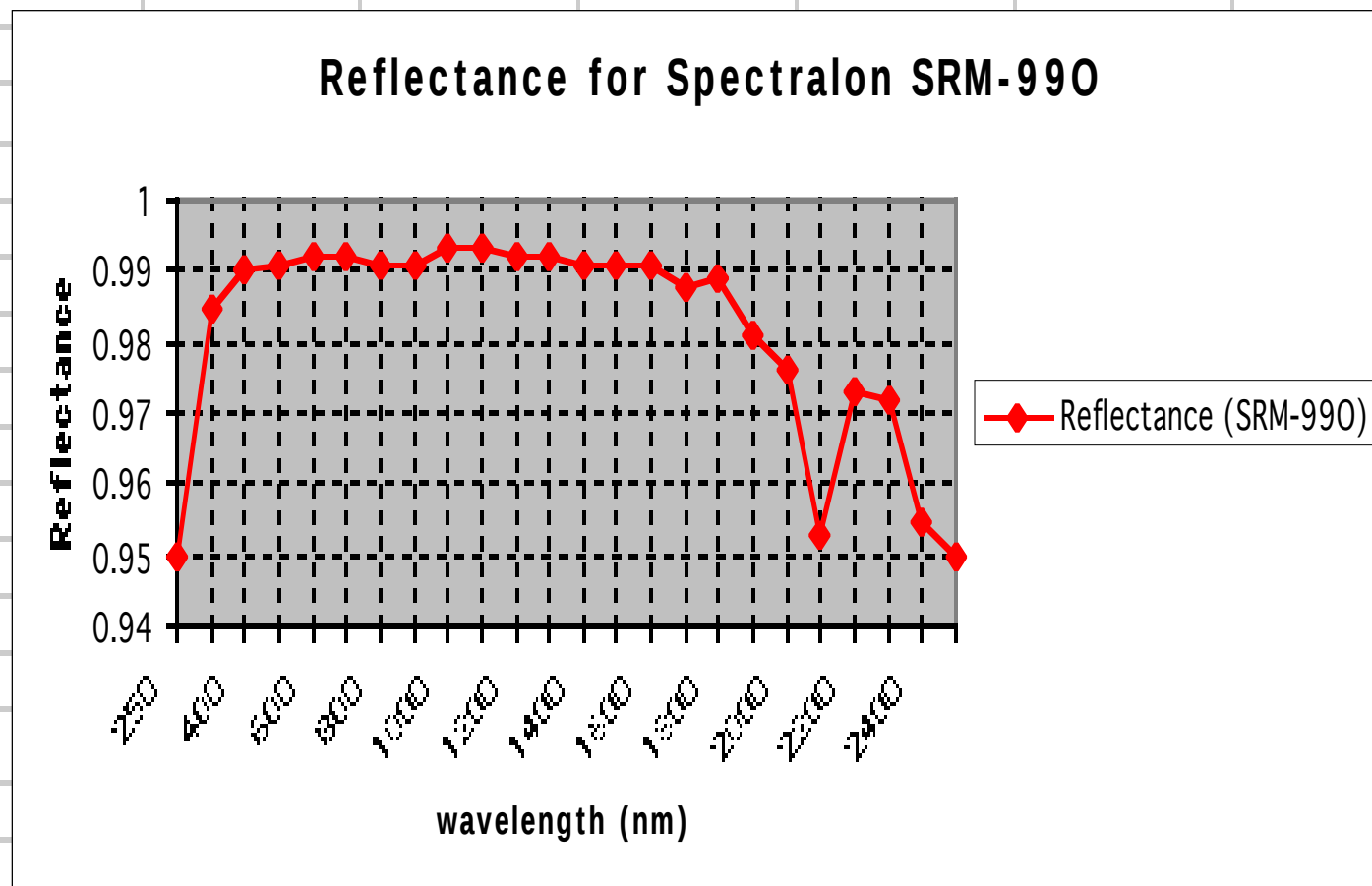
# OFC DBS USED IN PFM HAS MORE POLARIZATION

- OFC DICHOIC SELECTED FOR LOWER SCATTER PROPERTIES

COMPARISON OF OFC AND OCLI DBS #1



# Reflectance of MODIS Solar Diffuser Material





- Review of TV3 data uncovered non-responsive detectors on band 6, and 1 non-responsive detector on band 5
- Of 20 channels on Band 6, during TV3
  - 6 are inoperable at 83K (channels 1,2,5,7,8,19)
  - 2 additional channels are inoperable at 88K (channels 4,9)
  - 3 additional channels are inoperable between 88K and 118K (11,16,17)
- Not present in TV2 test
- No direct method of analyzing effects other than in TV
  - Bands do not operate in ambient/bench test cooled environment
  - Band evaluation occurs only in a radiatively-cooled environment
- Indirect method developed recently (DC restore offsets/ECAL ramps) allows for operability assessment in ambient environment



- **PFM –vs- FM1 –vs- FM2 design, manufacturing, testing review**
  - Materials, processes, tests, test configurations, fabrication history/performance, yields, major events during fabrication and test cycle, personnel changes, equipment/facilities changes, etc
  - SWMIR FPA –vs- other Modis FPAs comparison
  - Susceptibility to failure mode
- **Process and production history review**
  - Other customers, performance history (before/after Modis), process maturity
- **Modis life/qualification test program review**
  - Sample/population size, tests performed/results, further test requirements/needs
- **Models/simulations**
  - Structural, thermal predictions, reliability analysis
- **FM2 FPA reliability/performance/operability assessment**
  - Available data
  - Additional testing

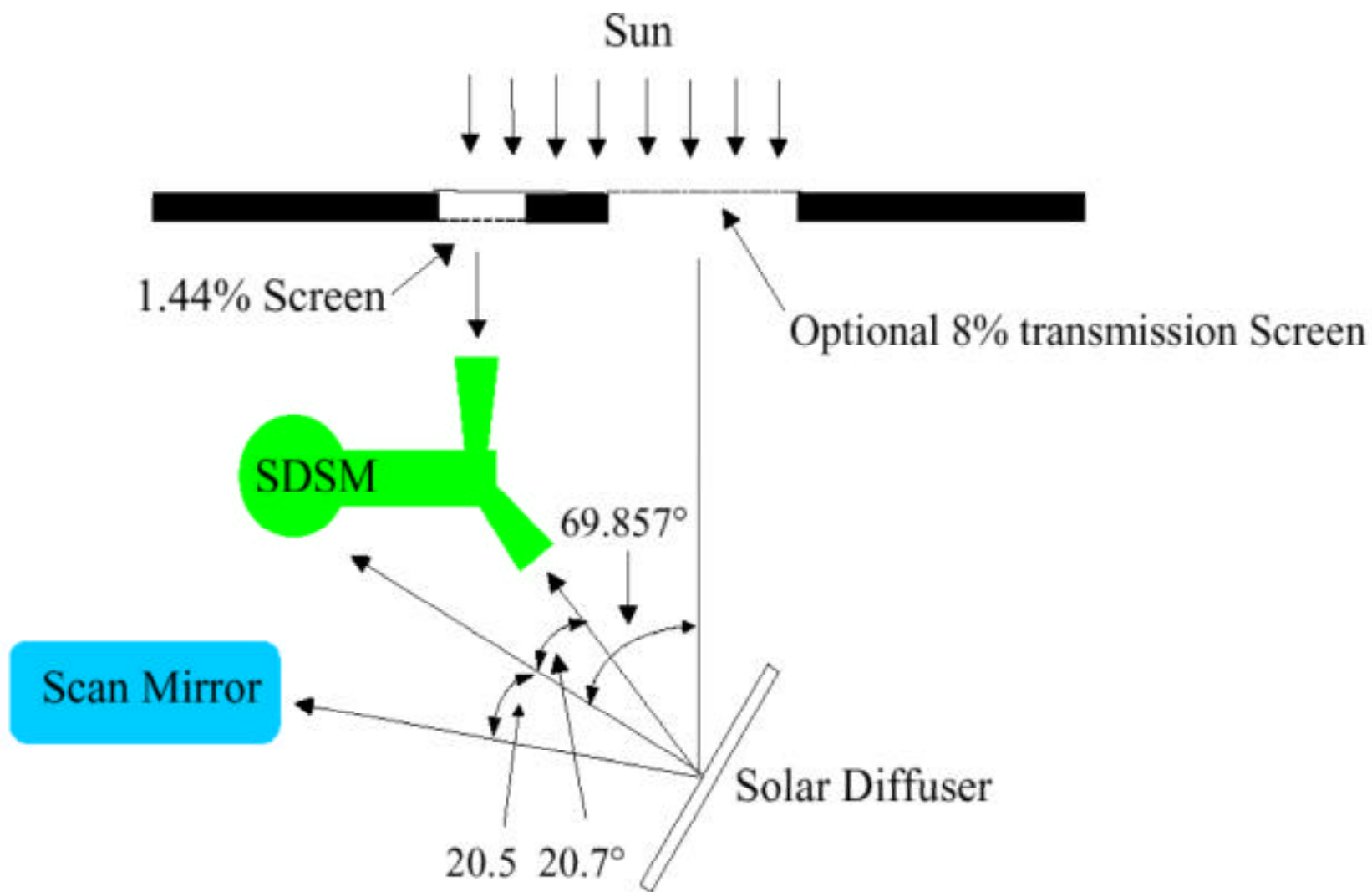
# FM1 SMWIR B5/B6 Operability Results to Date



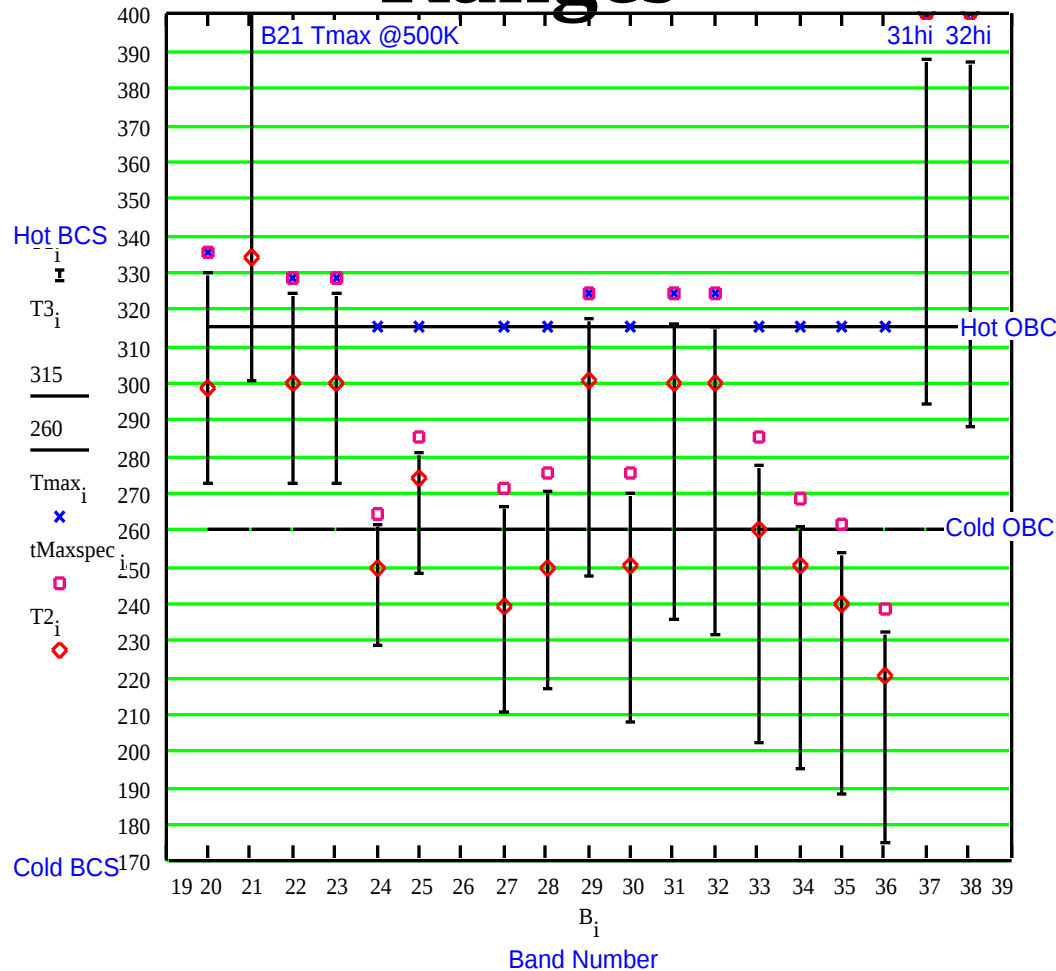
- **Anomaly most likely caused by deficiency of Indium bump interface**
  - **Between the detector and readout chip**
    - Electrical tests have supported this conclusion
  - **Manifestation of deficiency accelerated by thermal cycling**
    - FM1 test history review shows onset at about thermal cycle #70 (vs a 150 cycle spec/expectation)
  - **Some evidence of stress relief is apparent**
    - Phenomenon has been stable for 15+ thermal cycles
  - **Exact cause would require a complete disassembly/destructive diagnostic on the FPA**
    - Studies must consider a number of possible causes, along with associated reliability predictions
  - **To date, failure investigation results do not imply catastrophic consequences for the FPA**
  - **Results from planned analyses may give a more quantitative risk assessment**
- **Effort being coordinated with NASA GSFC**
  - **Impact to mission objectives**
  - **Long term stability of operational channels**



# Solar Diffuser Stability Monitor (SDSM)

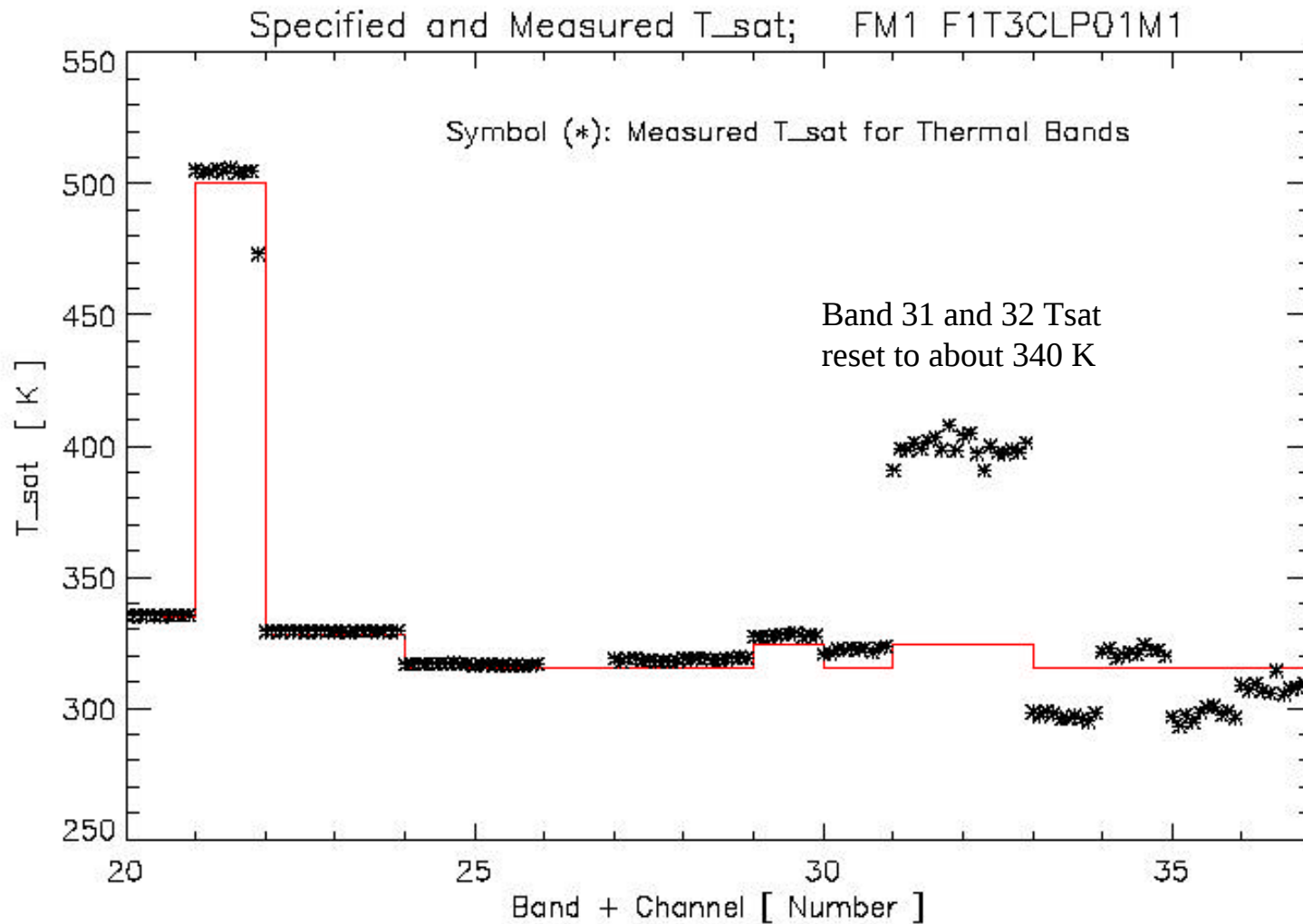


# MODIS Operational and Calibration Ranges



Vertical bars represent temperatures of 0.3Ltyp and 0.9 Lmax. Box symbols represent  $T_{max\_spec}$ , X symbols represent the value of  $T_{max\_set}$  (the higher of  $t_{Maxspec}$  or 315K), and the diamond symbols represent the temperature of  $L_{typ}$ .

*For FM1, Bands 33 and 35, we will not get science data when OBC-BB temperature above about 300 K*

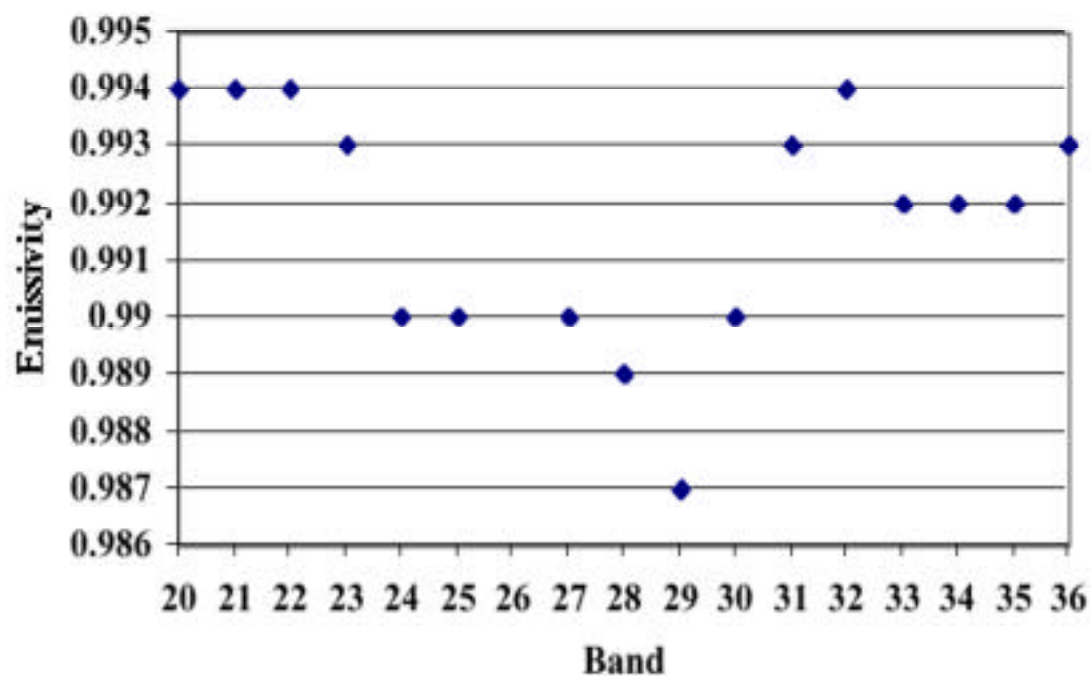




## OBC Blackbody Emissivity



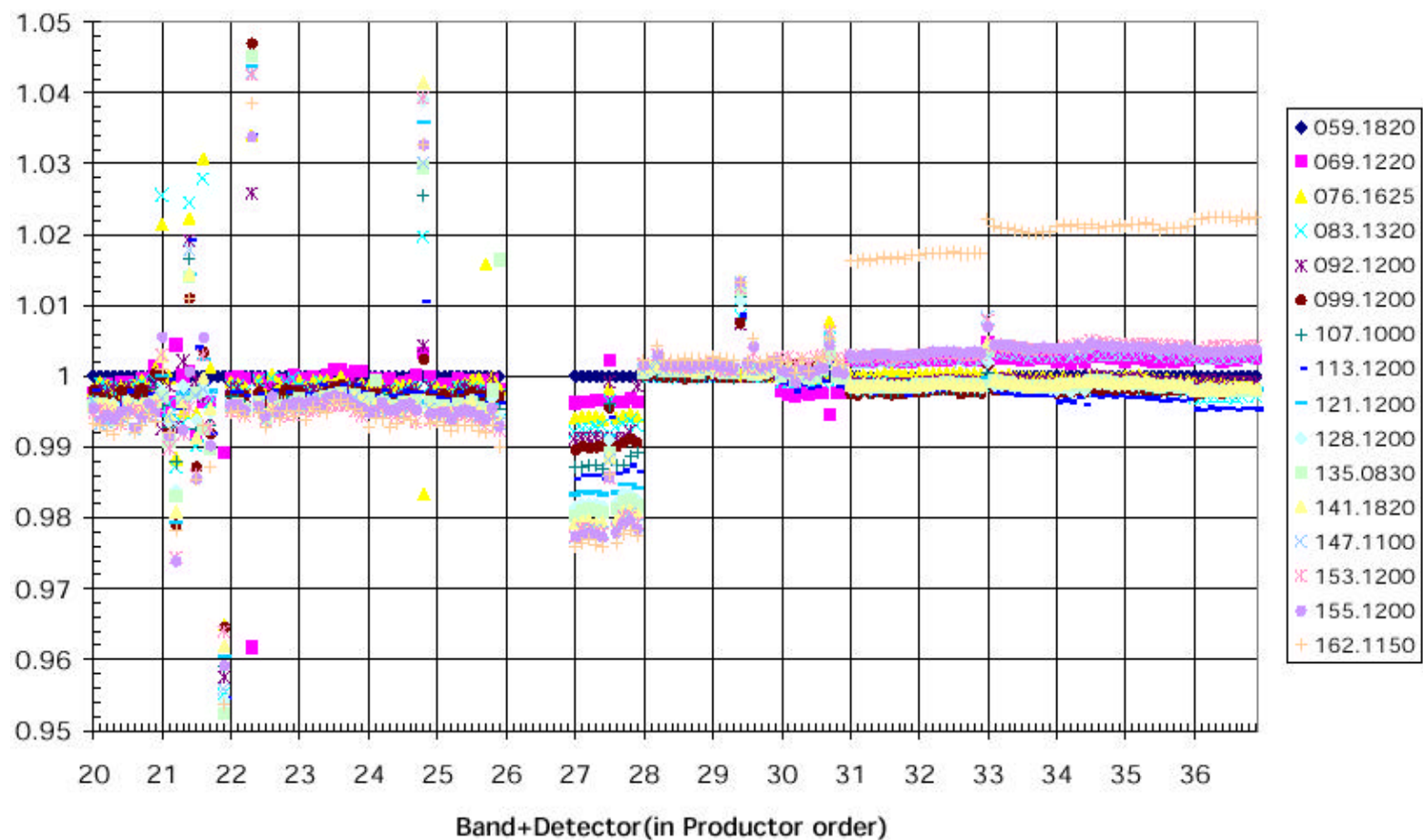
MODIS PFM Pre-launch OBC BB Emissivity



# MODIS PFM On-orbit TEB System Response(b1) Diffierence

(ltwk/Vdet = 110/226,T\_BB = 290 k, Mirror side 1)

(Normalized to Day 059.1820)

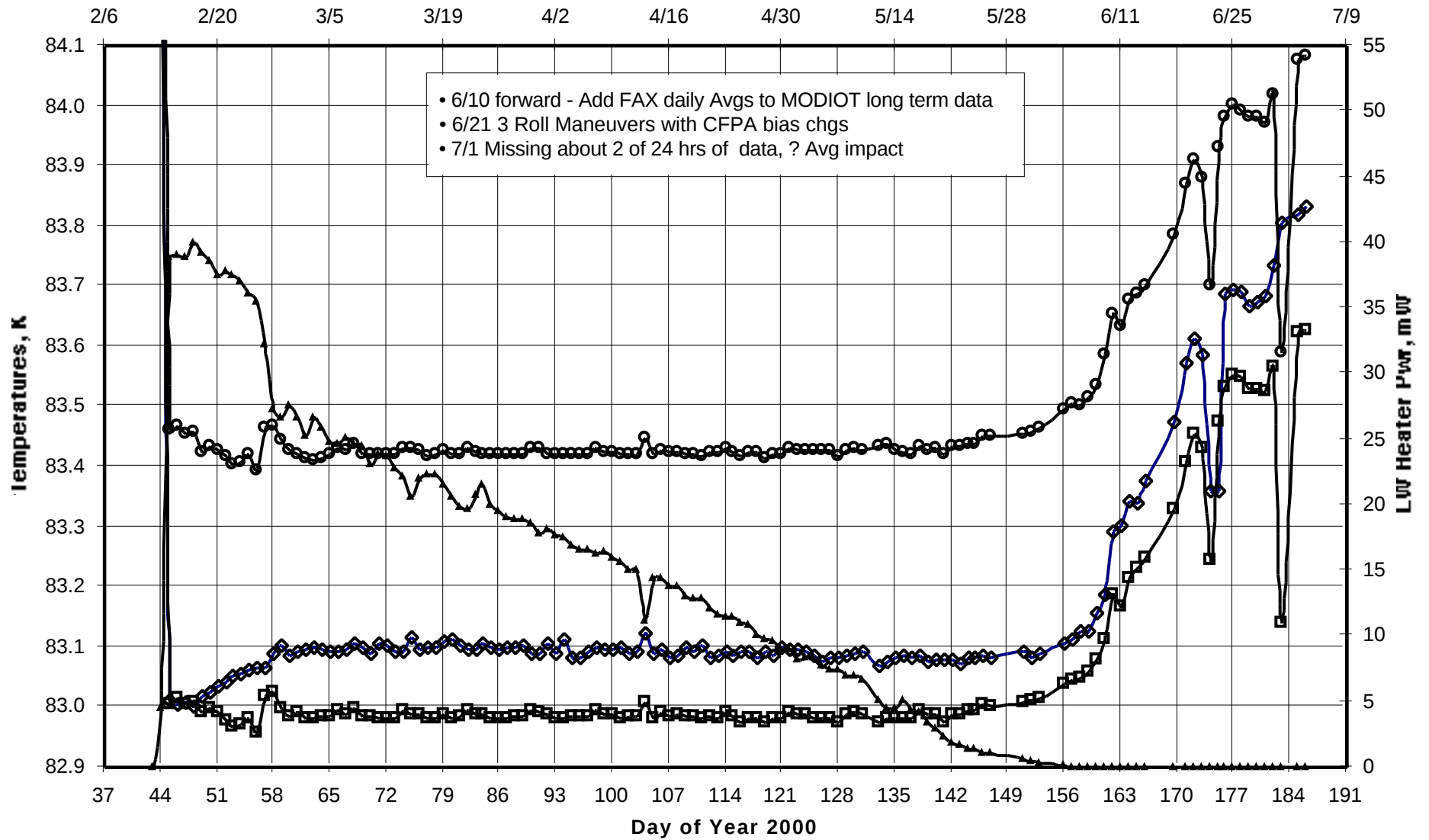


# TERRA MODIS PFM CFPA Anomaly Chart 1

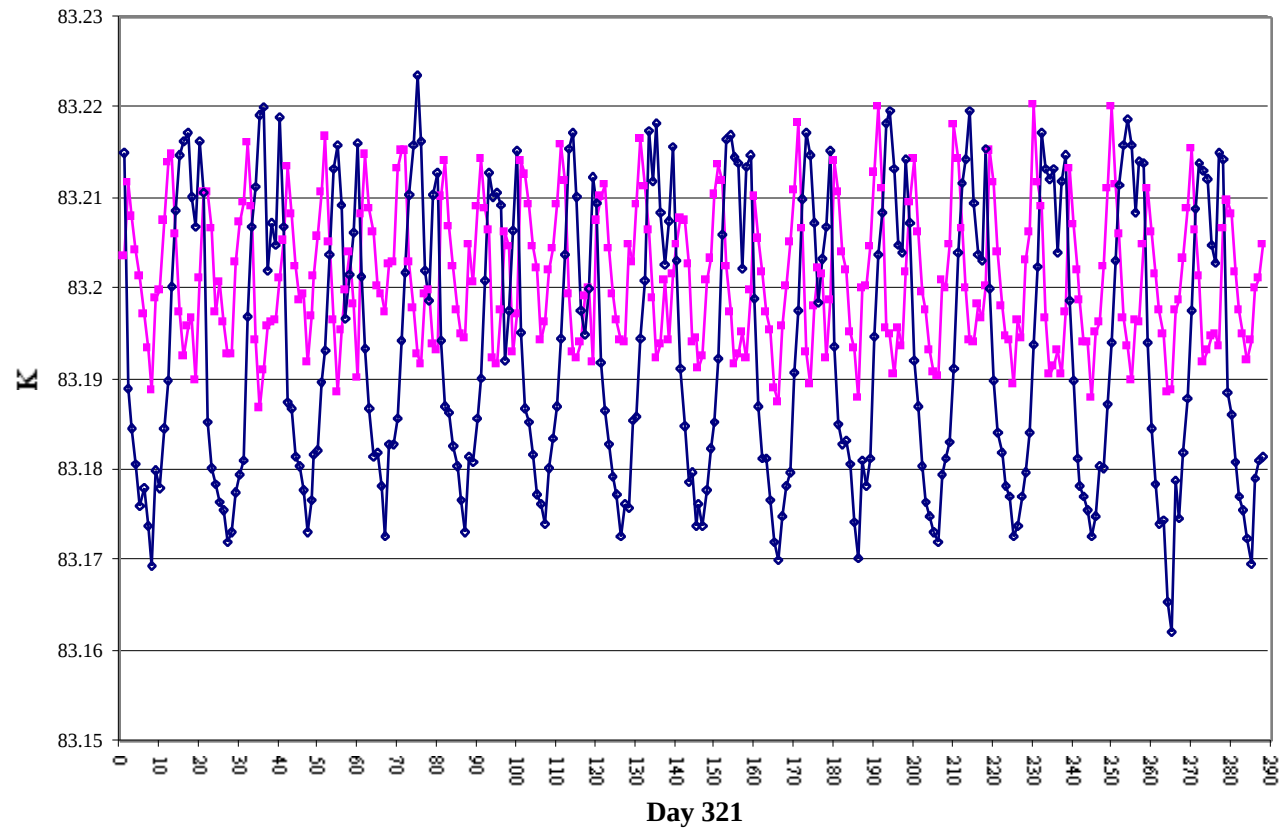
## Long Term Daily Avg Plot

◆ T\_CS\_OG   
 ■ T\_LWCFPA   
 ● T\_SMCFPA   
 —▲— P\_LWHTR

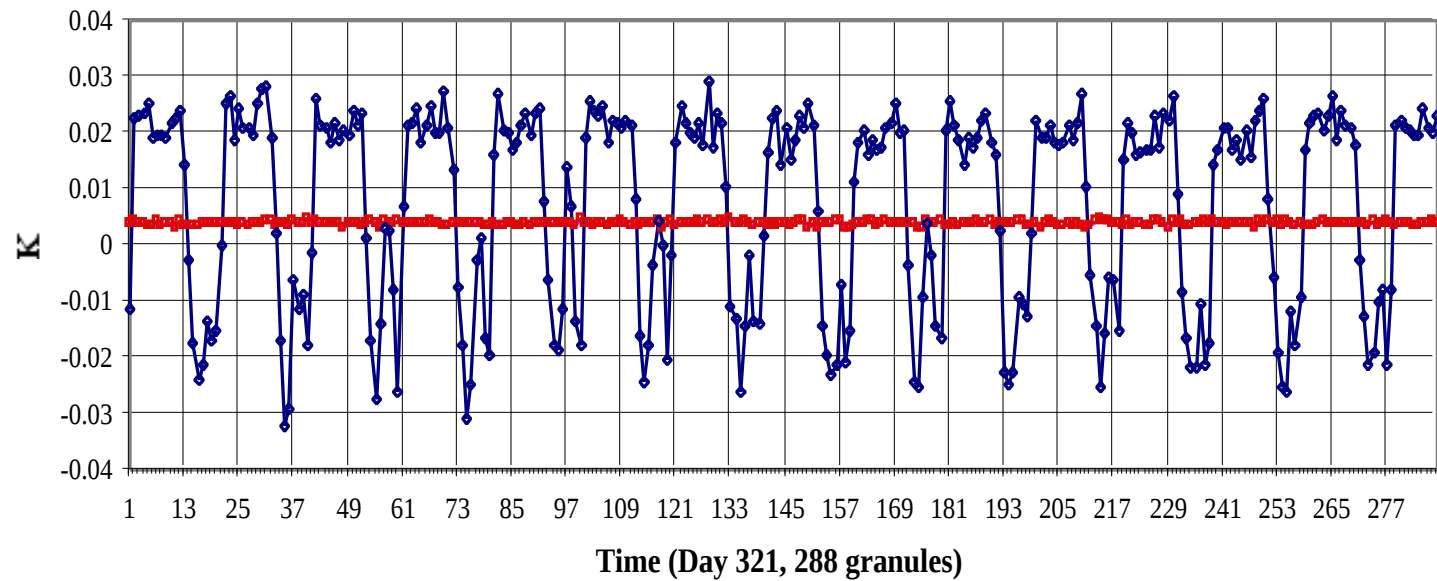
Date



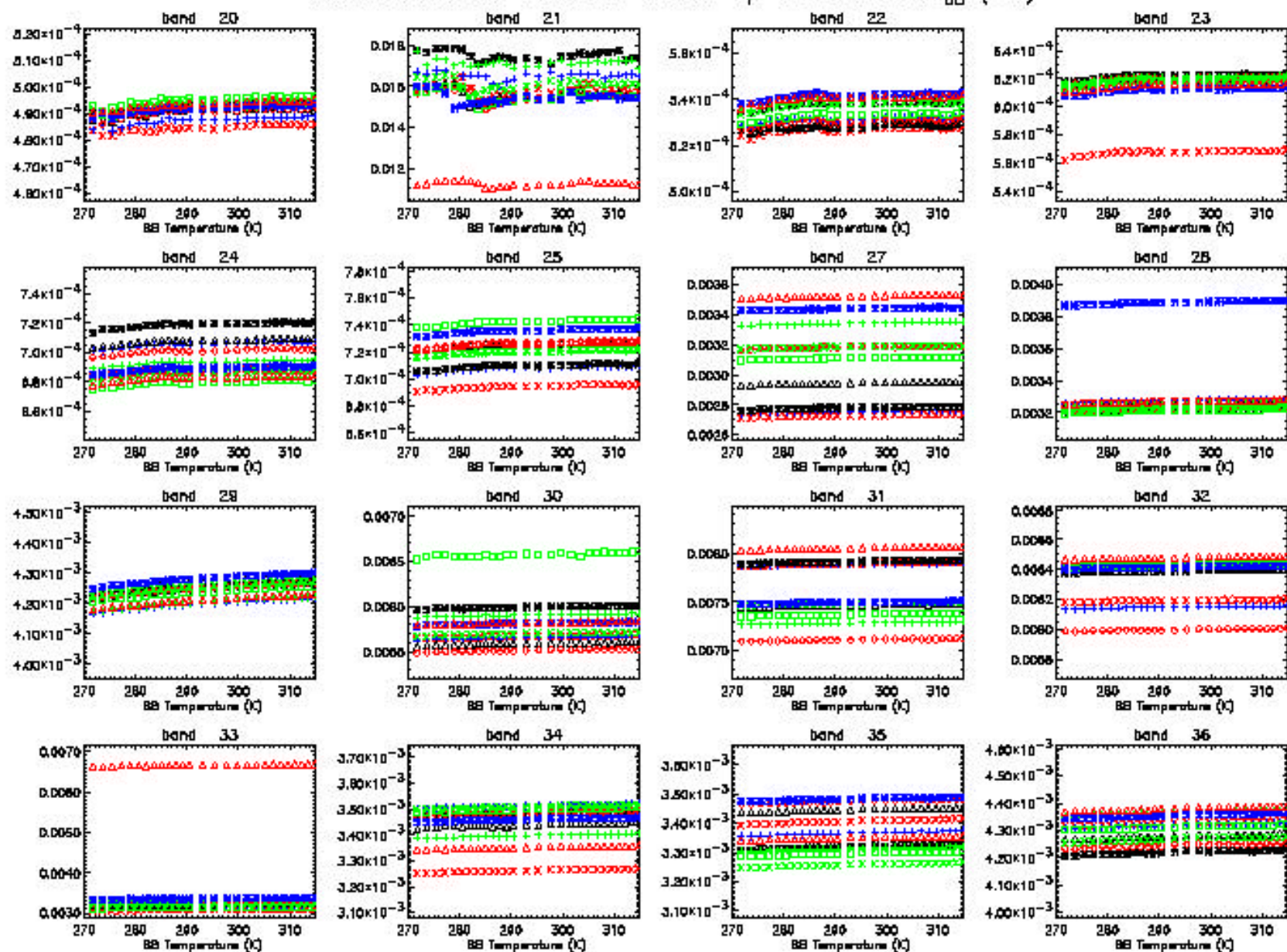
## MODIS PFM On-orbit Focal Plane Temperature (SMIR)



**MODIS PFM On-Orbit Averaged Focal Plane Temperature Difference**  
( $T_{\text{fpa@ms1}} - T_{\text{fpa@ms2}}$ , SMIR - Blue, LWIR - Red)



# MODIS PFM On-orbit BB Warm-up TEB b1 vs T<sub>BB</sub> (M1)



Ch1:Red x Ch2:Blu + Ch3:Blk \* Ch4:Blk Δ Ch5:Red ◊ Ch6:Gm □ Ch7:Grn x Ch8:Gm + Ch9:Blu \* Ch10:Red Δ

Data collect time: 305.2300 to 306.0230

detector in SBRS order

B\_side Itwk/Vdet = 79/110